

# Theoretical and experimental analysis of circularly polarized luminescence spectrophotometers for artifact-free measurements using a single CCD camera.

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## Article

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# Theoretical and experimental analysis of circularly polarized luminescence spectrophotometers for artifact-free measurements using a single CCD camera.

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## Abstract

We demonstrate the feasibility of artifact-free measurements of Circularly Polarised Luminescence (CPL) spectra on an imaging spectrometer. The polarization is spatially separated on two measurement channels using an achromatic quarter-wave plate associated with a polarizing beam splitter. The two channels are then imaged on a spectrometer to give the left and right circularly polarized spectra. The artifacts have been highlighted via the Mueller formalism and verified experimentally. As they add spectral contributions from either the shape of the luminescence or its derivative, they are almost impossible to eliminate by pre-calibration. We show that, at the cost of a second measurement that reverses the role of the two measuring arms, these artifacts are reduced in proportion to the stability of the lamp. The experimental results show the ability of our setup to measure CPL spectra on different kind of molecules in an accurate way with a dramatic (3 orders of magnitude) reduction in the recording time compared to a standard step by step measurement.

The interactions between circularly polarized light and chiral material are unequal for left-handed and right-handed circular polarizations and are revealed by different phenomena such as optical rotation (refractive index difference), electronic circular dichroism (ECD) (UV-VIS absorption difference), vibrational VCD (IR absorption difference) and so on [1]. The luminescence from chiral luminophore is also asymmetric in term of handedness which lead to the Circular Polarized Luminescence  $CPL = \Delta I = I_L - I_R$  corresponding to the difference between the left handed circular polarization (LHCP) and the right handed circular polarization (RHCP) luminescence [2]. CPL spectra of the enantiomers are opposite and the intensity of the signal is quantified by the dissymmetry factor  $g_{lum} = 2(I_L - I_R)/(I_L + I_R)$ .  $g_{lum}$  is of the order of  $10^{-3}$  to  $10^{-5}$  for most organic molecules [3, 4]. CPL is a valuable complementary tool to the more standard ECD spectroscopy because it allows to access different electronic transitions and can be detected in non-transparent media [5, 6]. During the last years, we observe a craze for CPL because of the synthesis of molecules or chiral systems with  $g_{lum}$  higher than  $10^{-3}$  making CPL a valuable alternative for the detection of chiral environmental probes or to conceive devices producing circularly polarized light [6, 7, 8, 9, 10, 11, 12, 13]. Thus, the need for reliable, robust and fast CPL measurements is becoming crucial.

Most of the CPL spectrometers are based on a polarization modulation scheme: the polarization is modulated between the two LHCP and RHCP at some tens of KHz frequency with a photo-elastic modulator (PEM) [2]. The Fourier analysis of the signal allows to get the excess of circular polarization. The limit of detection is determined by the integration time of the lock-in amplifier used to demodulate the signal. This system requires a fast mono-channel detector (photo-multiplier or photodiode) with a time constant lower than the modulation period (20 $\mu$ s for common PEM). Thus, it prevents the use of CCD camera which have, generally, millisecond long integration times. The spectra are necessarily recorded step by step and the recording times may last several tens of minutes making this technique very heavy and limited to fundamental studies. This technique has proven its efficiency with measured spectra of  $g_{lum} \leq 10^{-4}$  and its validity and related artifacts have been well investigated [14].

In this work, we investigate the use of a CCD camera for a fast recording of full CPL spectra. Because, high speed modulation can not be used (except for non commercially available specialized camera [15]), CCD based spectra must rely on a spatial polarization separation scheme. Spatial beams polarization splitting is

| Defect   | Origin  | Parameter  | magnitude                          |
|--|---|--|------------------------------------|
| QWP retardance<br>$\Psi \neq \frac{\pi}{2}$              | QWP specifications<br>- chromaticity            | $\psi = \Psi - \frac{\pi}{2}$                                      | $\pm \frac{\pi}{50}$ over<br>500nm |
| QWP azimuth<br>$\theta \neq \pm \frac{\pi}{4}$           | Mis-alignment                                   | $\alpha = \theta - \frac{\pi}{4}$                                  | $\pm \frac{\pi}{200}$              |
| Unequal arms transmission:<br>$T_{+1} \neq T_{-1}$       | PBS, fibers, grating                            | $\Delta T = \frac{T_{+1} - T_{-1}}{T_{+1} + T_{-1}}$               | $10^{-2}$                          |
| Polarization mixing                                      | imperfect extinction ratio (PBS),<br>straylight | $\varepsilon_p, \varepsilon_s, \sigma_l$                           | $10^{-2}$ - $10^{-3}$              |
| Pump source variations:<br>$\phi_{+1} \neq \phi_{-1}$    | LED stability                                   | $\Delta\phi = \frac{\phi_{+1} - \phi_{-1}}{\phi_{+1} + \phi_{-1}}$ | $10^{-2}$ - $10^{-3}$              |
| Wavelength mismatch:<br>$\lambda_{+1} \neq \lambda_{-1}$ | monochromator misalignment                      | $\Delta\lambda = \frac{\lambda_{+1} - \lambda_{-1}}{2}$            | 0.1nm                              |

**Table 1:** The different defects of the setup limiting the experimental measurements, their origin, the parameters describing them and their magnitudes.

commonly used for full polarization characterization, i.e Mueller imaging. It is generally dedicated to single wavelength and/or high polarization dissymmetry ( $\geq 10^{-2}$ ) phenomenon measurements, as complex procedures are required to calibrate each of the optical elements [16]. For measuring subtle spectral chiroptical signals, use of CCD camera after polarization beam splitting have been demonstrated for Raman optical activity [17] and CPL measurements [18, 19] and ourselves [20, 21]. The needed accuracy – of the order of 0.01% on the polarization difference – requires an equal control of the two analysis arms. This is almost impossible to achieve over the entire wavelength range and a very complex calibration procedure is proposed by Mackenzie [18]. Hug and Hangartner showed that by inverting the role of the two arms, the unequal arms transmission is auto-compensated [17]. Their setup was successfully commercialized by Biotools. On our side, we have also performed CPL measurements on a single camera by only separating the luminescence according to two polarization encoded optical paths [21]. However, even-though the use of only one camera reduces the artefacts sources, appropriate calibration procedure is still needed to avoid remaining defects and inaccurate signals recording.

In this study, we show that, at the cost of two successive measurements, corresponding to the swap of the two optical paths, and then their average, reliable spectra of CPL can be obtained. The performances and limitations of this method are fully investigated. The article is organized as follow. In the first part, after describing the experimental setup, we theoretically analyze the recorded spectra via the Mueller matrix framework. Three configurations are investigated: simultaneous measurements on two arms (spatial separation of polarization), successive measurements on one arm (time polarization's separation) and combination of the two previous ones (space-time combination of polarization's separation). First order calculations show that a first order false signal appears for spatial and time polarization's separations. These signals multiply together in the space-time combination and consequently the related artifacts become very low.

