

# Outer Van-Allen Radiation Belt Response to VLF Plasma Waves: Application of Granger Causality Method to Radiation Belt Storm Probe-A Data

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## Research Article

**Keywords:** Granger causality analysis, Van-Allen radiation belts, very-low-frequency (VLF) waves, relativistic electrons flux, wave-particle interactions, radiation belt storm probes (RBSP) data.

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# 1 Outer Van-Allen Radiation Belt Response to VLF Plasma Waves:

## 2 Application of Granger Causality Method to Radiation Belt

### 3 Storm Probe-A Data

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#### 6 Abstract

7 Granger causality analysis is applied to very-low-frequency (VLF) waves and relativistic  
8 electrons interaction in an outer Van-Allen radiation belt. The VLF wave power and electrons  
9 flux data are one hour observations of radiation belt storm probe A (RBSP-A) during 14-3-  
10 2015. The considered VLF band comprises of 24 waves and electrons flux data at 8 pitch  
11 angles is considered. It is found that every wave is not Granger causing flux gradients for all  
12 pitch angles; rather some VLF waves are more efficient than others. Specifically, flux  
13 gradients at 69 degrees pitch angles are Granger caused by 14 kHz VLF wave. Contrarily, for  
14 90 degrees pitch angles flux gradients are Granger caused by 6 waves. These results may be  
15 helpful for understanding the dynamics of outer radiation belt.

16 **Key Words:** Granger causality analysis, Van-Allen radiation belts, very-low-frequency  
17 (VLF) waves, relativistic electrons flux, wave-particle interactions, radiation belt storm  
18 probes (RBSP) data.

#### 19 I. Introduction

20 The plasma particles (electrons, protons) trapping requires gradients parallel to the direction  
21 of magnetic field. This condition is naturally fulfilled by Earth's magnetic field, which is  
22 dipole-like up to few L-shells. Here L-shell represents the geocentric distance of a point

23 where magnetic field line intersects geomagnetic equator (that is 11 degrees tilted with  
24 respect to its geographic equator, for example  $L=2.5$  means a point in magnetosphere that is  
25 lying on magnetic equator and is at distance of 2.5 Earth's radii from centre of Earth). In  
26 1958, the first space mission (Explorer spacecraft) discovered two regions of trapped plasma  
27 in Earth's magnetosphere. These are known as Van-Allen radiation belts [1-2]. The inner  
28 radiation belt exists between  $L$ -shell= 1.5 – 3 and is mostly dominated by protons having  
29 energies 0.1 MeV – 40 MeV. The outer radiation belt occupies the region  $4 < L$ -shell  $\lesssim 6.0$   
30 and its electrons have energies in keV – MeV range. These radiation belts are separated by a  
31 slot region ( $L$ -shell=3-4), that is devoid of high intensity plasma particle radiations under  
32 normal conditions [3-4]. Electron fluxes in the outer radiation belt fluctuate by several orders  
33 of magnitude but its dynamics are yet not fully understood. Some intense storms can lead to  
34 the formation of a temporary third radiation belt (also known as storage ring) in slot region,  
35 which can survive for four weeks [5-6].

36 The space missions have confirmed the existence of many plasma waves in Earth's  
37 magnetosphere [7-12]. These include very-low-frequency (VLF) waves, which are whistler  
38 waves in frequency range 3-30 kHz [13-18], ultra-low-frequency (ULF) waves that are  
39 Alfvén waves, electromagnetic ion cyclotron (EMIC) waves in vast frequency range (1Hz –  
40 1mHz), their periods range from 1- 1000 seconds [19-20]. The VLF waves can be divided  
41 into two categories: upper band chorus (having frequencies  $\leq 0.5\omega_c$  (cyclotron frequency)),  
42 and upper band chorus waves (with frequencies  $> 0.5\omega_c$ ) [21].

43 The plasma particle sources of radiation belts include wave-particle interactions leading to  
44 radially inward diffusion of electrons. Wave-particle interactions also act as loss mechanism  
45 either by pitch angle scattering, or outward radial diffusion [22-29].

46 Many studies have focused on the wave-particle interactions in Van-Allen radiation belts and  
47 third radiation belt formation in slot region [26-29]. But the studies so far have not yet  
48 applied the Granger causality method to understand the relationship of VLF waves and outer  
49 radiation belt electrons pitch angle resolved flux. Therefore, this study for the first time  
50 applies Granger causality method to radiation belt storm probe A (RBSP) data of 24 VLF  
51 waves and 8 pitch angles resolved electrons flux. Using Granger causality method, this work  
52 finds answers to two basic research questions. The first research question states, do all 24  
53 VLF waves in 10-30 kHz range have similar effects on outer radiation belt flux at pitch  
54 angles 16-90 degrees? The second research question states that, can VLF waves in frequency  
55 range 10-30 kHz non-resonantly affect the electrons flux? The answers for the first and  
56 second research questions are no, and yes respectively. The manuscript is organized as  
57 follows, section II is about data and methods. Section III describes the VLF plasma waves  
58 and electrons flux observations of radiation belt storm probe A (RBSP-A) for one hour.  
59 Section IV is devoted to the discussion of results by applying Granger Causality method. The  
60 summary of main findings is presented in section V.

