

The Deep Roots of Geology: Tectonic History of Australia and its Margins expressed by Mantle Anisotropy

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1 The Deep Roots of Geology: Tectonic History of Australia and 2 its Margins expressed by Mantle Anisotropy

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5

6 **Abstract**

7 The Australian continental crust preserves a rich geological history, but it is unclear to what
8 extent this history is expressed deeper within the mantle. Scattering of surface waves
9 predominantly between 100-200 km depth by lateral gradients in seismic anisotropy, termed
10 Quasi-Love waves, offer potential new insights. Across Australasia over 275 new scatterers are
11 detected, and are found to be preferentially located near (1) the passive continental margins, and
12 (2) the boundaries of major geological provinces within Australia. Such lateral anisotropic
13 gradients imply pervasive fossilized lithospheric anisotropy within the continental interior, on a
14 scale that mirrors the crustal geology at the surface, and a strong lithosphere that preserves this
15 signal over billions of years. Along the continental margins, lateral anisotropic gradients may
16 indicate either the edge of the thick continental lithosphere, or small-scale dynamic processes in
17 the asthenosphere, such as edge-drive convection, tied to the transition from oceanic to
18 continental lithosphere.

19 **Introduction**

20 The Australian continent preserves a vast tectonic history, stretching from the Archean to
21 the present day ^{1,2}. By the Neoproterozoic, supercontinent cycles had assembled the three major
22 Australian cratons (North, South, and West), while later continental accretion added the eastern
23 orogens in the Phanerozoic ³. Since the break-up of eastern Gondwana, and the formation of its
24 rifted passive margins, the Australian continent has moved northwards colliding with SE Asia ^{4,5}.
25 Today, Australia is the fastest moving continent on Earth ⁶, exerting significant shear on the
26 underlying mantle asthenosphere. Such deformation is thought to cause a lattice preferred
27 orientation (LPO) in olivine generating strong azimuthal anisotropy aligned with plate motion ^{7,8}.
28 Intriguingly, this strong azimuthal anisotropy is not detected by shear-wave splitting studies,
29 with weak and complex splitting generally observed across Australia ⁹⁻¹¹. Instead a contrasting
30 contribution from fossilized anisotropy in the lithosphere has been proposed, frozen-in by past
31 deformational events ^{10,12}.

32 Observations of Love-to-Rayleigh surface wave scattering, termed Quasi-Love (QL)
33 waves ¹³, can be used to pinpoint the location of lateral gradients in seismic anisotropy in the
34 upper mantle (Fig. 1), such as those that may exist along present (and/or past) plate boundaries.
35 When a Love wave encounters such a boundary, a portion of its energy is scattered from Love
36 into Rayleigh wave motion, generating a QL wave (Fig. 1). The QL wave retains a similar
37 waveform shape as the fundamental Love wave (G1), hence ‘Quasi-Love’, but travels with the
38 slower velocity of the fundamental Rayleigh wave (R1). This principle can be exploited to
39 determine the distance (δx) the QL wave has travelled and thus estimate the location of the
40 anisotropic gradient. An anisotropic medium may also cause a small rotation of surface wave
41 polarization ¹⁴⁻¹⁶. These polarization-distortions are sometimes referred to in the literature as

42 ‘quasi-Love waves’ also ¹⁷, however their characteristics are distinguishable from QL wave
43 scattering as described here (see Methods for further details). Previously QL wave scattering has
44 been detected across a wide range of tectonic settings from continental collision ^{18,19}, seafloor
45 spreading ²⁰, subduction ^{13,21–23}, hotspot volcanism ^{24,25}, and former continental rift zones ²⁶.
46 Most recently QL waves have been recorded along the continental margin of eastern North
47 America ²⁷, implying localized dynamic processes and a disruption of the underlying mantle
48 flow-field at the margin, potentially induced by the transition from continental to oceanic
49 lithosphere. As of yet, however, there have been no such studies targeting the rich history of the
50 Australian continent and its extensive passive margins.

51

52 **Results**

53 Using seismic stations from the Australian National Seismograph Network, and shallow
54 earthquakes (<50 km depth) at 70°-180° epicentral distances, robust QL wave detections were
55 measured on 22% of the 2248 available event-station pairs (Fig. 2, Supplementary Data Table
56 S1) with 276 QL scatterers falling within the study area, as shown in Fig. 3. The QL waves
57 detected are of varying amplitude (represented in Fig. 3 by their symbol size) which is not only a
58 function of the strength of the anisotropic gradient but also the geometry of the anisotropy
59 relative to the angle of the raypath. The back-projected scatterers are widely distributed,
60 comprising one of the largest and most geographically diverse QL datasets produced to date, as
61 well as the first results within Australasia. While the QL scatterers are widespread and numerous
62 (the reliability of any single measurement should be treated with caution), when considered
63 collectively, they display intriguing relationships with certain tectonic features at the Earth’s
64 surface (Figs. 3-4 and Figs. S4-S11). Most noteworthy is the continental margins (cyan symbols;

65 Fig. 3a-b), that define the boundary between continental and oceanic crust ²⁸. Almost half of the
66 scatterers (47%) fall within 200 km of this boundary. This is higher than expected considering
67 just under two-fifths (38%) of the study area falls within 200 km (Fig. 4). The pattern is
68 particularly intriguing around the Tasman Sea, along the former rift boundary between Australia
69 and Te Riu-a-Māui/Zealandia (Fig. 3b). The scatterers appear to closely follow curvatures in this
70 boundary and even coincide with micro-continents such as the East Tasman Plateau (ETP) and
71 Gilbert Seamount Complex (GSC), which were isolated from the mainland during the rifting
72 process ²⁹.

73 The statistical significance of such a relationship between the location of scatterers and
74 the ocean-continent boundary can be explored further. Fig. 4a shows the cumulative distribution
75 function (CDF) of the minimum distance between each scatterer and the ocean-continent
76 boundary (cyan line). This is compared with the average CDF (solid black curve) for 100
77 distributions (thin grey lines) of 276 randomly located points across the study area, the same as
78 the number of QL scatterers in the dataset. The comparison indicates that there are more QL
79 scatterers closer to the ocean-continent boundary than seen in the averaged properties of the 100
80 random distributions (i.e. the cyan curve is above the upper dashed black curve out to a distance
81 of ~600 km). A two-sample Kolmogorov–Smirnov test ^{30–32} is performed to test the null
82 hypothesis that the two CDFs are from the same continuous distribution. A p-value of 0.04 is
83 found, suggesting that the null hypothesis can be rejected at 4% significance level, indicating that
84 the QL scatterers can indeed be distinguished from a random distribution. Furthermore, sub-
85 dividing the dataset into oceanic versus continental scatterers (Fig. S10) indicates that the spatial
86 correlation is even stronger within the oceans (p value = 0.01), with 38% of the oceanic

87 scatterers located within 100 km of the ocean-continent boundary, which is more than 2.5
88 standard deviations higher than the mean of the random distributions.

89 Away from the continental margins, scatterers can also be found within the Australian
90 continental interior coinciding with the boundaries between major geological provinces³³ (yellow
91 symbols; Fig. 3). All 73 scatterers that fall within the Australian mainland are located less than
92 156 km away from such a boundary, and 95% are located within 100 km (Fig. 4c-d). If the same
93 number of scatterers were distributed randomly across the continental mainland, we would
94 typically expect the maximum distance of scatterers to be twice as far away, up to 310 km from
95 the boundary, and the number within 100 km to be 15% less. A two-sample Kolmogorov–
96 Smirnov test^{30–32} between the CDF of the locations of the observed scatterers and the mean CDF
97 of 100 random distributions returns a p-value of 0.0006. This suggests that the null hypothesis;
98 that the randomly distributed points and the continental mainland scatterers are from the same
99 continuous distribution, can be rejected at the 0.06% significance level. It can therefore be
100 implied that the spatial correlation is significantly better than random.

101 Looking more closely at where the continental scatterers are found, several appear to be
102 associated with the Proterozoic orogenic blocks and deformation belts of central and northern
103 Australia (Fig. 3b). These tectonic features are associated with prominent gravity anomalies (Fig.
104 3c), indicating that they represent major crustal structures with potentially deeper expressions in
105 the mantle lithosphere also³⁴. Multiple scatterers can be found on the edge of the strongest
106 gravity anomalies, including the east-west trending Musgrave (Mu) and Arunta (Ar) blocks in
107 central Australia (dashed white circle Fig. 3c), and the north-south trending Darling Fault at the
108 southwest corner of the Yilgarn Craton (YC). Interestingly, there are several large amplitude
109 scatterers associated with the Kimberley (K) region of northwestern Australia, which has been

110 argued to be a small cratonic block underlain by an Archean mantle root ^{2,35}. Within the
111 submerged continental area of Te Riu-a-Māui/Zealandia a correspondence with topographic
112 variations at the surface can also be seen (Fig. 3a), particularly near New Caledonia (NC) and
113 along the New Caledonia Trough and Norfolk Ridge. This hints at an association with the
114 internal geological architecture of Te Riu-a-Māui/Zealandia similar to that found within
115 Australia.

116 As well as old tectonic boundaries within the continental interior and along its margins,
117 as would be expected multiple scatterers can be found at present day plate boundaries (white
118 dashed lines and white circles; Fig. 3a), such as the New Hebrides Trench (NHT).