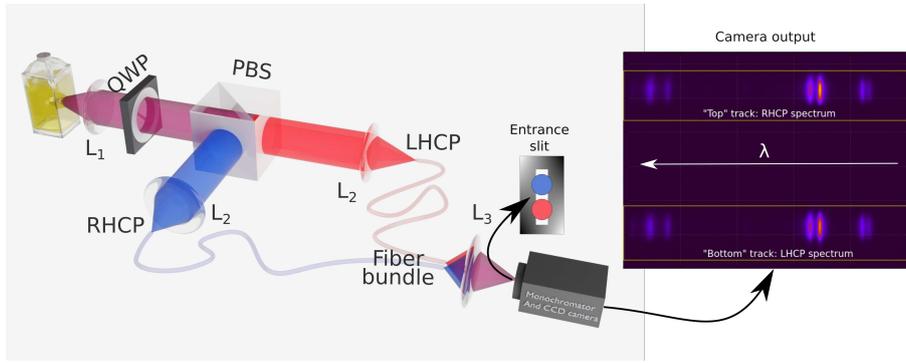
These theoretical finding are experimentally illustrated in the second part of the paper, on broad ( $\sim 50$  nm) and narrow line ( $\sim 1$  nm) spectra from camphorquinone and  $\text{Eu}^{3+}$  complexes, respectively. Moreover, in order to assess the limits of our method and to verify its performance, we compare the spectra obtained with the camera to those obtained with our "home-made" single channel PEM based setup described elsewhere [7, 22, 23].

This work shows that CPL free from first order artifacts can be recorded in a few seconds over a wavelength range between 300 and 1050 nm with a multi-channel coverage of 150-300 nm depending on the grating and optical magnification. The only limitation is the analog to digital converter of the CCD which gives a limit of detection  $g_{lum} \sim 5.10^{-4}$ . This works opens perspectives on fast/automatic measurement of CPL in automated processes.

## 1 Optical set-up

### 1.1 Single CCD-based CPL apparatus

The home-made single CCD based CPL spectrometer is schematically represented in Figure 1. Basically, the handedness of the circularly polarized luminescence is spatially encoded into two geometrical paths before



**Figure 1: Schematic of the CCD-based CPL spectrometer:** The fluorescence light is collected via lens  $L_1$ . After the quarter-waveplate (QWP, main axis  $45^\circ$ ) and the polarizing beam splitter (PBS, main axis  $0^\circ$ ) the LHCP and RHCP are transmitted and reflected respectively (as horizontally and vertically polarized light). The two beams are launched into two fibers with a lens  $L_2$ , for each arm. These two fibers merge as a bundle. The bundle output is imaged via lens  $L_3$  as two vertically separated spots on the entrance slit of the monochromator (see SI, Sec 1). The two spots are horizontally diffracted by the grating giving on the CCD camera the top/bottom spectrum coming from the top/bottom spots and thus related to the LHCP and RHCP emission. Here, the image is recorded from an  $\text{Eu}^{3+}$  complex **2**.

81 being spectrally dispersed by the spectrometer and recorded on the CCD camera. This is accomplished by  
 82 the association of a quarter wave plate (QWP, fast axis  $45^\circ$ ) and a polarizing beam splitter (PBS): the left  
 83 handed circularly polarized (LHCP) light is transmitted by the PBS ( $0^\circ$  linear polarization) while the right  
 84 handed circularly polarized (RHCP) light is reflected by the PBS ( $90^\circ$  linear polarization). The two handedness-  
 85 encoded beams are then routed at the adjustable entrance slit of the spectrometer by means of a dual core  
 86 fiber bundle (200 $\mu\text{m}$  diameter). The two beams are spatially imaged as two vertically aligned spots on the  
 87 entrance slit. Inside the spectrometer (see SI, Sec 1 for details), the input slit is imaged on the CCD camera  
 88 (Andor iDus-420) after being horizontally diffracted by the grating leading to two vertically split spectra: the  
 89 top (bottom) one corresponds to the top(bottom) fiber and consequently to the RHCP (LHCP) luminescence.

90 Using a single spectrophotometer with only one CCD array camera avoids stability issues between two detec-  
 91 tors and reduces the price of the setup, by saving one spectrophotometer, as regard CPL setups based on two  
 92 of them [18, 19].

## 93 1.2 Three ways of measuring CPL

94 Here we experimentally investigate CPL recording using time, spatial and time-spatial polarization separation.  
 95 For this, the QWP is held in a computer-controlled rotating holder. The measurement first involves setting the  
 96 orientation of the QWP fast axis to  $+45^\circ$  to place the RHCP channel on the top track of the CCD and the  
 97 LHCP on the bottom track. These first two polarization-encoded spectra are recorded. In a second step, the  
 98 fast axis of the QWP is set to  $-45^\circ$  to invert the polarization channels and the two polarization-encoded channels  
 99 are recorded. This allows to measure the CPL along three ways:

100 **Spatial separation** :  $I_L$  and  $I_R$  simultaneously measured as the spectra coming from the two polarization encoded  
 101 paths for one given QWP orientation (Figure 2-a);

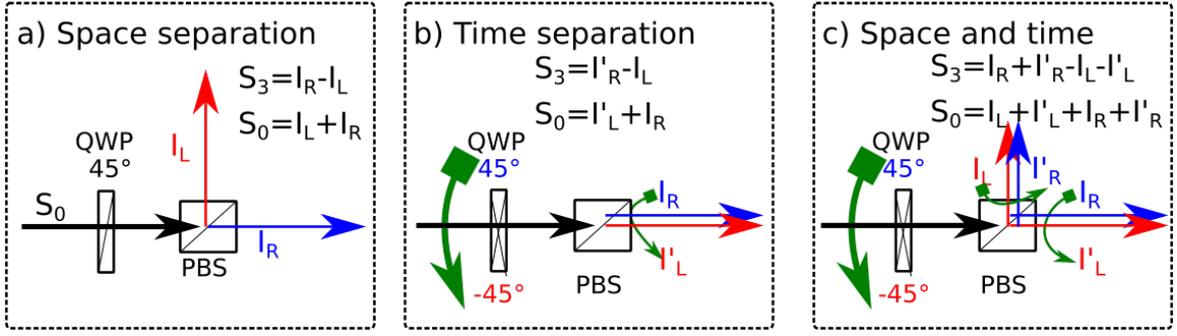
102 **Time separation** :  $I_L$  and  $I_R$  successively measured on one optical path for two QWP azimuth  $\pm \frac{\pi}{4}$  (Figure 2-  
 103 b);

104 **Spatial-Time combination** :  $I_L$  (resp.  $I_R$ ) is taken as the sum of the two left- (resp. right-) handed polarization  
 105 measurements for the two QWP orientations (Figure 2-c).

## 106 2 Theoretical analysis

### 107 Principle

108 Ideally, all the configurations illustrated in Figure 2 lead to the same true CPL spectra. However, since the  
 109 optical elements along the two paths and the pump source are not ideal, we must consider different experi-  
 110 mental limitations. The detailed steps to obtain the intensity of the recorded light are given in the Supporting  
 111 Informations. In here, we give only the major steps. Light is described using the 4-components of the Stokes



**Figure 2:** Three measurement's configurations to obtain CPL: Left and right circular polarisations encoded spectra are measured: (a) simultaneously through the two arms, (b) on the same arm after rotation of the QWP and (c) as the sum of the two previous ones.

| Polarization separation                   | Signals combination                             | ZO                                 | 1 <sup>rst</sup> non null correction  | False CPL signals                      |
|---|---|------------------------------------|---|--|
| Time: QWP (rotation) + PBS (One channel)  | $I_{p,+1} - I_{p,-1}$                           | $p(1 - \epsilon_p)F_3S_3(\lambda)$ | $\Delta\phi S_0$  | Lum.                                   |
| Spatial: QWP (fixed) + PBS (two channels) | $I_{+1,q} - I_{-1,q}$                           | $q\eta_{mix}F_3S_3(\lambda)$       | $(\Delta T + \frac{\epsilon_+ - \epsilon_-}{2})S_0$<br>$\Delta\lambda \frac{\partial S_0}{\partial \lambda}$<br>$\eta_{mix}F_1S_1 + \eta_{mix}F_2S_2$   | Lum.<br>Lum. derivative<br>Linear Lum. |
| Time + Space                              | $(I_{1,1} - I_{-1,1}) - (I_{1,-1} - I_{-1,-1})$ | $\eta_{mix}F_3S_3(\lambda)$        | $\eta_{mix}(\Delta\phi\Delta T + \frac{\epsilon_+ - \epsilon_-}{2})S_0$<br>$\eta_{mix}\Delta\phi\Delta\lambda \frac{\partial S_0}{\partial \lambda}$<br>$\eta_{mix}\Delta\phi(F_1S_1 + F_2S_2)$ | Lum.<br>Lum. derivative<br>Linear Lum. |

**Table 2:** CPL calculated at different orders of the experimental limitations using Equation 2 for three signal combinations. ZO is the result at zero order. First non nul order terms are in fourth column. The last column shows the corresponding unwanted spectral signals present in the main signal.