61

## 62 **II. Data and Methods**

63 The VLF waves and electrons flux data analysed in this work are one hour observations of  
64 radiation belt storm probe A (RBSP-A) during 00:00:10.148 - 00:59:57.914, on 14-March-  
65 2015. The electrons flux data at 2 MeV energy and varying pitch angles from 16-90 degrees  
66 pitch angles and very-low-frequency (VLF) electric field power density data at 10 -30 kHz  
67 frequencies is analysed.

68

69 The flux data is from relativistic electron proton telescope (REPT), basically from NASA  
70 cdaweb (<https://cdaweb.gsfc.nasa.gov/>) at RBSP/ECT REPT pitch angle resolved electron

71 and proton fluxes with electron energies, 2-59.45 MeV with a logical source rbspa\_rel03\_ect-  
72 rept-13. The time resolution of data values is varying from 10-11 seconds. The VLF waves  
73 power data comes from radio and plasma wave instrument (EMFISIS-L2, electric and  
74 magnetic field instrument level 2) with time resolution 6 seconds at <http://rbps-ect.sr.unh.edu>.

75

76 The data file for Granger analysis is produced by selecting only those VLF power values  
77 which are almost simultaneous with the flux data. The gradients of flux are calculated by  
78 subtracting previous value of flux from its current value. A similar procedure is applied for  
79 calculating power density gradients in the time series for each of 24 VLF waves. The initial  
80 values of flux gradient and VLF power gradients are taken zero.

81

82 This work applies Granger causality tests to study the cause and effect relationship of 24  
83 waves in very-low-frequency (VLF) band and 2 MeV relativistic electrons of outer Van-  
84 Allen radiation at 8 types of pitch angles.

85

86 The Granger causality test is a statistical procedure of hypothesis testing. There are two types  
87 of hypothesis: null and alternate. For this study, the null hypothesis states that temporal  
88 gradients of electrons flux time series at given pitch angle are independent of the temporal  
89 gradients in the power density time series of very-low-frequency electric field. Contrarily, the  
90 alternate hypothesis states that the temporal gradients of electrons flux time series are caused  
91 by the temporal gradients of VLF electric field power density gradients. Basically, there is  
92 not a single procedure to determine Granger causality. Different methods can be adopted to  
93 reject the null hypothesis and prove the causality of VLF waves and electrons flux. The  
94 lengthy details are avoided [already presented in Refs. 30, 31] here and a brief description is  
95 presented for readers. The causality judgement here is based on P-Value of F-statistics.

96 Basically, the formula of P-value reads as [30-31],

97 
$$P - value = 1 - F^{cdf} \quad (1).$$

98 The superscript *cdf* in equation (1) indicates cumulative probability. However, F is

99 calculated by the following relationship [30-31],

100 
$$F = \frac{\left[ \frac{SSR_r - SSR_u}{N_p} \right]}{\frac{SSR_u}{(N_0 - N_p - 1)}} \quad (2)$$

101 Here  $SSR_r$  is sum square of residuals of the restricted model and  $SSR_u$  represent sum square  
102 of residuals of the un-restricted model,  $N_p$  is the number of predictor variables and  $N_0$  is  
103 equal to the number of observations. This work takes a confidence interval of 95% which  
104 translates to a critical value of 0.05. When equation (1) returns a P-value less than 0.05, then  
105 null hypothesis is rejected and alternate hypothesis is accepted. Contrarily, when P-value is  
106 greater than 0.05 then null hypotheses is accepted and alternate hypothesis is rejected. Using  
107 this procedure the cause and effect relationships of the gradients in time series of VLF  
108 electric field power density and gradients in the time series of electrons flux are determined  
109 in figures (4)-(5) and explained in the discussion section IV.

110

### 111 **III. Results**

112 Figure 1 (a) displays the L-shell of radiation belt storm probe A with time. As the outer Van-  
113 Allen radiation belt exists between L = 4-6 and the L-shell of RBSP-A changed from L=5.8  
114 to 6.4, therefore, the observations are at the outer edge of outer Van-Allen radiation belt and  
115 the findings of this study are more significant for understanding the relativistic electrons  
116 dynamics in the presence of VLF wave turbulence in that region of Earth's magnetosphere  
117 near geostationary orbit (at L=6.6).