119

120 **Discussion**

121 *Errors and uncertainties on the interpretation of Quasi-Love waves*

122 It is important to first consider the uncertainties associated with the analysis of QL waves
123 before making any physical interpretations. A brief summary is provided here in terms of the
124 mechanism, magnitude, depth, and location (for a more detailed description please see Methods).
125 Numerous studies have demonstrated that QL waves are most efficiently generated by lateral
126 gradients in seismic anisotropy (i.e. changes in either the strength or orientation of seismic
127 anisotropy), and are difficult to reproduce with purely isotropic velocity changes ³⁶⁻³⁹. Higher
128 mode Rayleigh waves with velocities faster than the fundamental mode may potential interfere
129 within the selected time window, but shallow earthquakes (<50 km) should preferentially excite
130 the fundamental mode rather than higher modes ⁴⁰. In addition all QL wave detections are
131 visually inspected to ensure similarity in waveform shape with the Love wave (G1).

132 The magnitude of the QL wave (i.e. size of the symbols in Fig. 3), is expressed as the
133 relative amplitude between QL and G1. This magnitude is a function of both the strength of the
134 lateral anisotropic gradient and the propagation azimuth relative to the anisotropic geometry. If
135 the propagation direction is parallel or perpendicular to the symmetry axis then the amplitude
136 will be zero. A particularly large QL amplitude is therefore indicative of the presence of a strong
137 gradient but a small amplitude, or even absence of a QL wave entirely, is not conclusive
138 evidence of a weak or non-existent gradient. It is therefore not possible to directly interpret the
139 amplitudes of the observations.

140 Depth sensitivity kernels calculated for a QL wave with a dominant frequency of 0.01 Hz
141 (100 s) suggest a broad range of sensitivity throughout the upper mantle, with a peak around
142 100-200 km depth (refer to Fig. S1 of Levin et al. (2007), Fig. 9 of Chen and Park (2013)). While
143 potential contributions from shallower or deeper depths cannot be entirely ruled out, the source
144 of the QL waves most likely resides within the upper mantle, and not within the crust.

145 The location of the QL scattering source is found by back-projection along the great-
146 circle path (Figs. 1, S2, S5). Potential location errors are estimated to be on the order of 100 km,
147 considering the typical wavelength of 100s surface waves, the results of numerical
148 experiments¹⁸, potential deviations from the great-circle path, and assumptions made the
149 uniformity of the average Rayleigh and Love phase velocities between source and receiver (refer
150 to Methods for details). This is consistent with the spatial distribution of results found relative to
151 certain tectonic and geological features (Figs. 3-4, S9).

152

153 *New implications for the upper mantle beneath the Australian continent and its margins*

154 The results of this study reveal the widespread presence of lateral anisotropic gradients
155 capable of generating QL waves. The location of such gradients appear closely linked to past
156 tectonic boundaries, both within the continental interior and along the rifted margins. Such
157 features visible from the Earth's surface, by inference, must also extend to depth within the
158 upper mantle. Considering the peak sensitivity of Love-to-Rayleigh scattering at 100s period is
159 expected to be between 100-200 km depth ^{18,22}, this depth range would sample the old thick
160 cratonic lithosphere in central and western Australia (light colors; Fig. 3d), while for eastern
161 Australia the lithosphere is younger (Phanerozoic) and much thinner, ~75km ⁴¹. Except in
162 regions of thick continental lithosphere, the QL scatterers should therefore correspond to lateral
163 anisotropic gradients in the asthenosphere, where olivine LPO is expected to develop through
164 mantle flow ⁴².

165 A large number of scatterers in this dataset are associated with continental margins (Figs.
166 3a-b, 4a-b), similar to the findings of Servali et al., (2020) who detected pervasive scatterers
167 along the eastern North America passive margin. For western and southern Australia the margin
168 is co-located with the edge of the thick continental lithosphere (Fig. 3d, Fig. 5). QL scatterers
169 along this boundary may therefore simply reflect the transition from active flow in the
170 asthenosphere to anisotropy fossilized within the lithosphere, which is supported by the first-
171 order patterns seen in tomographic models of azimuthal anisotropy (Figs. 3d, S11). Alternatively
172 a significant step in lithospheric thickness may trigger edge-driven convection (EDC) at the
173 margin ⁴³, which can generate high-strain rates and localize dislocation-creep deformation
174 capable of overprinting LPO anisotropy to beneath the lithospheric step ⁴⁴. For the southern
175 Australian margin, geodynamical models have suggested the formation of broad upwellings in
176 the region behind the trailing edge of the continent due to EDC ⁴⁵, which may potentially explain

177 the broad zone of QL scatterers in this region. Numerous scatterers along the eastern margin of
178 Australia, and the western margin of Te Riu-a-Māui/Zealandia, are however less easy to explain
179 given that the lithosphere appears consistently thin (<100 km) across this region. This may
180 instead hint at a fundamental difference between how continental versus oceanic lithosphere
181 couples to the underlying asthenosphere, perhaps in terms of differences in viscosity ⁴⁶.
182 Alternatively finer-scale variations in lithospheric thickness along the margins may exist than
183 those currently resolved in Figure 3d. Such small-scale 3D topography on the lithosphere-
184 asthenosphere boundary may locally divert or disrupt the mantle flow-field, such as localized
185 EDC processes proposed at the Newer Volcanics Province in SE Australia ^{44,47}. Recent modeling
186 work ⁴⁸ suggests that even when the edge of the continent is parallel with the background mantle
187 flow field (as is generally thought to be the case for eastern Australia) dynamical instabilities in
188 mantle flow can still be found at the edges of thick continental lithosphere, and may be enhanced
189 where there are smaller-scale variations in the geometry of the boundary. It remains to be seen
190 however whether the same effect could be generated beneath eastern Australia with only small
191 changes in lithospheric thickness and/or viscosity.

192 Another significant finding of this study is the presence of numerous QL scatterers
193 associated with the major geological provinces of Australia (Fig. 3). Many of these scatterers are
194 located within thick cratonic lithosphere (Fig. 3d), and are associated with Archean-Proterozoic
195 terranes and past orogenies (Fig. 3b). This suggests QL scatterers are sensitive to lithospheric
196 anisotropy frozen-in since the last deformational episode, and therefore that they reflect the
197 evolution of the continent. Furthermore this implies that geological boundaries are preserved
198 over billions of years not only in the crust but also at deeper depths as anisotropic gradients in
199 the lithospheric mantle. If fossilized lithospheric anisotropy is pervasive, yet heterogenous with

200 strong lateral gradients, including the potential for multiple layers to exist within cratonic
201 lithosphere⁴⁹, this would add interference to the SKS splitting accrued from the strong active
202 flow in the underlying mantle, causing a loss of coherency in the overall signal. This may explain
203 the long-standing conundrum as to why SKS splitting is perplexingly weak and complex across
204 Australia ⁹⁻¹¹, despite fast Australian plate motion, and strong azimuthal anisotropy imaged
205 beneath the continent (Figs. 3d, S11) ^{7,8,12}. Further studies to constrain lithospheric anisotropy
206 may therefore help to decode the complex SKS splitting, as well as provide further insights on
207 the past deformational history of continent.

208

209 **Conclusion**

210 Observations of QL waves have revealed extensive lateral anisotropic gradients beneath
211 the Australian continent and its margins that have otherwise, until now, been impossible to
212 resolve with smooth seismic tomographic methods (e.g. Fig. 3d). Based on the newly acquired
213 results, it is speculated that at the continental margins small-scale convective processes, such as
214 EDC, or localized disruptions to the mantle flow field, may be occurring where the lithosphere
215 transitions from continental to oceanic. Alternatively QL scattering may be explained by the
216 difference between fossil anisotropy in old thick continental lithosphere and active flow in the
217 oceanic asthenosphere. However, this is harder to justify where the continental lithosphere is
218 younger and thinner on either side of the Tasman Sea. In the future QL scatterers may be used as
219 a predictive tool, helping to constrain the geophysical signature of major crustal/lithospheric
220 structures at depth, such as those currently hidden under cover ⁵⁰, and in particular in identifying
221 those with deep mantle connections that may suggest potential for mineralization ^{31,51}.

222

223 **Methods**

224 *Overview of data selection and processing*

225 Earthquakes from 1994-2020 of magnitude >6.5 , depth <50 km, and in the epicentral
226 distance range of 70° - 180° were selected to ensure large amplitude surface waves and enough
227 separation to distinguish QL from R1. All data were low-pass filtered at 100s (0.01Hz) to
228 amplify the fundamental Love and Rayleigh wave components and suppress overtones.
229 Following previous studies, the delay time (δt) between the QL and G1 is calculated using the
230 cross-correlation technique^{18,20,27}. In the final step the delay time is converted to distance (see
231 Fig. 1b), and back-projected along the great-circle path to determine the location of the scatterer.
232 If the polarization of the surface waves suggests a deviation from the great-circle path of more
233 than 10° then the result is discarded (Fig. S8). In this study a few modifications have been made
234 to the cross-correlation technique to improve detection and measurement quality. These are
235 outlined in further detail below along with a more detailed description of the methodology.