122 vector which are: the total intensity  $S_0 = I_L + I_R$ , the difference between the circularly polarized intensities  
 123  $S_3 = I_R - I_L$  (opposite to the CPL) and the differences between the linearly polarized intensities  $S_1 = I_{0^\circ} - I_{90^\circ}$   
 124 and  $S_2 = I_{45^\circ} - I_{-45^\circ}$ . References axes ( $0^\circ$  and  $90^\circ$ ) are chosen identical to the PBS ones.

125 Using the Mueller matrices of the different elements encountered along the optical paths, the Stokes vector  
 126 at the detector can be derived as the product of the different Mueller matrices  $\mathbf{M}_i$  representing each optical  
 127 element:

$$\prod_i \mathbf{M}_i \cdot \mathbf{S}(\lambda) \quad (1)$$

128 where  $\mathbf{S}(\lambda)$  is the Stokes vector of the emitted light.

129 We have calculated the four signals corresponding to the top and bottom arms measured twice with the  
 130 QWP azimuth at  $+\frac{\pi}{4}$  and at  $-\frac{\pi}{4}$ . Different experimental limitations have been taken into account: QWP phase  
 131 retardation  $\Psi \neq \pi/2$  and azimuth  $\theta \neq \pm\pi/4$ , unequal transmission of the two arms, spectral mismatch, cross-  
 132 talk between the two arms and pump instability. Several reasons make the overall transmission of the two arms  
 133 unequal: unbalance of the beamsplitter, unequal launching into the fibers, non-equal polarization response of  
 the grating, ... We have gathered all these imperfections in a transmission factor called  $T_{\pm 1}$  for the top/bottom  
 arms.

It is also important to consider the potential wavelength mismatch between the two spectra recorded simul-  
 taneously. It may come from a mis-alignment between the grating, the entrance slit and the CCD pixel matrix  
 and induces that the two intensities recorded on the same vertical channel of the CCD correspond to the two  
 slightly different wavelengths  $\lambda_{\pm 1}$  for the top/bottom arms.

PBS imperfections, as well as straylight inside the monochromator, induce cross-talk between the two po-  
 larization encoded channels: a small relative part of one polarization falls into the detection area of the other  
 one's.  $\epsilon_{+,-}$  denote the PBS inverse rejection rates and we assume that the straylight mixes a fraction  $\sigma_l$  from  
 one channel to the other one's.

134 Finally, because we measure spectra at two different times, a possible drift of the pump source may occur.  
 135 We model this pump instability as two exciting ratio  $\phi_{\pm 1}$  for the  $\pm \frac{\pi}{4}$  QWP azimuth orientation. All these  
 136 sources with their quantification parameters are gathered in Table 1 and the corresponding Mueller matrices are  
 137 described in SI section 5.

Taking into account all these effects, the spectra recorded for each arm following the two QWP azimuth orientations, write at first order:

$$I_{p,q} = \frac{\phi_q T_p}{2} \left[ (1 + \epsilon_p) S_0 + p(1 - \epsilon_p)(F_1 S_1 + F_2 S_2) + pq(1 - \epsilon_p) F_3 S_3 \right] (\lambda + p\Delta\lambda) \quad (2)$$

where the subscript  $p = \pm 1$  denotes the polarization channel after the PBS (+ and - for the horizontal and vertical polarization respectively) and the subscript  $q = \pm 1$  denotes the orientations of the QWP fast axis ( $\pm$  for  $\pm\pi/4$ ) and therefore the pumping efficiency  $\phi_{\pm 1}$  corresponding to the pump efficiency at the time of the measurements. The  $F_i$  coefficients take into account the departure of the polarization elements from the ideal case:

$$F_1 = \sin^2 2\alpha - \cos^2 2\alpha \sin \psi \approx -\psi \ll 1 \quad (3)$$

$$F_2 = -\sin 2\alpha \cos 2\alpha (1 + \sin \psi) \approx -2\alpha \ll 1 \quad (4)$$

$$F_3 = -\cos \psi \cos(2\alpha) \approx 1 \quad (5)$$

138 For each of the three configurations, combining equation 2 with the respective  $(q, p)$  measurements pairs allows  
 139 to derive the experimental luminescence (by summation) and CPL spectra (by differentiation) (SI sections 4.4-  
 140 7). These calculations are performed as a power development with  $S_0 \gg S_{i \neq 0}$  and the experimental defects taken  
 141 a small first order effect (i.e.  $\Delta T$ ,  $\Delta\phi$ ,  $\Delta\lambda\partial/\partial\lambda$ ,  $F_1$  and  $F_2 \ll 1$ ). The results are gathered in Table 2 for the  
 142 CPL, the luminescence is given in the SI. In Table 2, the measured signals are ordered following the zero and  
 143 the first non null terms of the power development.

144 As expected, the zero order corresponds to the expected signals, i.e. CPL to the nearest  $\eta_{mix} F_3$  factor where  
 145  $F_3$  denotes the QWP imperfection and :

$$\eta_{mix} = 1 - \frac{\epsilon_+ + \epsilon_-}{2} - 2\sigma_l \quad (6)$$

146 quantifies the reduction of the CPL coming from the mixing between the two arms. The higher order terms come  
 147 from the experimental limitations that introduce unwanted signals which mix with the good one's. First order  
 148 term in the luminescence signals are generally much smaller than the luminescence as they imply the product of  
 149 the small CPL and linear birefringence signals modulated by the instrumental defects. We can therefore, safely  
 150 assume that the sum signals are the luminescence ones'.

151 **Trueness of the CPL** At first order, the measured CPL is scaled down by two factors:  $F_3$  coming from the  
 152 QWP imperfections and  $(1 - \epsilon_p)$  or  $\eta_{mix}$  for the time and spatial separation schemes respectively. As a product  
 153 of cosinus of small angles (see Table 1), the  $F_3$  coefficient is very close to one. With usual alignment procedures  
 154 between to cross polarizers,  $F_3$  may be keep higher than 0.99 on a 400nm range with standard achromatic QWP.  
 155 CPL reduction coming from the PBS is about 1% and we measure in our home made monochromator a stray  
 156 light coefficient lower than 4% over the whole spectral range. It gives an overall reduction factor of 0.95 on the  
 157 measured CPL. The wavelength dependance of this factor is very smooth as it can be seen from retardance and  
 158 rejection rate specifications given by different suppliers. Our own measurements one rejection rates and stray  
 159 light validate this assumption (see SI Fig. 2). Therefore, we do not take into account this correction in the rest  
 160 of the paper.