118 The electrons in Van-Allen radiation belts can have three adiabatic invariants [33]. This study  
119 does not concentrate on the first and third adiabatic invariant and the figure 1 (b) displays  
120 second adiabatic invariant (SAI) for electrons. The SAI in figure 1(b) varied from 0.034 to  
121 0.024. This indicates a 30 % reduction during the considered time interval. Based on this  
122 observation, SAI is not conserved for electrons. Under such conditions one cannot apply the  
123 adiabatic theory for finding the source of electrons flux changes. Although we have not  
124 provided the figure of first adiabatic invariant (FAI), but a simple argument indicates that  
125 FAI cannot be conserved. Because, the observed VLF waves frequencies (10-30 kHz) are  
126 comparable to the cyclotron frequencies of electrons (18.87 - 25.84 kHz) in the region of  
127 observations. This implies that the phase space density approach is also not applicable,  
128 because, it is only valid for conserved first adiabatic (FAI). Consequently, a new approach,  
129 Granger causality is adopted in this work to explore the influence of VLF waves on the  
130 relativistic electrons flux variations.

131 The power densities in 24 waves of very low frequency band are presented in the figures 2  
132 (a)-(c). In figure 2(a) power densities for 8 types of VLF waves at frequencies 10.0, 10.5,  
133 11.0, 11.5, 12.1, 12.7, 13.3, 14.0 kHz are presented. It shows that the highest power densities  
134 correspond to the 10 kHz waves and power density diminished as frequency increased,  
135 smallest power densities occurred for 14 kHz waves. The figure 2(b) displays the power  
136 densities for waves at 14.7, 15.4, 16.2, 17.0, 17.8, 18.7, 19.6, 20.5 kHz. Except 19.6, 20.5  
137 these VLF waves have frequencies smaller than electron cyclotron and some of these can be  
138 recognized as lower band chorus waves when the criteria  $f < 0.5f_c$  is satisfied. Here  $f$  and  
139  $f_c$  denote wave frequency and an electron cyclotron frequency respectively. The VLF waves  
140 in figure 2(c) at frequencies 21.5, 22.6, 23.7, 24.9, 26.1, 27.4, 28.7, 30.1 kHz are upper band  
141 chorus waves, because these satisfy the criteria  $f > 0.5f_c$ .

142 Figure 3 (a) shows the scatter plot of the 2 MeV relativistic electrons flux at 8 pitch angles,  
143 16, 26, 37, 47, 58, 69, 79, and 90 degrees. The flux is decreasing with the time for electrons  
144 at these considered pitch angles. To highlight extra information, figure 3 (b) presents a  
145 contour plot of the 2 MeV relativistic electrons with time and pitch angles. The highest flux  
146 peaks occurred for 37 degrees and 47 degrees pitch angles. Temporal trend shows that flux at  
147 all pitch angles is diminishing with time. The presence of VLF waves in figures 2 and flux  
148 changes in figure (3) can be linked together. One can assume that flux changes in figure (3)  
149 are due to VLF wave interactions with electrons. Such possibilities are explained in the  
150 following section.

151

#### 152 **IV. Discussion of Results:**

153 Radiation belt storm probe A (RBSP-A) simultaneously observed VLF waves (figure 2) and  
154 relativistic electrons flux reduction (figure 3), this implies that the flux changes are triggered  
155 by VLF waves. However, there are some basic research questions, such as, are the electron  
156 flux changes at 16 degrees pitch angles are caused by all the 24 waves or by only one wave?  
157 Similar, research questions can be asked for electrons flux changes at other pitch angles and  
158 combinations of waves. The RBSP-A observations in figures (2) and (3) cannot provide  
159 answers to such research questions. Therefore, further analysis of data in figures (2) and (3) is  
160 required. This work applies the Granger causality tests to explore the roles of considered VLF  
161 waves. The critical value 0.05 is used for judgement. When P-value is smaller than 0.05, then  
162 we reject the null hypothesis and alternate hypothesis is accepted. Otherwise, when P-value is  
163 greater than 0.05, then null hypothesis is accepted and alternate hypothesis is rejected. In this  
164 study, the null hypothesis states, the electron flux gradients at pitch angle (= 16, 26, 37, 47,  
165 58, 69, 79, 90 degrees) are independent of the gradients in powers of VLF wave (10, 10.5,