236

237 *Strategy for Detection and Measurement of Quasi-Love Waves*

238 Data for each event is requested from the IRIS Data Management Centre (ds.iris.edu/) for
239 5000 seconds following the origin time. After low-pass filtering at 0.01Hz, seismograms are
240 rotated into the vertical, radial, and transverse, and normalized by the maximum amplitude on
241 each component. The maximum on the transverse signifies the fundamental Love wave (G1),
242 and on the vertical the fundamental Rayleigh wave (R1), indicated by the vertical dashed green
243 and dashed blue lines in Fig. S1. Initial quality control restricts events to those with a signal-to-
244 noise ratio (SNR) of the Love wave greater than 5, where the ‘noise’ amplitude is determined by
245 the first 300 seconds of the seismogram. A cut-off of 5 was determined empirically from the

246 data. In general large amplitude Love waves are a necessary pre-requirement to generating
247 observable QL waves, and thus detections are more likely in the Love radiation maxima (and
248 Rayleigh radiation minima).

249 Most previous studies only measure QL waves on the vertical component even though
250 the particle motion is elliptical and distributed across both the vertical and radial components
251 (Fig. S1b-c). Same as for R1, for QL, the vertical and radial components are 90° out of phase.
252 The first derivative of the radial component (dashed red line Fig. S1d) matches the vertical
253 component. This is a useful diagnostic feature. Any seismograms that do not display this
254 characteristic during the visual inspection process are discarded, and the QL wave must be
255 clearly visible on both components, not just the vertical. Furthermore, stacking of the vertical
256 component with the derivative of the radial component amplifies the QL (and R1) while
257 dampening the noise (magenta line Fig. S1e). This improves both the visibility and confidence of
258 the QL detection.

259 The next step is to cross-correlate this QL stack with G1 on the transverse component.
260 Fig. S1f shows the absolute amplitude of the cross-correlation (thin black line) as a function of
261 lag time relative to the position of G1. Peaks of this cross-correlation function (i.e. maxima,
262 where the derivative equals zero) are identified by blue crosses and connected by a new blue
263 curve. Peaks in this new curve are then selected and highlighted by blue circles. Typically the
264 largest peak is associated with the correlation between G1 and R1 (as seen in Fig. S1f). The QL
265 peak, if present, is a secondary peak located before the highest R1 peak but after the G1 arrival
266 (indicated by dashed green line). The inferred QL delay time is marked by a green asterisk and
267 dotted green line. The time shifted Love wave is plotted for comparison with the QL stack in Fig.
268 S1e.

269 All seismograms and cross-correlation functions are visually inspected for quality based
270 on the criteria outlined above. Only those with a clearly visible and distinct QL wave are kept. It
271 is possible however that multiple QL waveforms may appear in one seismogram if multiple
272 scatterers are located along the great circle path. In such cases, for simplicity, the first QL wave,
273 closest to G1, and therefore closest to Australia is recorded.

274 The last step in the process is to convert the delay time (δt) into distance and back-project
275 the location of the scatterer along the great-circle path. Considering that the QL wave travels
276 distance, δx , of the total path, Δ , (Fig. 1a) with the slower Rayleigh wave phase velocity (V_R),
277 the delay time can be expressed as:

$$278 \quad \delta t = \frac{\delta x}{V_R} - \frac{\delta x}{V_L} \quad (1)$$

279 where V_L is the Love wave phase velocity. To negate having to use either an average guess for
280 the phase velocities or calculate individually for every seismogram, the following substitution
281 can be made,

$$282 \quad \delta t = \frac{\delta x(t_R - t_L)}{\Delta} \quad (2)$$

283 where Δ is the epicentral distance between source and receiver, and $t_R - t_L$ is the time delay
284 between the Rayleigh (R1) and Love (G1) waves. Simply rearranging this equation for δx gives,

$$285 \quad \delta x = \frac{\delta t \Delta}{(t_R - t_L)} \quad (3)$$

286 The scatterer distance is therefore independent of the absolute phase velocities or the arrival
287 times, and only depends on the relative time difference between G1 and R1. This is useful given
288 the exact start of the R1 is difficult to determine, especially considering likely overlap with QL.
289 Instead the time difference between the maximum amplitude of G1 and R1 can be utilized, which
290 is much easier to calculate from the seismogram.

291

293 In this study the term ‘Quasi-Love waves’ refers to surface wave polarization anomalies
294 that are characteristic of Love-to-Rayleigh scattering. This is distinct from distortions or
295 rotations of the polarization of surface waves by azimuthal anisotropy that are sometimes also
296 called ‘quasi-Love’ waves or ‘quasi-Rayleigh’ waves elsewhere in the literature ¹⁴⁻¹⁶. Such
297 polarization distortions are analogous to the terminology of ‘quasi-P’ and ‘quasi-S’ when body-
298 wave polarizations are altered by Earth 3D heterogeneity ¹⁷. In such polarization-distortion cases,
299 the ‘quasi-Love’ wave will exhibit particle motion that is still mostly transverse polarized but
300 with a small added vertical component (see Figure 2 of ¹⁶). In contrast, Love-to-Rayleigh
301 scattering generates Quasi-Love waves that are predominantly polarized in the vertical-radial
302 plane, displaying ellipticity characteristic of Rayleigh waves with a 90° phase lag between the
303 vertical and radial components ¹⁷. By comparing the waveform shape of the QL wave on both the
304 vertical component and the first-derivative of the radial component during the measurement
305 process (e.g. Fig. S1d) ensures the presence of characteristic Rayleigh wave motion as expected
306 for QL wave scattering. Incorrect interpretations due to slight polarization-distortions are thus
307 considered unlikely in the present dataset.

308 Previous studies have shown that Love-to-Rayleigh scattering is most efficiently
309 generated by lateral gradients in seismic anisotropy (i.e. changes in seismic anisotropy either in
310 strength or orientation) rather than changes in isotropic structure ³⁶⁻³⁹. Lateral anisotropic
311 gradients of only 3-5% are enough to generate QL waves with 5-10% amplitude of the
312 fundamental Love wave ^{19,38}, compared to unrealistic perturbations of 20% or more which would
313 be required to produce the same effect with isotropic velocity changes only ^{36,37}. Additionally QL
314 waves are not typically seen in synthetic seismograms generated from velocity models without

315 lateral variations in azimuthal anisotropy (Fig. S6) ^{20,26,27}. It is worth noting that lateral
316 anisotropic gradients will also cause Rayleigh-to-Love scattering, generating Quasi-Rayleigh
317 waves, in addition to Quasi-Love waves ^{17,52}. Such Quasi-Rayleigh waves will appear on the
318 transverse component trailing G1 and thus are more challenging and less commonly studied
319 compared to QL waves that arrive ahead of R1 on the vertical and radial components ⁵². As a
320 relatively large amplitude of G1 is a pre-requirement of this study (SNR>5), the amplitude of R1
321 is often comparatively less, and thus Quasi-Rayleigh waves are not expected to be visible on the
322 same seismograms as prominent QL waves (Fig. S1).

323 The relative amplitude between the QL wave and the fundamental Love wave is
324 represented via the cross-correlation amplitude, which is used to determine the symbol size in
325 Fig. 3. QL amplitude is dependent not only on the strength of the anisotropic gradient but also on
326 the anisotropic geometry relative to the propagation path. Maximum QL conversion will occur
327 when the propagation azimuth is 45° to the symmetry axis, and zero conversion when the
328 propagation azimuth is either parallel or perpendicular ^{20,53}. This creates a symmetric four-lobed
329 radiation pattern with both positive and negative polarities of the QL wave, which in theory
330 could be used to infer not only the location of an anisotropic gradient but also the geometry of
331 anisotropy. It is safe to say that large QL amplitudes can be interpreted as the existence of a
332 strong gradient in anisotropy but smaller amplitudes do not necessarily equate to a weaker
333 gradient. Likewise a lack of detection of QL waves is not conclusive evidence that anisotropic
334 gradients do not exist along the great circle path. The angle between the propagation path and the
335 geometry of the anisotropy may simply be non-optimal for Love-to-Rayleigh scattering to occur.

336 While lateral variations in seismic anisotropy are the most likely cause of the QL wave
337 scatterers, an important consideration is at what depth they are generated. Depth sensitivity

338 kernels calculated for QL wave scattering with a dominant frequency of 0.01 Hz (100 s) suggest
339 a broad range of sensitivity throughout the upper mantle but with a peak around 100-200 km
340 depth (refer to Fig. S1 of Levin et al. (2007), Fig. 9 of Chen and Park (2013)). While potential
341 contributions from shallower or deeper depths cannot be entirely ruled out, the source of the QL
342 waves most likely resides within the upper mantle, and not within the crust.

343 As well as the depth, there is also a degree of uncertainty associated with the horizontal
344 location of the QL wave source. Firstly, the width of the lateral anisotropic gradient cannot be
345 exactly determined. For example, a wide gradational zone can produce the same signal as a sharp
346 localized gradient, but both are represented as a point scatterer at the center of the lateral gradient
347 distribution²⁰. Secondly, the smallest feature that can be resolved is likely on the order of 1/4th of
348 the wavelength¹⁸. Surface waves with a period of 100s have a wavelength of ~400 km,
349 suggesting ~100 km resolution. This is consistent with numerical experiments based on simple
350 anisotropic models with an equatorial band of 4% azimuthal anisotropy, in which the source of
351 the QL scatterer could be correctly located to within 100 km following a similar methodology to
352 that applied here¹⁸.