161 **Time separation of the polarizations** is presented in the first row of Table 2. It induces a false CPL signal  
 162 ( $\Delta\phi S_0$ ) coming from the variation of the excitation source between the two  $I_L$  and  $I_R$  measurements. Indeed, a  
 163 pump variation between the two measurements cannot be discriminated from a true variation of  $I_L$  with respect  
 164 to  $I_R$ . This signal having the spectral shape of the luminescence can lead to erroneous interpretation. It cannot  
 165 be subtracted as a baseline because, it is carried on by the signal itself. Therefore, designing a CPL apparatus  
 166 requires that this first order term is negligible compared to  $S_3$ . As  $S_3 = g_{lum} S_0$ , we obtain the requirement on  
 167 the pump stability to get CPL free from first order artifacts:

$$\Delta\Phi \ll g_{lum} \quad (7)$$

168 Practically, it means that for standard  $g_{lum} \approx 10^{-3}$  the measurement of CPL by temporally inverting the  
 169 polarization requires stabilized excitation sources with relative variation  $\Delta\phi \leq 10^{-4}$ . As the weakness of the  
 170 CPL is the need for long integration time, it is practically impossible to avoid the excitation source variation  
 171 except for the fast modulation techniques. Indeed, PEM modulation allows to invert the polarization selection  
 172 at a very short time (a few  $\mu\text{s}$ ) preventing any pump drift. Moreover, averaging over a large number of cycles  
 173 increases the signal to noise ratio.

174 **Spatial separation of the polarization** is presented in the second row of Table 2. Simultaneous measurement  
 175 over two polarization encoded paths completely remove the pump variation effects but three other artifacts  
 176 appear. The first one ( $\sim S_0$ ) is related to the non equal transmission between the two optical paths that can  
 177 not be differentiate from a true non equal  $I_L$  over  $I_R$  luminescence. Less intuitive, the wavelength mismatch  
 178 of the two recorded spectra induces a CPL signal proportional to the luminescence derivative ( $\frac{\partial S_0}{\partial \lambda}$ ). QWP  
 179 imperfections mix the linear luminescence into the CPL. The non-perfect retardance ( $\psi$ ) couples with  $S_1$  (0–90°  
 180 polarization excess). This signal may appear in solid samples and also in solutions due to the polarization  
 181 selectivity of the excitation light, for a 90° excitation/detection scheme. The wrong alignment of the QWP  
 182 azimuth induces a coupling with  $S_2$  ( $\pm 45$  linear polarization excess) and should be only relevant for anisotropic  
 183 solid state samples. Because we work only with solution, we neglect these linear anisotropy terms in the following.

184 It is important to emphasize, that the artifacts described here are carried on by the fluorescence signal itself  
 185 ( $S_0$  and  $\frac{\partial S_0}{\partial \lambda}$ ). They are not unwanted background and can not be simply subtracted from a recorded spectra.  
 186 Therefore, they add ambiguous contributions to the CPL with spectral feature having similar shapes as the CPL  
 187 itself.

188 As  $\frac{\partial S_0}{\partial \lambda} \approx S_0/\delta\lambda$  where  $\delta\lambda$  is the luminescence spectral width, the requirements for the experimental parameters  
 189 in order to get signal artifacts lower than x% of  $g_{lum}$  are:

$$\Delta T \leq x\% \cdot g_{lum} \text{ and } \Delta\lambda \leq x\% \cdot g_{lum} \cdot \delta\lambda \quad (8)$$

190 For example, at 10%  $g_{lum}$  with small organic molecules ( $g_{lum} \sim 10^{-3}$  and  $\delta\lambda \sim 50\text{nm}$ ), we get  $\Delta T \leq 10^{-4}$  and  
 191  $\Delta\lambda \leq 0.005\text{nm}$ . For a higher  $g_{lum} \sim 10^{-2}$  obtained for rare earth complexes ( $\delta\lambda \sim 1\text{nm}$ ), the constraints are  
 192  $\Delta T \leq 10^{-3}$  and  $\Delta\lambda \leq 10^{-3}\text{nm}$ . Whatever the kind of sample to measure, the required properties of the optical  
 193 instruments (control of the transmissions better than 0.1%) and the need for monochromators with sampling  
 194 steps lower than 0.001nm are unrealistic and could explain the discrepancies observed between the measurements  
 195 made by camera and PEM-based set-ups for narrow lines rare-earth emitters[18]. A way to overcome this issue  
 196 consists in a calibration procedure with a non polarized light. But, as the artifacts are proportional to the shape  
 197 of the luminescence of the studied molecules, these calibration must be done with a racemic mixture of the  
 198 molecule at the very same concentration. Our own practice showed us that these calibrations are very difficult  
 199 to perform and must be re-evaluated almost before every measurement to be effective.

200 **Combination of time and spatial configurations** allows to reduce dramatically the artifacts related to the  
 201 CPL measurement by our setup. Inspection of the first two rows of the Table 2 reveals that the first order  
 202 terms do not depend on the  $p$ ,  $q$  parameters while CPL does. Consequently, artifacts add or subtract to the  
 203 CPL depending on the two measuring schemes. Therefore, by adding the two CPL signals measured for the  
 204 two configurations (last row of Table 2), the first order terms are canceled and the main remaining artifacts are  
 205 products of  $\Delta\phi$  with  $\Delta T$  or  $\Delta\lambda$  (see SI section 4.7 for the detailed calculation). The requirements to avoid false  
 206 CPL are then:

$$\Delta\Phi\Delta T \leq x\% \cdot g_{lum}, \quad \Delta\Phi\Delta\lambda \leq x\% \cdot g_{lum} \cdot \delta\lambda \quad (9)$$

207 If spatial or temporal constraints alone are very difficult to maintain at very low level, the combination of  
 208 spatial and temporal polarizations separations strongly reduces the effect of these constraints, rendering the  
 209 experimental setup much easier to perform and thus reliable CPL measurements much easier to obtain.

210 A final important point is that the arm inversion must be done without changing the polarization before the  
 211 monochromator. Indeed, gratings are very sensitive to the polarization with complicated wavelength depen-  
 212 dence. Therefore, inserting rotating polarizers in front of the monochromator will mix arm mis-balance with  
 213 monochromator polarization sensitivity. This is why we use the QWP before a PBS in such a way that the  
 214 polarizations are unchanged on each detecting arm.

## 215 3 Experimental validation

### 216 3.1 Molecules and their fluorescence and CPL spectra

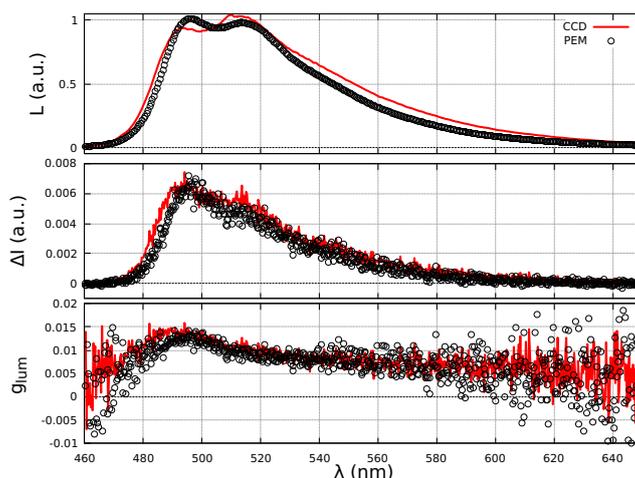
217 Two kind of molecular species have been used for this study : Camphorequinone **1**, a small organic molecule  
 218 with broad emission band (reference molecule used by the CPL spectrometer manufacturer: JASCO [24]) and  
 219 rare earth chiral complexes with narrow emission lines in order to put into evidence the CPL artifacts related  
 220 to fluorescence and fluorescence derivative.