166 11.0, 11.5, 12.1, 12.7, 13.3, 14.0, 14.7, 15.4, 16.2, 17.0, 17.8, 18.7, 19.6, 20.5, 21.5, 22.6,  
167 23.7, 24.9, 26.1, 27.4, 28.7, 30.1 kHz). The alternate hypothesis states that flux gradients are  
168 dependent on the gradients of the wave power. These hypotheses are tested with Granger  
169 causality test and relationships between waves and flux changes at considered frequencies,  
170 and pitch angles are determined. Figure 4 shows P-values of Granger causality tests at 4-lags  
171 for the 2 MeV electrons at pitch angles 58 degrees (4a), 69 degrees (4b), 79 degrees (4c), and  
172 90 degrees (4d). In figures 4 (a) –(d) the horizontal red line shows a critical level of 0.05 and  
173 any point that lies under this level leads to rejection of null hypotheses and all points above it,  
174 accepts the null hypothesis. The P-values in figure 4(a) show that 8 waves Granger caused the  
175 flux changes for 58 degrees pitch angle electrons. Contrarily, gradients in power of only one  
176 wave (at 14 kHz) Granger caused (resulted in P-values less than 0.05 for first and second lag)  
177 flux gradients at 69 degrees pitch angles, as shown in figure 4(b). The flux changes in figure  
178 4 (c) for 79 degrees pitch angle electrons are Granger caused by two waves. Figure 4 (d)  
179 shows that power gradients of 6 waves Granger caused flux variations of 90 degrees pitch  
180 angle electrons. Figure 5 shows that the flux gradients of 16 (figure 5a), 26 (figure 5b), 37  
181 (figure 5c), and 47 degrees (figure 5d) pitch angle electrons are Granger caused by 2, 5, 3,  
182 and 5 waves respectively. The results in figures (4) and (5) also suggest that VLF waves also  
183 interact non-resonantly with outer Van-Allen radiation belt electrons, because, all these  
184 waves does not satisfy the cyclotron resonance condition. For example, consider the case of  
185 90 degrees pitch angle in figure 4 (d). In this case, 18.87 kHz VLF wave frequency is  
186 comparable to electron cyclotron frequency (18.87 - 25.84 kHz) in region of observation,  
187 therefore, the flux changes due 18.87 kHz VLF wave can be recognized as cyclotron resonant  
188 effect. Because, the cyclotron resonance condition reads [32-33],

189 
$$\omega - kV_{parallel} = n\omega_c \quad (3).$$

190 For 90 degrees pitch angle, parallel velocity  $V_{Parll} = 0$ , and resonance condition (1) reduces  
191 to  $\omega = \omega_c$  when  $n = 1$ . The effects of 20.5, 23.7 kHz waves can also be attributed to  
192 cyclotron resonance for  $n = 1$  because, these waves are bounded in the domain of electron  
193 cyclotron frequencies (18.87 - 25.84 kHz). However, the VLF waves at 12.1, 27.4 kHz and  
194 28.7 kHz does not satisfy the resonance condition for any  $n (= \pm 1, \pm 2, \pm 3, \dots)$  and therefore,  
195 these waves non-resonantly Granger caused flux gradients for the 90 degrees pitch angle  
196 electrons.

## 197 **V. Summary and Conclusions:**

198 The radiation belt storm probe (RBSP-A) simultaneous observations of 24 VLF waves and  
199 diminishing fluxes of 2 MeV electrons at 8 pitch angles suggest that flux changes may be due  
200 to VLF waves. By applying, the Granger causality tests, we find an answer to the basic  
201 research question, does all 24 VLF waves equally contributed to electrons flux changes? The  
202 answer is that all these waves do not contribute equally to electron flux changes in outer  
203 radiation belt at the considered pitch angles. This work also answers another research  
204 question, does VLF wave only Granger caused flux gradients when cyclotron resonance  
205 condition is satisfied. It is found that in some cases flux changes are Granger caused by VLF  
206 waves satisfying cyclotron resonance condition but not always. This implies that VLF waves  
207 also interacted non-resonantly with electrons and caused flux changes. The findings of this  
208 study can be helpful for understanding the dynamics of outer radiation belt.

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210 measured electrons flux data. We are also thankful to C. Kletzing for EMFISIS-L2 data of  
211 VLF wave electric field power density data.

212 **Data availability statement:** The data is available upon request to corresponding author at  
213 [masifkf@gmail.com](mailto:masifkf@gmail.com).

214 **Conflict of Interest:** The author has no conflict of interest.

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## 312 **Figure Captions**

313 **Figure 1** shows (a) L-shell of radiation belt storm probe A (RBSP-A) during one hour  
314 observations (00:00:10.148 - 00:59:57.914) on 14-March-2015. (b) Exhibits temporal  
315 variations in the electrons second adiabatic invariant.