353 The real Earth however is much more complex than such idealized models and location
354 errors may be influenced by 3D heterogeneity. In particular 3D velocity structure may cause
355 deviations from the great circle path (e.g. multi-pathing), which could result in a location error
356 when the QL wave source is back-projected. In order to limit this, a comparison is made between
357 the polarization of the surface waves and the source-receiver back-azimuth. Any events with
358 deviations greater than 10° are thrown out from the dataset, but most are in agreement to within a
359 few degrees (Fig. S8). Slight rotations away from the radial/vertical (Rayleigh) and transverse
360 (Love) polarizations are to be expected when in the presence of seismic anisotropy¹⁴⁻¹⁷.

361 Nonetheless, a maximum deviation of 10° from the great circle path would result in a location
362 error of 390 km for a QL wave scatterer at a typical distance of 20° from the receiver (Fig. S8).
363 A typical deviation of 2° would however cause a location error of <100 km.

364 A location error may also result regarding the assumption that the average Rayleigh and
365 Love phase velocities between the QL scatterer and the seismic station are the same as the
366 average phase velocities along the entire great-circle path. Considering that many of the
367 propagation paths cross an ocean basin before generating a QL wave and then continue through
368 mainly continental regions after scattering at the ocean-continent boundary, this may cause the
369 phase velocities along the respective path segments to substantially differ, either due to changes
370 in isotropic or anisotropic properties or both. If the average phase velocities were 5% higher
371 between the QL scatterer and the receiver (i.e. δx), than between the earthquake source and the
372 receiver (i.e. Δ), then this would directly correspond to a 5% error in the scatterer location as a
373 function of distance along the great circle path. For example, if a typical QL scatterer is
374 estimated to be 20° away from the receiver, then a 5% velocity error would equate to a distance
375 error of 1° (~ 110 km).

376 Overall based on the above considerations, typical location errors are expected to be on
377 the order of 100 km but the further away the QL scatterers are the size of the error likely
378 increases. This may be supported by the data as Fig. 4 suggests that QL scatterers are
379 preferentially located near the ocean-continent boundary to within 400-500 km, and with
380 geological provinces to within 100-150km, based upon the steepness of the CDF curve relative
381 to the random distributions.

382 Of further note, the methodology applied here is based upon detecting a single scatterer
383 located closer to the receiver than to the source. This ensures enough separation from the

384 Rayleigh wave on the radial and vertical components in order to be able to identify the QL wave.
385 If a QL scatterer is located near the source then the QL arrival will overlap with the Rayleigh
386 wave making it hard to distinguish. Likewise, in the cases where more than one potential QL
387 wave is identifiable on the seismogram, potentially due to multiple lateral gradients in seismic
388 anisotropy along the propagation path, only that which is closest to the receiver is included in the
389 dataset as it is the easiest to distinguish. If the seismogram contains multiple closely arriving QL
390 waves generating a complex overlapping signal, then this will be discarded during the quality
391 control process as the QL wave is required to have a similar waveform shape as G1 (Fig. S1).

392 Lastly, it is possible for QL waves to resemble higher-mode Rayleigh waves^{26,54}. In this
393 study only shallow events less than 50 km depth are considered are these should primarily excite
394 the fundamental modes rather than overtones, and all seismograms are low pass filtered at 100s
395 which should aid separation of the modes⁵⁵. During the quality control process, the requirement
396 that the QL wave should have a similar waveform shape to G1 (see Fig. S1), should also help
397 distinguish potential interference from higher mode Rayleigh waves, as these would not be
398 expected to possess the same waveform shape as G1. Following these conditions, synthetic
399 seismograms demonstrate that the observed QL waves are unlikely to be the result of higher
400 mode Rayleigh waves (Fig. S6)^{18,26,27}.

401

402 **Data Availability**

403 The seismic data used in this study is from the Australian National Seismograph Network
404 (network code AU; www.fdsn.org/networks/detail/AU/), and can be freely obtained via the Data
405 Management Center of Incorporated Research Institutions for Seismology (ds.iris.edu/). A full
406 list of results is provided in Supplementary Table S1.

407

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414

415 **Author Contributions and Competing Interests Declaration**

416 The sole author of this work, C.M. Eakin, conducted the study in its entirety. The author declares
417 no competing interests.