221 **1** have been purchased from TCI company in enantiopure form, 1(R)-(-)-Camphorequinone and 1(S)-(+)-  
 222 Camphorequinone (SI, Figure 4). Fluorescence and CPL measurements have been carried out on solution  
 223 samples by dissolving these molecules in ethanol. Typical concentration of  $8 \cdot 10^{-3} \text{ mol.l}^{-1}$  were prepared.

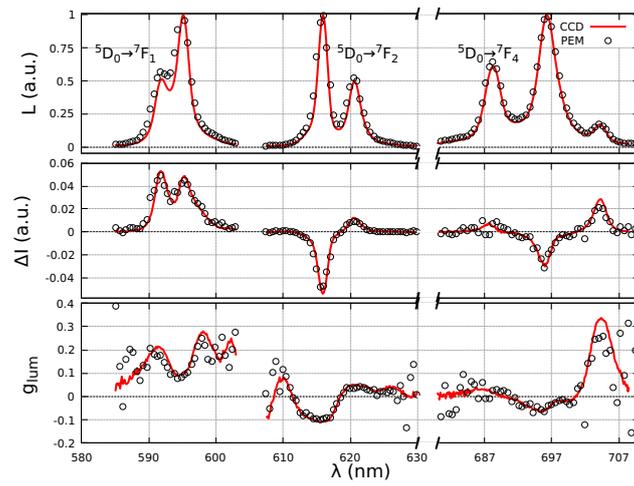
224 Complexes **2** and **3** are isostructural complexes of  $\text{Eu}^{3+}$  (**2**) and  $\text{Yb}^{3+}$  (**3**), respectively. They were prepared  
 225 in their enantiopure forms ( $[\text{Ln}(\mathbf{R},\mathbf{R})\text{-L}_3](\text{OTf})_3$  and  $[\text{Ln}(\mathbf{S},\mathbf{S})\text{-L}_3](\text{OTf})_3$ ), according to the procedure we  
 226 previously described[21]. Their structures (see SI, Figure 4) was shown identical in both solid state and solution  
 227 and their high brightness make them good candidates as a reference for the calibration of CPL set-ups.

228 For the measurements of **1**, with  $g_{\text{lum}} \sim 10^{-2}$ , we used a 830gr/mm grating leading to a dispersion over a 300  
 229 nm spectral range and a wavelength sampling interval of  $\sim 0.3\text{nm}$ , at the CCD side. The resolution was 1nm and  
 230 the integration time to get the presented spectra was 10minutes (5min for the acquisition of each configuration  
 231 plus 10s time to rotate the QWP). Camphorequinone solutions were excited with an Laser diode (10 mW,  
 232  $\lambda = 405\text{nm}$ ). The luminescence and CPL spectra displayed in Figure 3), present a broad single band extending  
 233 from 400 to 550nm. To compare the spectra features are also recorded with our step by step setup using a  
 234 standard PEM plus analyser system to differentiate via a lock-in amplifier the LHCP and RHCP. In this case,  
 235 the spectral resolution was 1nm, the step was 0.25nm and the integration time was 7s/step. Spectra recorded  
 236 with the two systems and corrected from the wavelength response, are very similar. The slight difference in the  
 237 relative intensities of the fluorescence and CPL bands show the residual errors effects which are more visible in  
 238 the case of wide emission bands, due to the different spectral response of the two setups. As for each of them,  
 239 the error on the fluorescence and CPL measurements is the same, it is cancelled in the  $g_{\text{lum}}$  spectra, as can be  
 240 seen in Figure 3. It took two hours for the PEM-based setup against 10 minutes for the CCD-based one, to get  
 241 spectra with equivalent signal to noise ratio.

242 For the measurements of complexes **2**, a (830gr/mm) grating was used with a twice optical magnification  
 243 providing a spectral range and a wavelength sampling of  $\sim 135\text{nm}$  and  $\sim 0.15\text{nm}$ , respectively. Spectra were  
 244 acquired under UV LED (4mW,  $\lambda = 365\text{nm}$ ), with an equivalent resolution of 0.5nm and acquisition time of  
 245 10s for each spectrum at the CCD side and 2.1s/step at the PEM one. The measured fluorescence and CPL  
 246 spectra display the usual features of  $\text{Eu}^{3+}$  with 3 main lines in the visible part of the spectrum (Figure 4).  
 247 Measurements using either the CCD (red line) or the step by step PEM (black circles) set-ups, give the same  
 248 results as expected. The acquisition time for the CCD setup is 30s (10s for the acquisition of each configuration  
 249 plus 10s time to rotate the QWP) against  $\sim 20\text{min}$  for the monochanel system, to measure equivalent spectra



**Figure 3:** Luminescence (top), CPL (middle) and  $g_{\text{lum}}$  (bottom) of **1** under 450 nm excitation. The scales luminescence and CPL are normalized to the maximum of emission. In red the spectra recorded with the CCD camera using the four signals procedure, in black circle the spectra recorded in a step by step set-up using a PEM+analyser system. Wavelength response correction has been applied for each system.



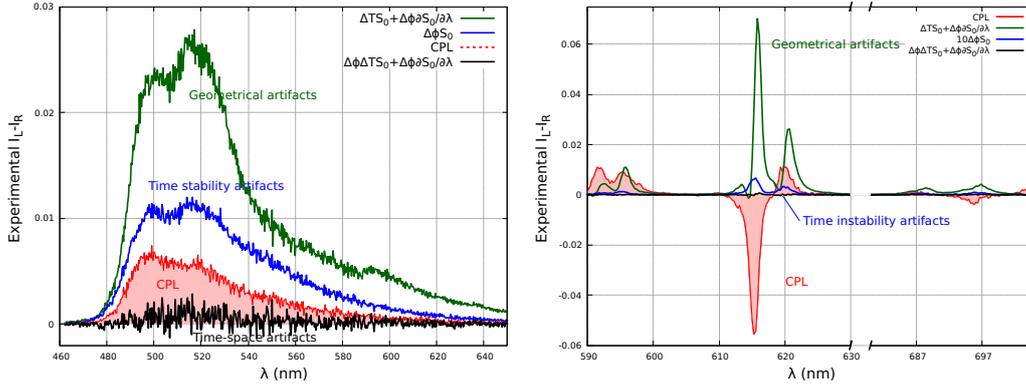
**Figure 4:** Three main visible transitions (luminescence (top), CPL (middle) and  $g_{lum}$  (bottom)) of  $\text{Eu}^{3+}$  complex **2** under 365nm excitation. In red the spectra recorded with the CCD camera using the time-spatial combination, in black circle the spectra are recorded in a step by step set-up using a PEM+Analyser system. The scales of each luminescence and CPL transitions are normalized to the maximum of emission.

250 with  $g_{lum}$  of few  $10^{-1}$ .

## 251 3.2 CPL artifacts

252 In order to characterize the measurements artifacts with the CCD-based set-up we proceed as follow. we first,  
 253 align our setup as well as possible. We take great care to make the two sensing paths as identical as can be  
 254 (same lenses, same relative position of the optical elements, top and bottom tracks centered on the camera). The  
 255 spectrometer is also precisely aligned: slit, grating and camera being parallel to each other, within one pixel.  
 256 After that, we measure the CPL spectra according to our spatial-time procedure (red line curves in Figure 5).  
 257 Then, we take off the quarter wave-plate and record exactly the same spectra. Assuming no linear fluorescence  
 258 effects, we must get a zero CPL signal. However, because of the geometrical and spatial imperfections described  
 259 in the theoretical part, non zero signal can arise. Mathematically speaking, taking off the QWP makes the  $\mathbf{Q}_q$   
 260 (SI Eq. 3) equal to identity matrice. Therefore,  $F_3=F_2 = 0$  and because  $S_1 = 0$  (no linear dicroism) only the  
 261  $S_0$  term remains and the recording signals are the artifacts listed in Table 2 (4<sup>th</sup> column).