316 **Figure 2** shows very- low-frequency (VLF) waves electric field power densities for (a) eight  
317 waves at 10.0, 10.5, 11.0, 11.5, 12.1, 12.7, 13.3, 14.0 kHz (b) eight waves at 14.7, 15.4,  
318 16.2, 17.0, 17.8, 18.7, 19.6, 20.5 kHz (c) eight waves at 21.5, 22.6, 23.7, 24.9, 26.1, 27.4,  
319 28.7, 30.1 kHz. These are observations of RBSP-A for same time interval as in figure 1.

320 **Figure 3** shows (a) scatter plot of 2 MeV energy electrons pitch angle resolved flux at eight  
321 pitch angles, 16, 26, 37, 47, 58, 69, 79, 90 degrees respectively (b) displays contour plot of  
322 the 2 MeV electrons flux with pitch angle and time. The flux measurements are based on the  
323 relativistic electron proton telescope (REPT) on board RBSP-A during one hour observations  
324 (00:00:10.148 - 00:59:57.914) on 14-March-2015.

325 **Figure 4** shows the P-values of Granger causality test, determined for gradients in VLF  
326 waves powers and gradients of electrons flux at pitch angles (a) 58 degrees (b) 69 degrees (c)  
327 79 degrees, and (d) 90 degrees. When P-value is less than critical value 0.05 (highlighted  
328 with red horizontal line) then null hypothesis is rejected and flux gradients are Granger  
329 caused by power gradients of VLF waves. Otherwise, null hypothesis is accepted and flux  
330 gradients are thought to be independent of power gradients of VLF wave.

331 **Figure 5** shows the P-values of Granger causality test determined for gradients in VLF waves  
332 powers and gradients of electrons flux at pitch angles (a) 16 degrees (b) 26 degrees (c) 37  
333 degrees, and (d) 47 degrees.