418

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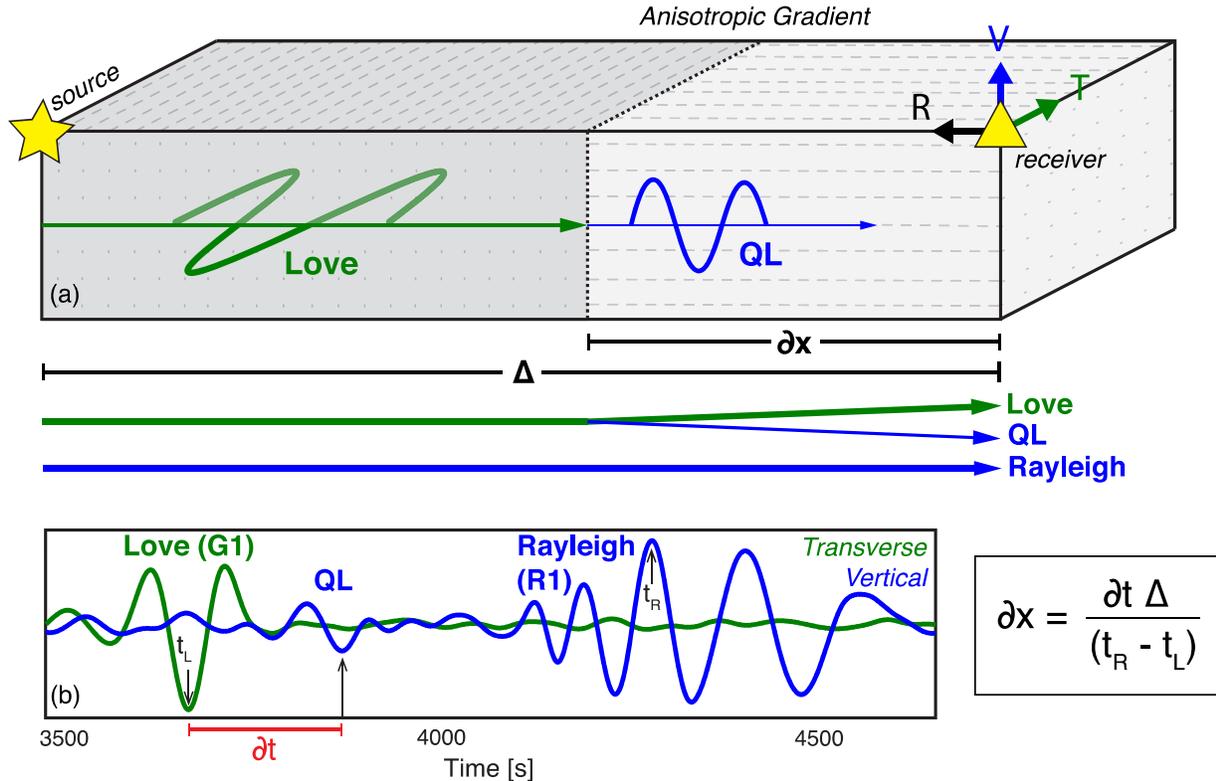
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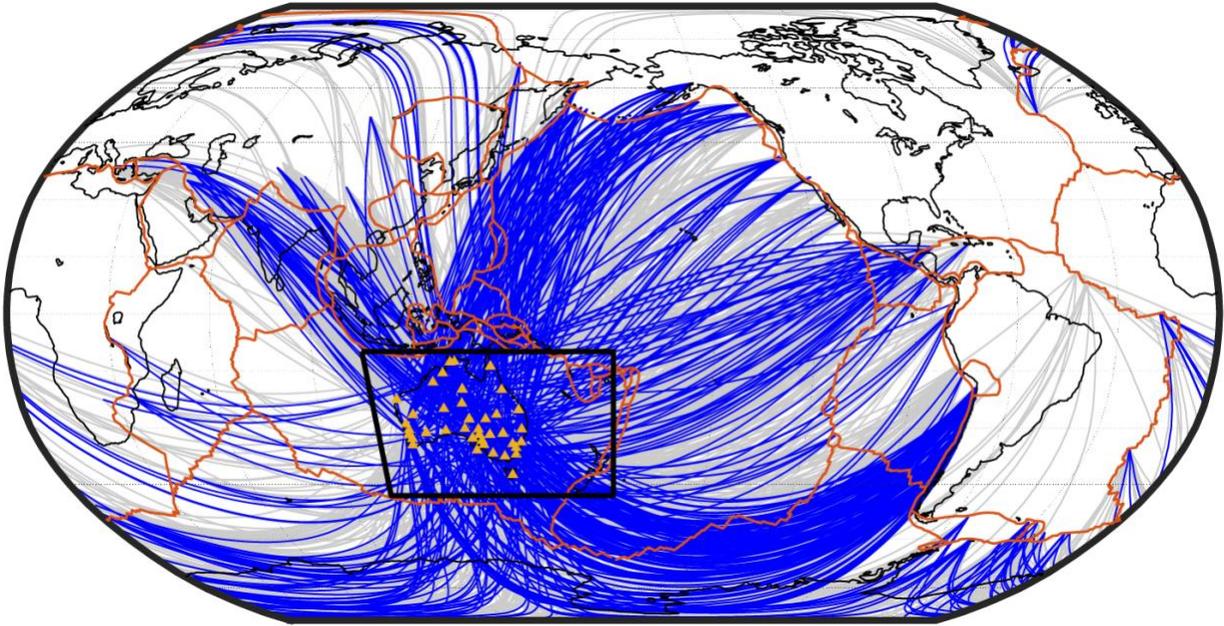
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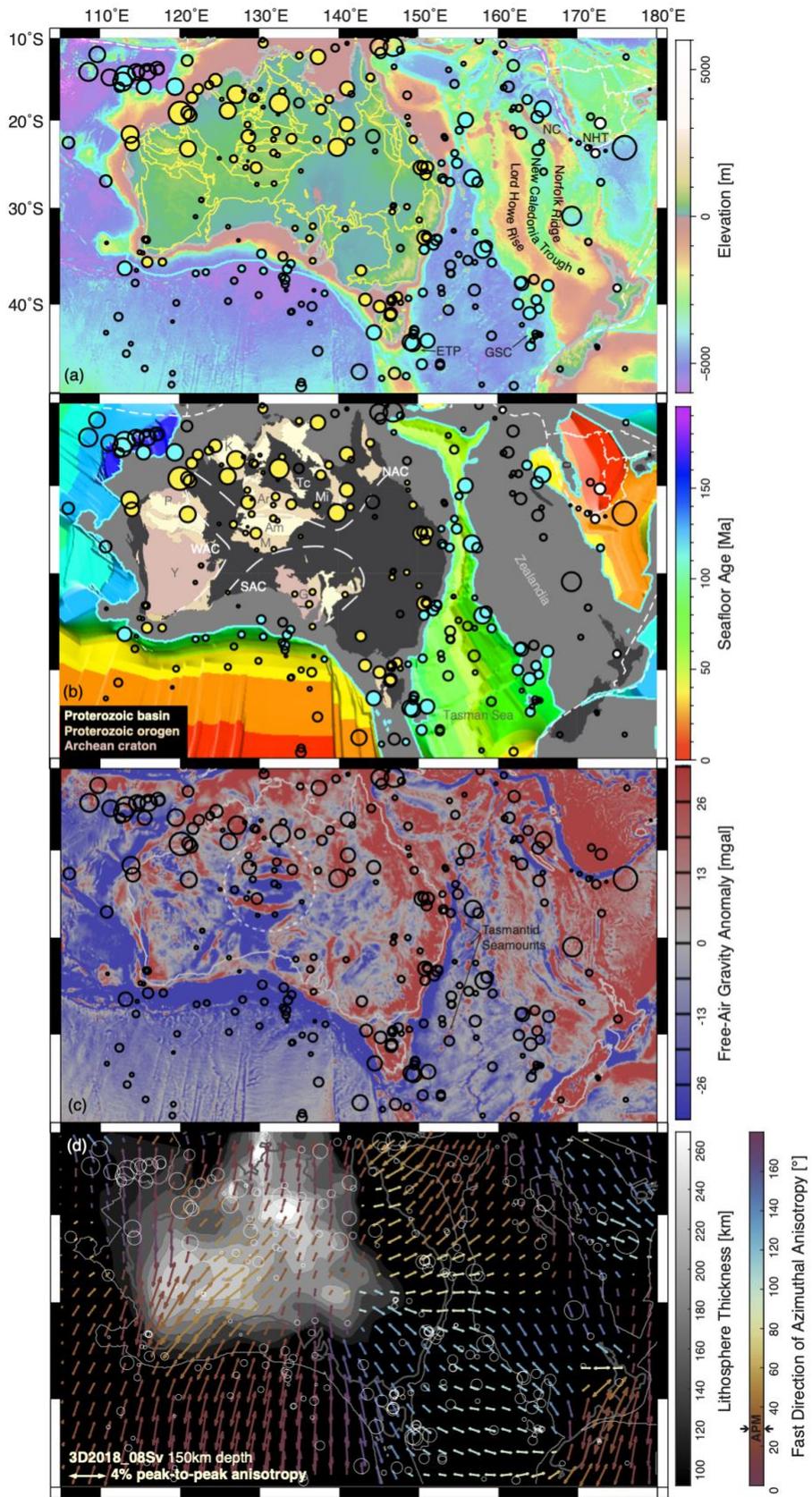
568 **Figure 1.** Process of QL wave scattering due to an anisotropic gradient. (a): Schematic
 569 illustration of a Love wave with transverse (T) particle motion (indicated in green), which upon
 570 encountering a lateral gradient in seismic anisotropy, partially converts into a QL wave with the
 571 same phase velocity and vertical (V) particle motion (indicated by blue) as a Rayleigh wave. The
 572 fundamental Love (G1) and Rayleigh (R1) waves travel a distance Δ between the source and
 573 receiver, whereas the QL wave only travels a distance δx between the scattering location and the
 574 receiver. The scattering is caused by a change in strength and/or orientation of the mantle
 575 anisotropy, indicated by the grey fabric pattern. (b): Example recording of a QL wave arrival that
 576 is found on the vertical component between G1 and R1. The distance to the scattering point (δx)
 577 is related to the time delay (δt) between the QL and G1 via the equation shown.