262 Figure 5 displays the CPL artifacts measured on **1** and **2**. The difference between the two optical paths  
 263 without QWP, result in the green curve showing a false CPL similar to the luminescence spectra of **1** whose  
 264 intensity is of about 2.5% that of the luminescence spectra (Figure 3). This is the optical unbalance of our setup.  
 265 For the  $\text{Eu}^{3+}$  complex, the recorded signal present a false CPL opposite to the good one with an intensity of  
 266 7% of the luminescence (Figure 4). This 2.5-7% of intensity variation is the optical limitation for the spatial  
 267 separated polarizations-based set-ups and only CPL with  $g_{lum} \gg 0.05$  can be safely measured. The stability  
 268 of the excitation source is given by the differential measurement on the same channel but at 20s interval (blue  
 269 curve). Depending on the source, high power laser for **1** or low power light emitting diode for **2**, we observe  
 270 intensity variation of about 1% and 0.5% respectively. This is the limitation of time separated polarizations-  
 271 based set-ups. Finally, as described in the theoretical part, the combination of four measurements (black curve)  
 272 through the time-spatial-based configuration dramatically decreases the artifacts down to  $10^{-4}$  as the product  
 273 of 2.5-7% times 0.5-1%. It allows the safe study of samples with  $g_{lum} \geq 10^{-3}$ .



**Figure 5:** CPL artifacts of **1** (left) and **2** (right) recorded without the QWP (green -blue and black). To compare the CPL measured by the four signals combination is plotted in red. **1** is excited with a  $\lambda = 365$  nm emitting diode with 2 s integration time. **2** is excited with a  $\lambda = 450$  nm laser with 15 s integration time. Signals are normalized by the maximum of the luminescence. The green curves are the signal difference between the two polarization encoded arms: it images the unbalance of the optical paths. In blue, the signal difference for two measurements delayed by 20s obtained on the same arm, this is the  $\Delta\phi$  light source relative variation. 20s represents the required time to rotate the QWP by  $90^\circ$ . Black curves result from spatial-time combination.

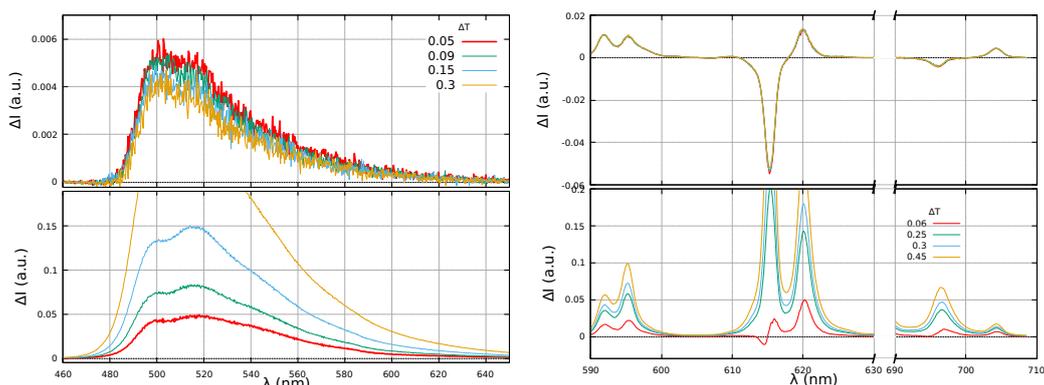
## 274 3.3 Robustness

275 Here, we experimentally investigate the robustness of the artifact-free CPL procedure. Starting from our best  
 276 aligned setup, we degrade on purpose some parts and record the corresponding spectra.

277 **Transmission misbalance** We place a variable neutral density filter in front of one of the two lenses  $L_2$  (see  
 278 optical setup figure 1), in order to lower the transmission along one polarization encoding path. The recorded  
 279 CPL spectra are displayed in Figure 6 for **1** (left) and **2** (right). The measured CPL spectra using the four signals  
 280 configuration are completely insensitive to the transmission difference between the two arms in the investigated  
 281 range up to  $\Delta T = 0.1$  for **1** and  $\Delta T = 0.4$  for **2**.

282 To compare, the CPL obtained for the spatial separated polarization are also displayed in the bottom panels.  
 283 By looking at the luminescence spectra (Figure. 3-4), the CPL artifacts clearly depends on the luminescence  
 284 as described by the term  $\Delta T \times S_0$  (Table 2). On the bottom left of the figure, the false measurement are

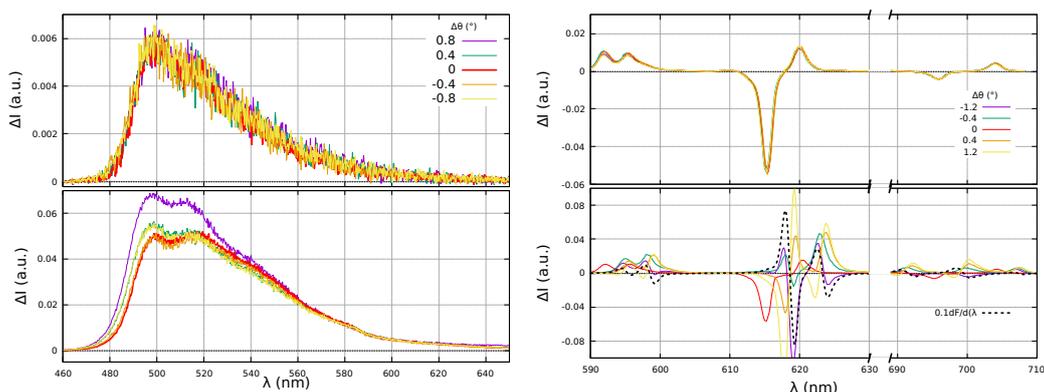
285 nearly proportional to  $\Delta T$  with a shape very similar to the luminescence spectra of the corresponding molecule  
 286 (Figure. 3) with some ripple effects attributed to the slow wavelength dependence of  $\Delta T$ . On the bottom right,  
 the narrow emission lines only let appear the shape of the luminescence in the false CPL.



**Figure 6:** CPL of **1** (left) and **2** (right) by using either the artifacts-free procedure (top) or taking one single  $I_L-I_R$  difference (bottom). Note the y-axis different range between the top and bottom. The different curves are recorded for different neutral density attenuation on one arm. The average transmission difference factor  $\langle \Delta T \rangle$  is calculated from the luminescence integral measured at each arm.

287

288 **Wavelength mismatch** To add a small wavelength mismatch between the two polarization encoded spectra,  
 289 we tilt a bit the camera away from the well aligned position. Again, we show that there is not any difference  
 290 in the recorded CPL spectra using the time-spatial procedure (top spectra of Figure. 7). For this experimental  
 291 mis-alignment, the CPL artifacts present in the spatial separated polarization procedure are related to the  
 292 fluorescence derivative as illustrated in the bottom part of Figure. 7. The order of magnitude of the measured  
 293 CPL compare to the luminescence derivative for the Eu complex (right bottom), shows that for our best set-up  
 294 (red curve) the wavelength mismatch is negligible around 595nm and is about 0.02 nm and 0.1 nm at 620 and  
 295 700 nm respectively. It indicates that, under best alignment, we can not guarantee a perfect spectral matching  
 296 on the whole spectral range (here 120 nm). Moreover, as the spectral sampling interval is 0.13 nm, it is not  
 297 possible to get a better accuracy and the CPL artefact related to this wavelength mismatch between the two  
 298 recorded spectra on two different part of the CCD is inevitable. However, our procedure is very stable in term  
 299 of CPL results with reproducible spectra even with strong misalignment (see the image distortion obtained for  
 300 a  $1.6^\circ$  tilt in the SI section 6).