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Figure 1

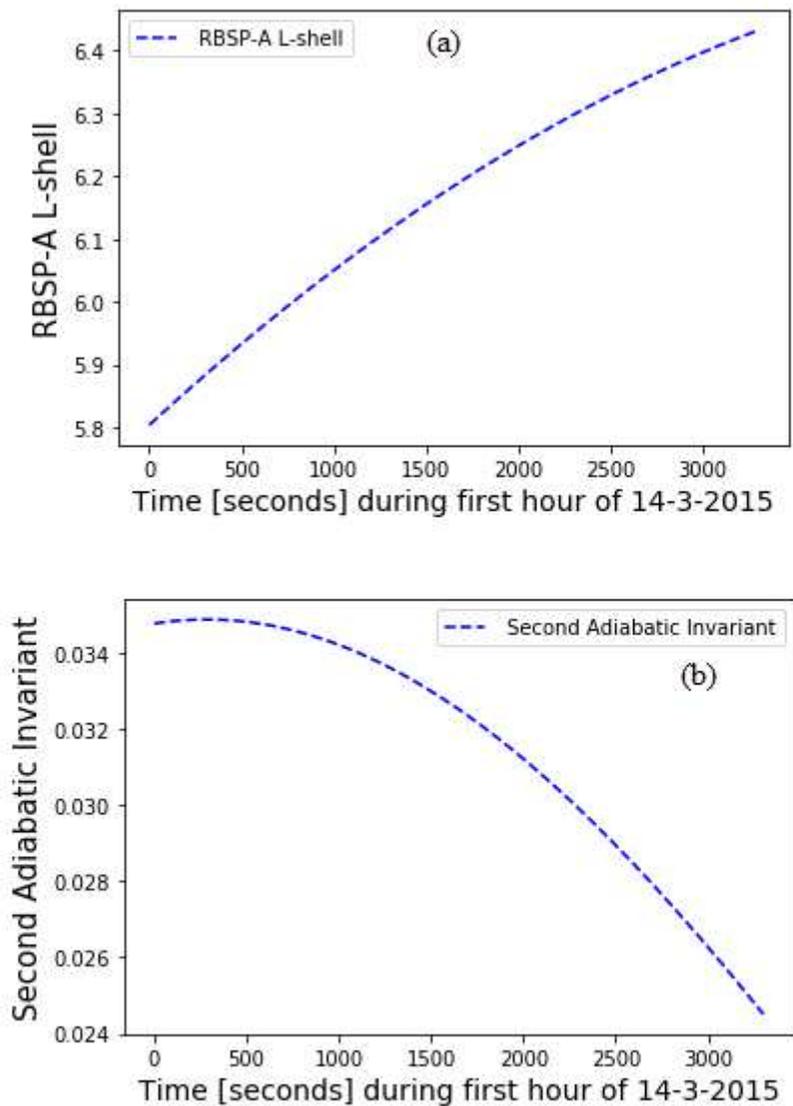


Figure 2

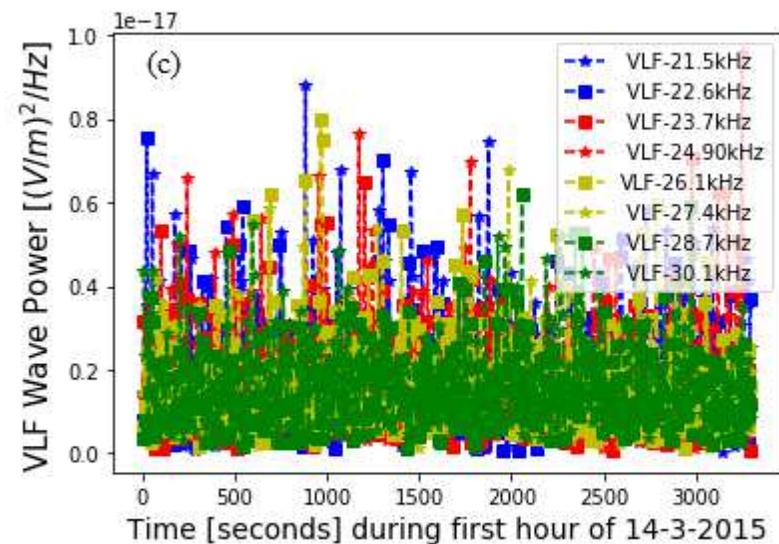
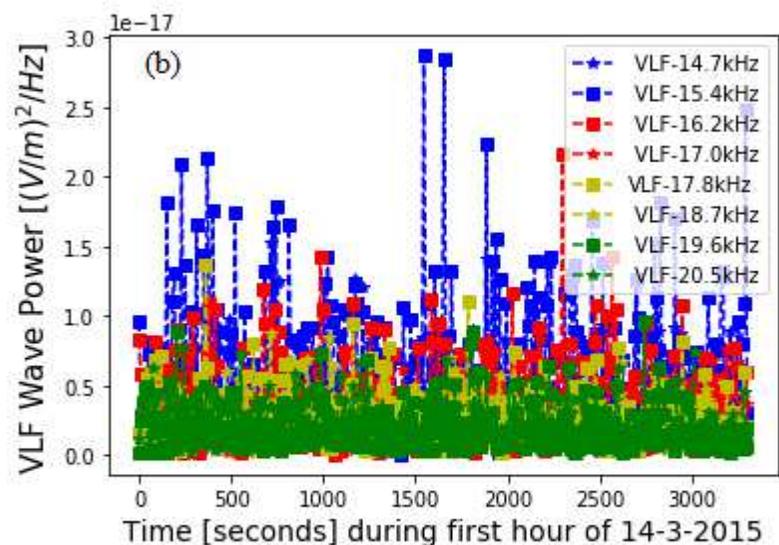
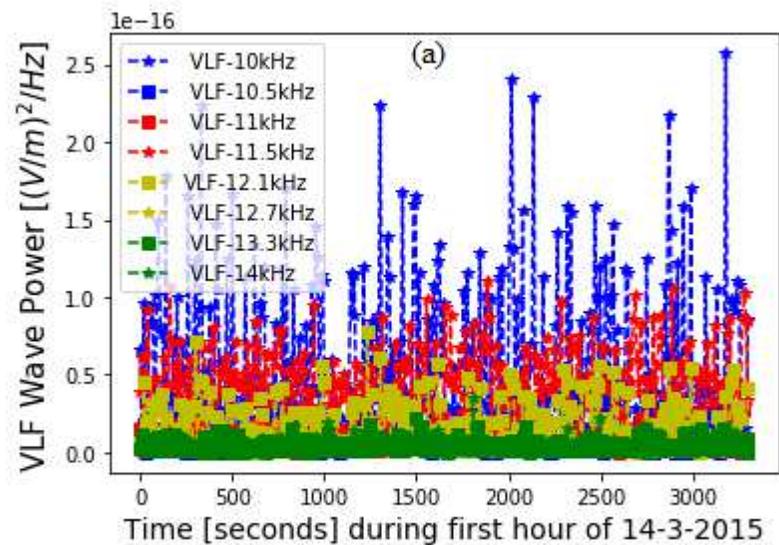