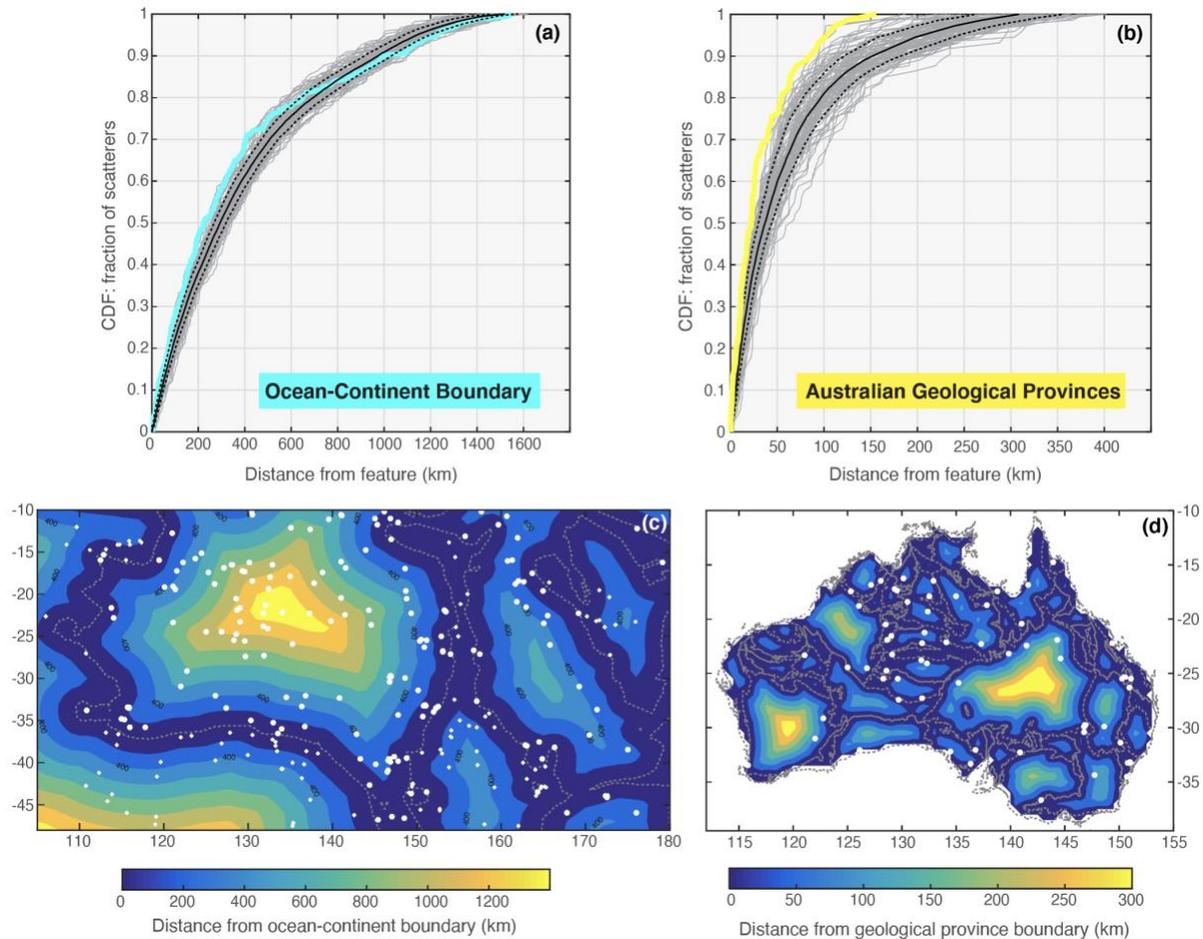
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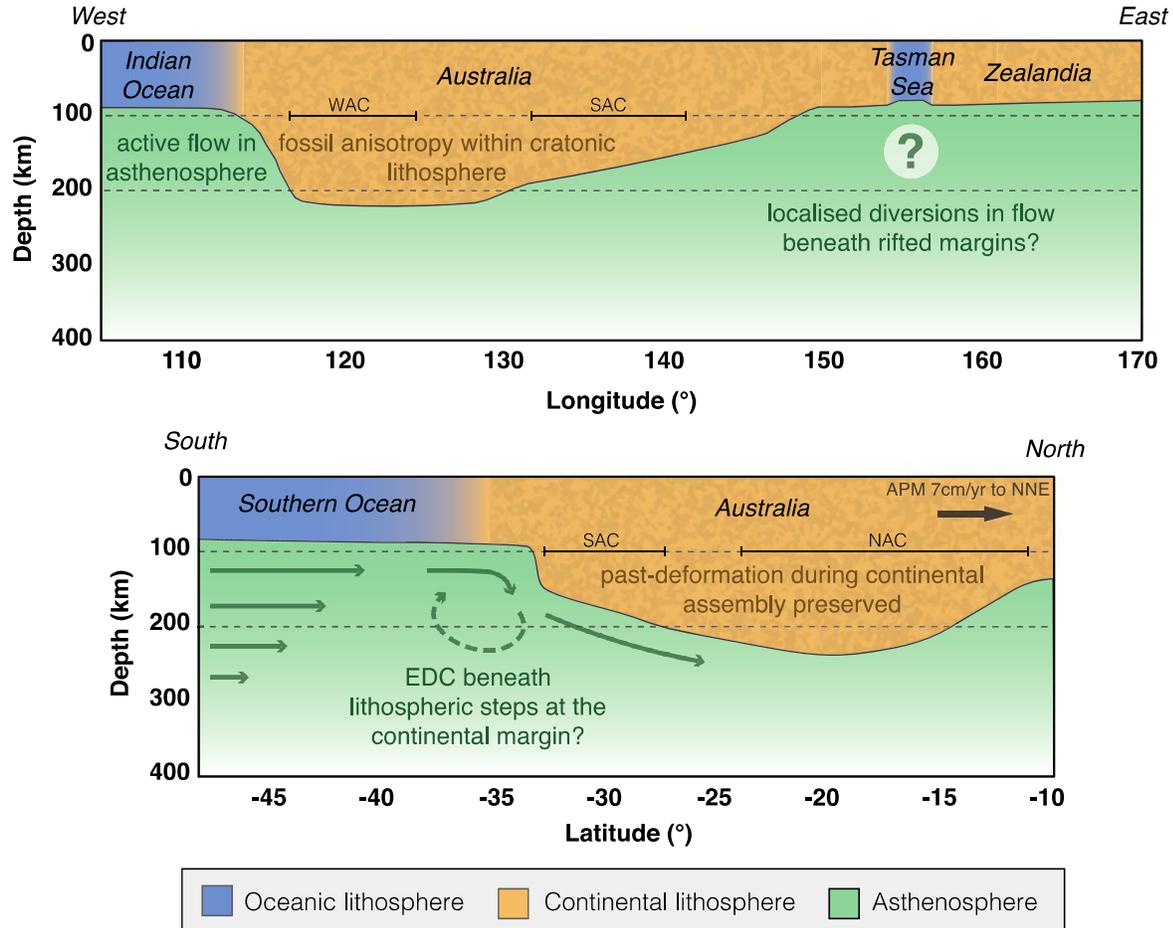
580 **Figure 2.** Global distribution of source-receiver paths analyzed for QL waves. Great-circle paths
581 on which QL waves were present are shown in blue, the remainder in grey. The location of
582 anisotropic gradients within the area of interest (black box, Fig. 3) is investigated using 43
583 seismic receivers (yellow triangles) based within Australia. Plate boundaries (in orange) from ⁵⁷.



585 **Figure 3.** QL wave scattering points (black circles) compared to (a): surface topography ⁵⁸, (b):
586 seafloor age ²⁸, the ocean-continent boundary, and the Precambrian tectonic elements of
587 Australia ⁵⁹, (c): free-air gravity ⁶⁰, and (d): azimuthal anisotropy from 3D2018_08Sv ⁸ and the
588 lithospheric thickness from AuSREM ^{31,41}. Circles are scaled in size relative to the amplitude of
589 the QL wave. Those which coincide (within a typical error of 100 km) with either a plate
590 boundary (dashed white lines), the ocean-continent margin (solid cyan lines), or geological
591 provinces within Australia (yellow lines) ³³, are highlighted by a solid circle with the respective
592 color in (a) and (b). In (a) abbreviations are as follows: ETP- East Tasman Plateau, GSC- Gilbert
593 Seamount Complex, NC- New Caledonia, NHT- New Hebrides Trench. In (b) abbreviations
594 refer to major Precambrian provinces: Am- Amadeus, Ar- Arunta, G- Gawler, K- Kimberley, M-
595 Musgrave, Mi- Mount Isa, P- Pilbara, Tc- Tennant Creek, Y- Yilgarn. White dashed lines show
596 the approximate extent of the West Australian Craton (WAC), North Australian Craton (NAC),
597 and South Australian Craton (SAC). In (c) the large dashed white circle over central Australia
598 highlights the location of prominent E-W trending gravity anomalies. In (d) azimuthal anisotropy
599 is illustrated by the arrows, with colors indicating the fast direction orientation, and the length of
600 the arrow relative to the strength of anisotropy (refer to legend for scale).



601
 602 **Figure 4.** Distance of QL scatterers relative to associated features. (a-b): Cumulative distribution
 603 functions (CDF) showing the fraction of all observed QL scatters as a function of distance from
 604 the ocean-continent boundary²⁸ (cyan line) and major geological provinces (yellow line)³³. The
 605 same CDFs are generated for 100 sets of the equivalent number of points randomly located
 606 across the study area (thin grey lines), alongside their mean (solid black line) and one standard
 607 deviation (dotted black line). (c-d) Maps showing spatial distribution of QL scatterers (small
 608 white symbols) compared to contours of the distance from either the ocean-continent boundary
 609 (c) or the geological province boundaries (d).
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612

613 **Figure 5.** Pattern and sources of seismic anisotropy in the upper mantle as inferred from QL

614 scatterers. Deformational processes are illustrated via two generic cross-sections across the

615 Australian continent: (i) east-west profile (top), and (ii) north-south profile (bottom). The 100-

616 200 km zone of peak depth sensitivity for 100 second QL wave scattering is indicated by the

617 dashed horizontal lines. The topography and depth of the lithosphere-asthenosphere boundary is

618 based on profiles from AuSREM⁴¹. Abbreviations are as follows: APM- Absolute Plate Motion,