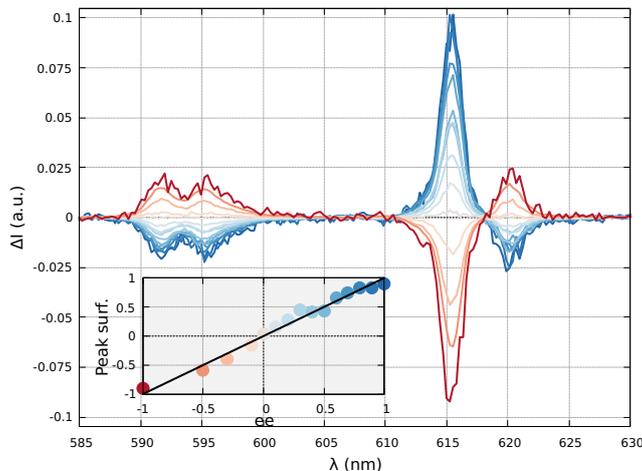


**Figure 7:** CPL of **1** (left) and **2** (right) by using either the artifacts-free procedure (top) or taking one single  $I_L-I_R$  difference (bottom). The different curves are recorded for different tilt mis-alignment of the CCD camera. Note the y-axis scale difference between the top and bottom panels. The dashed black curves are the derivative of the luminescence signal.

### 301 3.4 Linearity

302 We prepare two enantiopure solutions of  $\text{Eu}^{3+}$  complex **2** with the same concentration of ( $10^{-5}$  mol/l) but with  
 303 opposite handedness. We record a series of CPL spectra starting from one enantiopure solution and by adding,

304 gradually, the opposite enantiopure solution to the first one, we decrease the CPL of the mixture gradually until  
 305 it is cancelled for the racemic solution. The idea is to decrease the CPL signal gradually at a constant emission in  
 306 order to characterize the linearity of the measurements. Figure. 8 clearly shows this gradual racemization effect  
 and the decrease in intensity of the CPL, the inset in the figure attests for the linearity of our measurements.

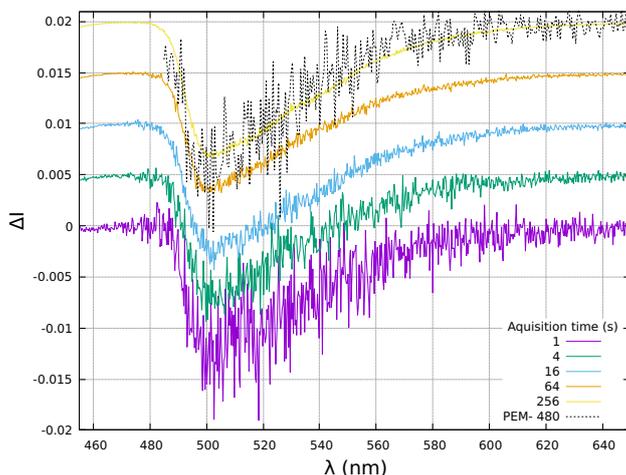


**Figure 8:** CPL spectra measured for different enantiomeric excess ee. The inset displays the surface of the 615nm peak versus ee.

307

### 308 3.5 Signal to noise ratio

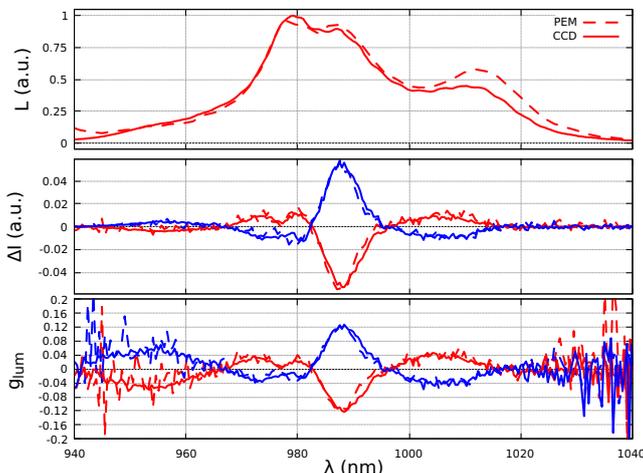
309 The CPL spectra of **1** is displayed in Figure 9 for different integration times from 1s to 256s using the CCD  
 310 apparatus (this is the integration time for one QWP configuration). As expected, the signal to noise ratio (SNR)  
 311 increases with the integration time, and this without any artifacts coming from the lamp stability. To compare,  
 312 we have used the same source and the same solution to record the CPL spectra with a step by step mono-channel  
 313 setup (PEM plus analyser). The overall experimental time of 480s (i.e. a step=2.5nm and an integration time=  
 314 2.1s/step ) results in the black dotted spectra plotted in the top of the figure. The SNR is similar to what  
 315 is obtained with the camera for 1s integration time spectra. For similar wavelength sampling interval, same  
 316 spectral range and a comparable SNR, the recording with the camera is ~500 times faster.



**Figure 9:** CPL spectra of camphorequinone excited with a  $\lambda = 450\text{nm}$  laser, recorded with the CCD spatial-time configuration at different integration times from 1-256s and the counterpart, in black-line, recorded with the step by step (PEM + analyser) based setup, with 1s/step integration time.

## 317 4 CPL in the near IR

318 Our setup can operate in a wide wavelength range from the UV to the near IR. Indeed, by changing only the  
 319 PBS cube and the transmission grating, we can easily switch from the UV-visible to the near infrared spectral  
 320 range. All other optical and detection elements (QWP, lenses, fibers and CCD) are chosen to cover the entire  
 321 spectral range from 300 nm to 1.1  $\mu\text{m}$ . Thereby, CPL of enantiopure complexes **3** ( $[\text{Yb}^{3+}(\mathbf{R},\mathbf{R})\text{-L}_3](\text{OTf})_3$  and  
 322  $[\text{Yb}^{3+}(\mathbf{S},\mathbf{S})\text{L}_3](\text{OTf})_3$ ) are recorded using the spatial-time procedure, for the two enantiomers and compared  
 323 with their counterparts obtained by our single channel PEM based setup with appropriate IR PM detector  
 324 (Hamamatsu H10330B-75).



**Figure 10:** Luminescence (top), CPL (middle) and  $g_{lum}$  (bottom) for the two enantiomers of **3** under 450nm excitation. The scales luminescence and CPL are normalized to the maximum of emission, so that the  $g_{lum}$  value at maximum luminescence, can be directly read on the y-axis. In continuous line spectra recorded with the CCD camera using the spatial-time procedure, in dashed line the spectra recorded in a step by step mode using a PEM+analyser system. Wavelength response correction has been applied for each system.