Figure 3

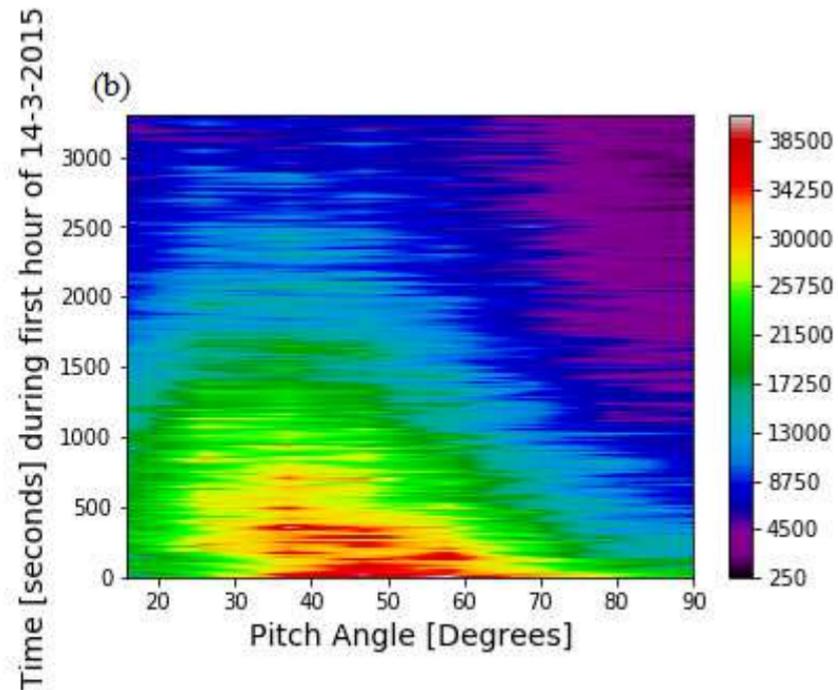
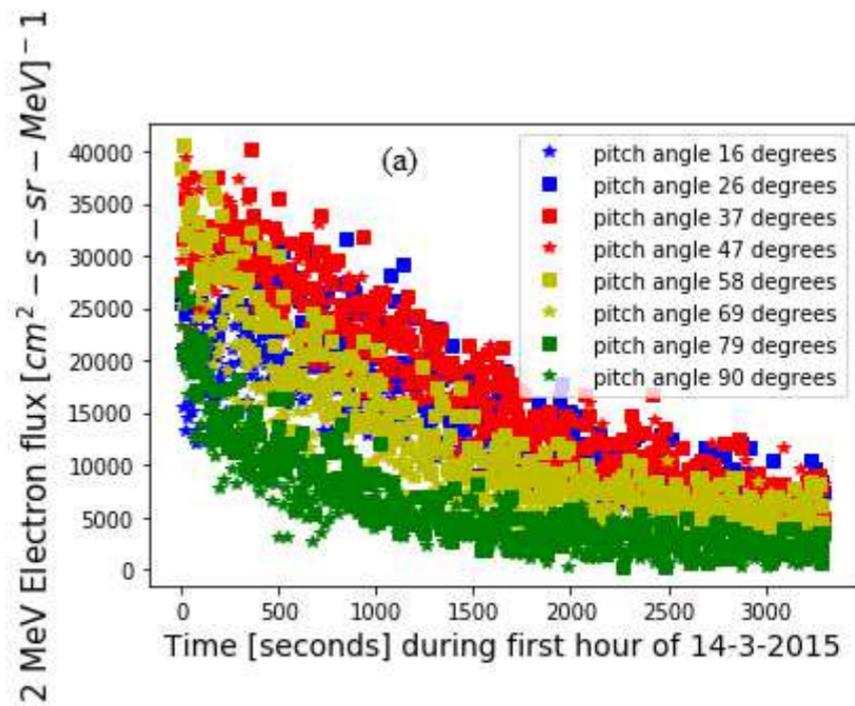