619 EDC- Edge Driven Convection, NAC- North Australian Craton, SAC- South Australian Craton,

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Figures

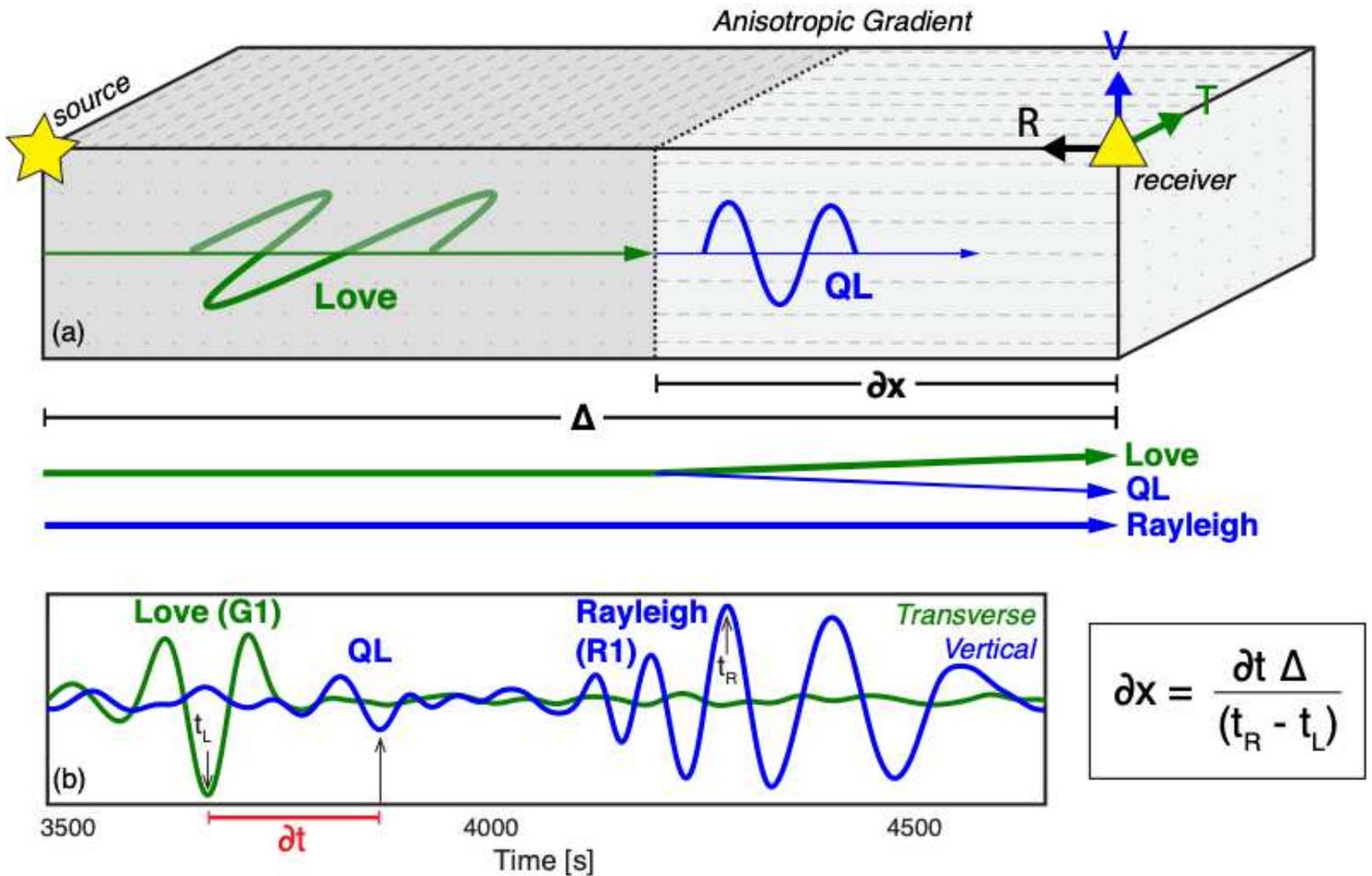


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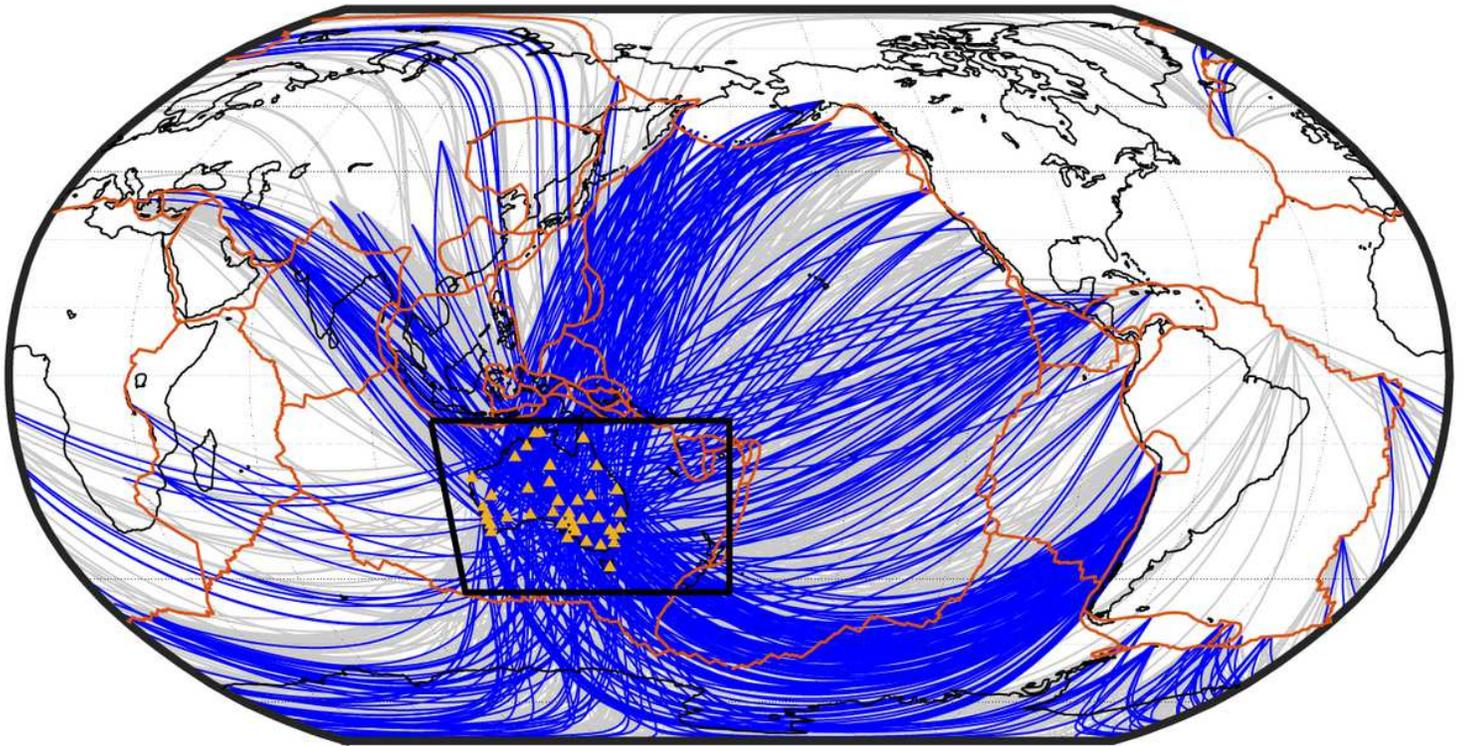


Figure 2

Global distribution of source-receiver paths analyzed for QL waves. Great-circle paths on which QL waves were present are shown in blue, the remainder in grey. The location of anisotropic gradients within the area of interest (black box, Fig. 3) is investigated using 43 seismic receivers (yellow triangles) based within Australia. Plate boundaries (in orange) from 57.

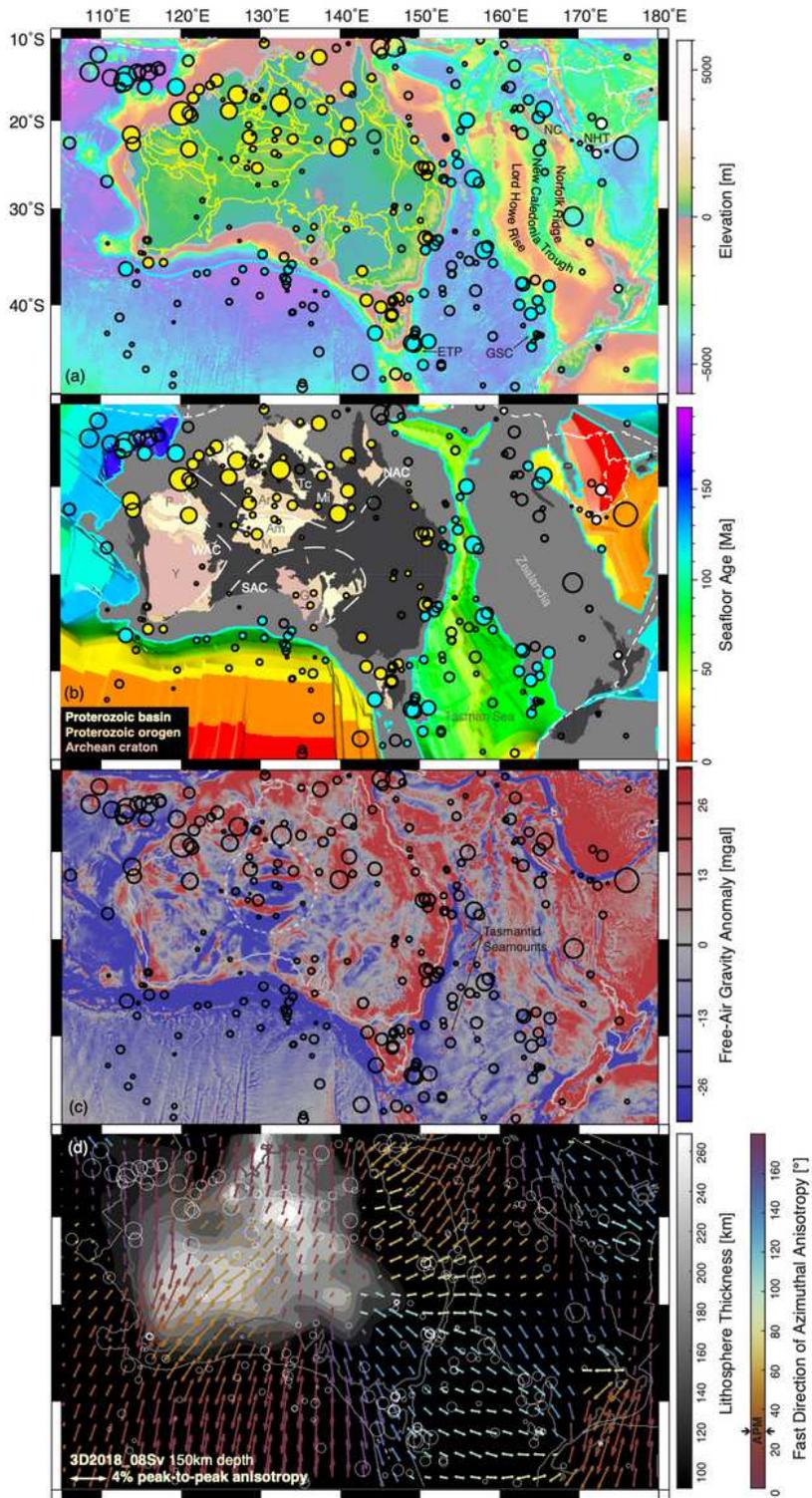


Figure 3

QL wave scattering points (black circles) compared to (a): surface topography 58, (b): seafloor age 28, the ocean-continent boundary, and the Precambrian tectonic elements of Australia 59, (c): free-air gravity 60, and (d): azimuthal anisotropy from 3D2018_08Sv 8 and the lithospheric thickness from AuSREM 31,41. Circles are scaled in size relative to the amplitude of the QL wave. Those which coincide (within a typical error of 100 km) with either a plate boundary (dashed white lines), the ocean-continent margin