Figure 4

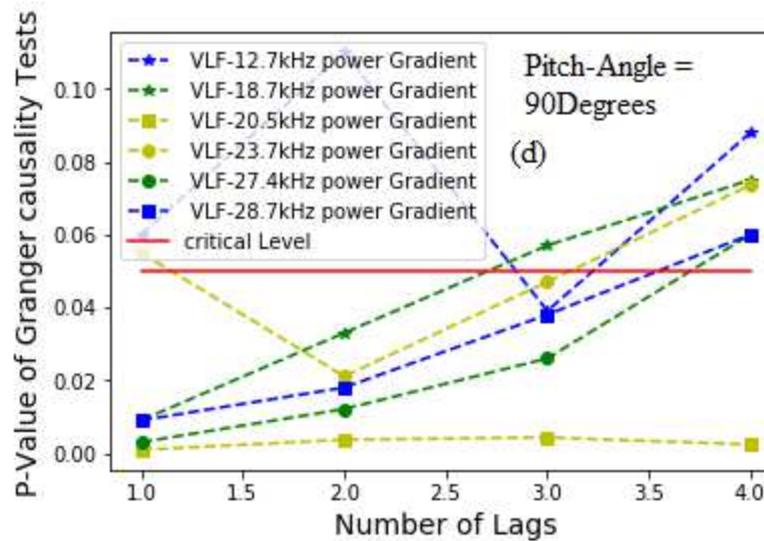
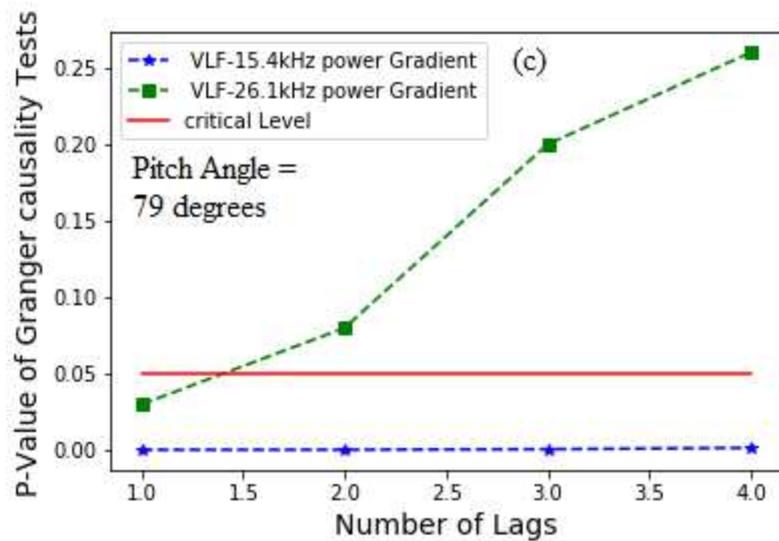
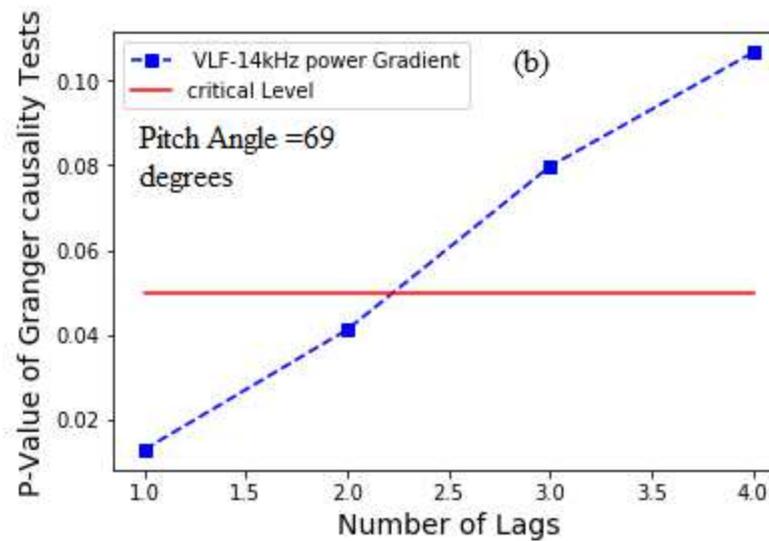
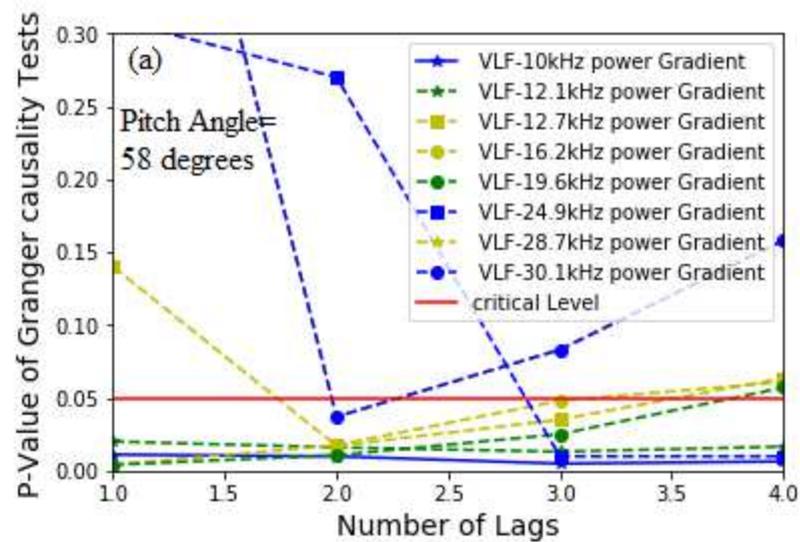


Figure 5