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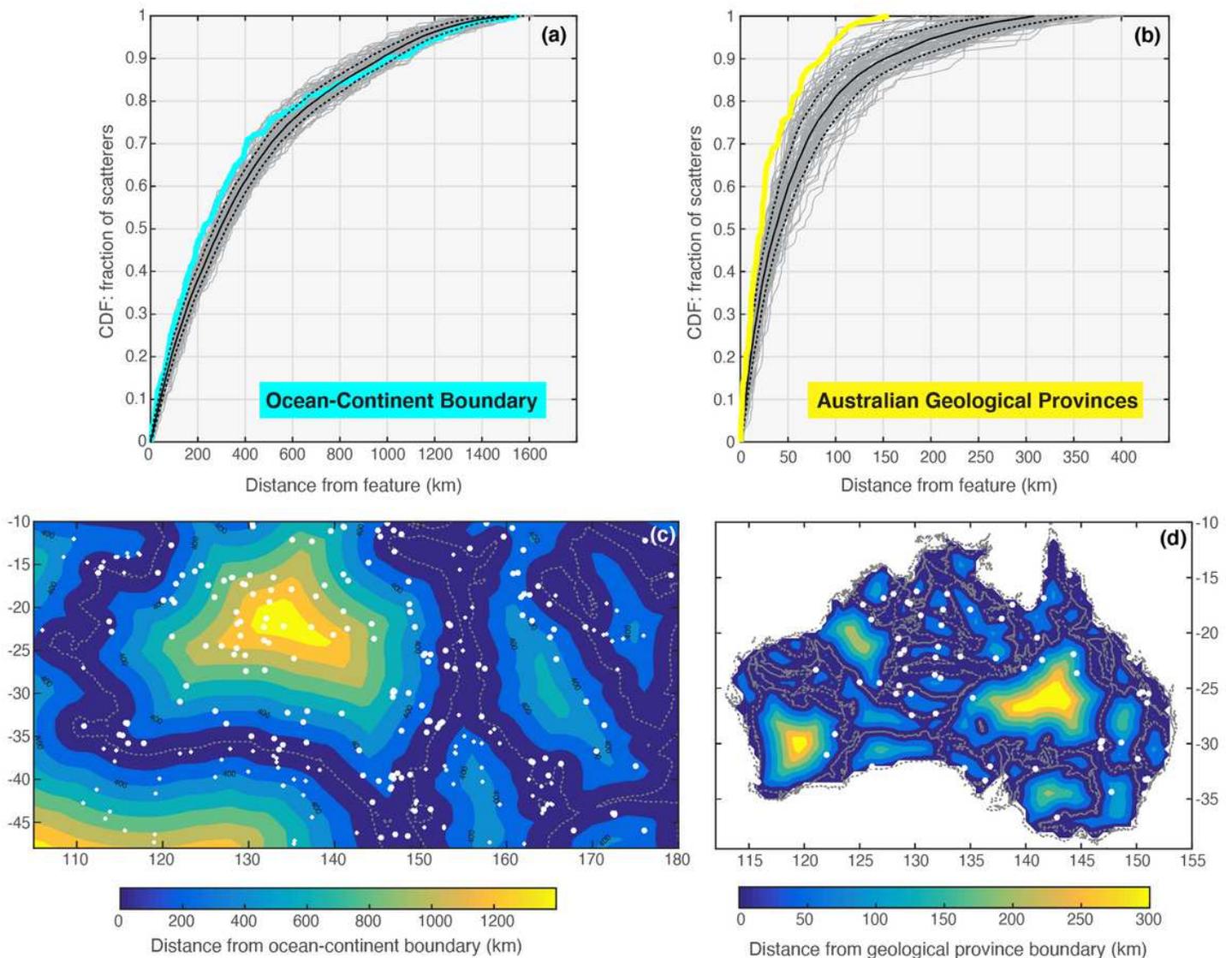


Figure 4

Distance of QL scatterers relative to associated features. (a-b): Cumulative distribution functions (CDF) showing the fraction of all observed QL scatters as a function of distance from the ocean-continent boundary 28 (cyan line) and major geological provinces (yellow line) 33. The same CDFs are generated

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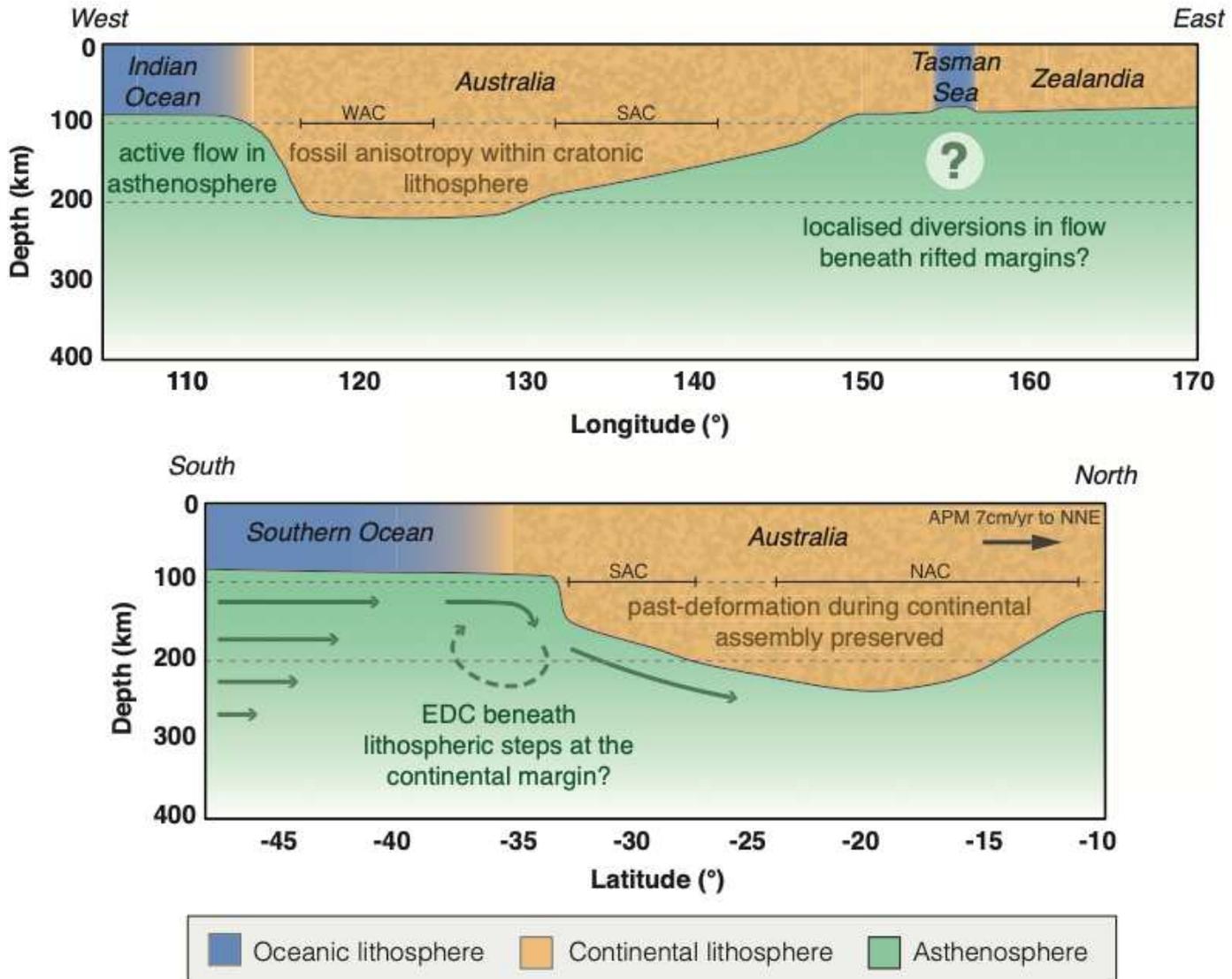


Figure 5

Pattern and sources of seismic anisotropy in the upper mantle as inferred from QL scatterers. Deformational processes are illustrated via two generic cross-sections across the Australian continent: (i) east-west profile (top), and (ii) north-south profile (bottom). The 100-200 km zone of peak depth sensitivity for 100 second QL wave scattering is indicated by the dashed horizontal lines. The topography and depth of the lithosphere-asthenosphere boundary is based on profiles from AuSREM 41. Abbreviations are as follows: APM- Absolute Plate Motion, EDC- Edge Driven Convection, NAC- North Australian Craton, SAC- South Australian Craton, WAC- West Australian Craton.

Supplementary Files

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