325 The spectra are recorded over 940-1040nm spectral range, under 450nm laser excitation (power 10mW). The  
 326 four obtained luminescence spectra are similar but to avoid crowding the graph, we only show on Figure 10,  
 327 one emission spectra for one enantiomere, measured by each setup. The normalised emission spectra are almost  
 328 identical and are typical of  $\text{Yb}^{3+}$ . The difference in intensity at long wavelengths, despite their correction by  
 329 the spectral response of the corresponding setup, is due to the detection limit of the CCD in this spectral range,  
 330 inducing a high error scale. The corresponding CPL bands are well mirror images for the two enantiomers and  
 331 for each of them, the spectra obtained by both setups are identical in terms of number of bands, their shapes  
 332 and their positions in wavelength as well as their relative intensities, as attested by Figure 10. Even though the  
 333  $g_{lum}$  value of the measured  $\text{Yb}^{3+}$ -complexes is rather high ( $10^{-1}$ ), the measurement of the CPL on the PEM  
 334 side is more time demanding because of the low signal to noise ratio of the detector in this wavelength range.

335 It took thus, 420s (step = 0.5nm/integration time = 2.1s/step) to record a CPL spectrum on the PEM side  
 336 while only 0.1s per spectrum (i.e. a global time of 10.2s: 2 spectra + QWP rotation) was needed to obtain the  
 337 same spectrum with, moreover, a better SNR, at the CCD side. Here again, our CCD-based setup with the  
 338 spatial-time method shows its strength for the measurement of fast and reliable CPL in the near IR range. This  
 339 range can be extended by the use of adapted CCD.

340 It is worth noting that to date, only a few pioneering works concerning the measurement of CPL in the near  
 341 IR region, using the conventional CPL spectrometer, have been reported [21, 23, 25, 26]. IR region remains  
 342 thus, a relatively new area to explore for CPL emitting materials and relative potential applications.

## 343 5 Discussion

344 The CPL measurement by combining four measures corresponding to spatial and temporal separation of the  
 345 circular polarizations allows a self-compensation of the artifacts and the direct measurement of  $S_3$  to the nearest  
 346  $\eta_{mix}F_3$ -factor (Table 2, third columns) with  $F_3$  and  $\eta_{mix}$  defined in Eq. 5-6.

347  $\eta_{mix}F_3$ -factor : This corrective-factor reflects the deviation from the linearity of the system due to the QWP  
 348 related errors and the polarization mixing in the two encoded optical paths (stray-light + PBS imperfection).  
 349 These errors cannot be suppressed for any of the three CPL measurements methods based on CCD . All we can  
 350 do , is to minimize them as much as possible.

- 351 • QWP related errors: phase retardation  $\Psi \neq \pi/2$  and azimuth  $\theta \neq \pm\pi/4$  are taken into account by  $\cos\psi$   
 352 and  $\cos(2\alpha)$  (with  $\psi = \Psi - \frac{\pi}{2}$  and  $\alpha = \theta - \frac{\pi}{4}$ ), terms respectively. As these two corrective factors are  
 353 *cosinus* dependent, their effects are minimal. Indeed, a phase delay  $\Psi = \frac{\pi}{2} \pm \frac{\pi}{100}$  (standard specification  
 354 of achromatic QWP over a few hundreds of nm) induces an error of 0.2% on the measured CPL. QWP  
 355 azimuth accuracy around 0.1deg can be easily achieved using standard alignment procedure by placing  
 356 the QWP between two cross polarizers. Even if this orientation is not accurate, a misalignment of  $\pm 2^\circ$   
 357 results in an error of 0.06%. Therefore, a standard quality of the QWP and its rough orientation does not  
 358 induce significant errors on the CPL magnitude.
- 359 • Mixing polarization related errors,  $\eta_{mix}$ : contrary to the QWP, the quality of the PBS or of any polarizer  
 360 used in a CPL measurement setup is very important. If the polarizer rejection ratio is 90%, for example,  
 361 then 10% of one polarization is injected in the arm transporting the other one. This results in an un-  
 362 derestimation of the CPL of 20%. Thus, the better the polarizer, the lower the error on CPL. Besides,  
 363 a small relative part of one polarization falls into the detection area of the other one's depending on the  
 364 straylight of the spectrometer. This leads again, to an underestimation of the CPL of 2 to 4% in our setup.  
 365 The straylight is the limit of CPL measurements setups based on a single spectrometer. The use of two  
 366 spectrometers improves this aspect but will introduce a new source of errors due to the responsivity drift  
 367 between the two detectors.

368 **Artifacts sources** : the artifacts sources for the three configurations of CPL measurements, have been identified.  
 369 The theoretical analysis of these artefacts show that their spectral shape is proportional to the fluorescence  
 370 signal and to its derivatives and therefore, cannot be subtracted as a simple baseline. Besides, their experimental  
 371 quantification, by taking off the QWP and measuring each of them, allow to asses the limits of each configuration.

- 372 • Artefacts related to the spatial separations of the polarization i.e. non equal transmission ( $\Delta TS_0$ ) and  
 373 wavelength mismatch ( $\Delta\lambda \frac{\partial S_0}{\partial \lambda}$ ) between the two optical paths (Table 2, second column) can not be sup-  
 374 pressed in the case of CPL measurements using only spatial separation. The constraints on  $\Delta T$  and  $\Delta\lambda$   
 375 are CPL dependent ( $\Delta T \ll g_{lum}$  ;  $\Delta\lambda \ll g_{lum} \cdot \delta\lambda$ ) and require unrealistic control of the transmissions,  
 376 better than 0.1% and spectrometer sampling step, lower than 0.001nm. Experimentally, with the best  
 377 alignment of our setup, the unbalance between the two arms induced not only false CPL whose intensity  
 378 is about 2.5 – 7% that of the luminescence, but also CPL band opposite to the good one, in the case of  
 379 narrow emission bands (Figure 5). Indeed, the artifacts related to the wavelength detuning are propor-  
 380 tional to the derivative of the luminescence signal and therefore more important for the narrow emission  
 381 lines. They introduced false CPL signal with intensity of  $\sim 8\%$  that of the luminescence.

382 Hence, in order to avoid as much as possible these false CPL contributions in the measured signal by  
 383 only the spatial separation configuration, adapted and rigorous calibration, using racemic sample and non  
 384 polarized light, is needed and have to be checked before each set of measurements.

- 385 • These constraints are no longer relevant in the case of CPL measurements by temporally inverting the  
 386 polarization, as only one optical path is involved. However, to get non-erroneous CPL spectra, the excita-  
 387 tion source stability is crucial with its variation between the two  $I_L$  and  $I_R$  measurements,  $\Delta\Phi \ll g_{lum}$ .  
 388 This stability is difficult to ensure during the two successive measurements, especially since long integra-  
 389 tion times are necessary for accurate CPL measurements. Using a LED excitation source with long term  
 390 stability of a few percents and standard optical elements, false CPL signal whose intensity is equal to 1%  
 391 of the luminescence, was measured.
- 392 • By combining spatial and temporal-separation of the circular polarization and performing two sets of  
 393 measurements where the role of the polarization encoded arms are inverted, then combining the four  
 394 obtained measures, the first order artefacts relative to the two precedent configurations vanish and only  
 395 second order terms remain, (see Table 2, third line). Apart from the CCD calibration for wavelength  
 396 accuracy, fast CPL spectra can then be safely recorded by the CCD-based spectrophotometer without the  
 397 need for tedious alignment and heavy calibration procedure each time a sample is changed or an optical  
 398 element is moved.

399 Compared to the standard modulation technique with lock-in amplifier, CCD-based devices do not have AC  
400 filter. Therefore, the intensity resolution is the same over the whole range of the measurement and small signals  
401 are difficult to extract. In our case, with a 16 bits camera, the analog to digital conversion leads to an accuracy  
402 of  $\frac{1}{2^{16}} = 1.5 \cdot 10^{-5}$  for a full signal range. This is the fundamental lower limit for  $g_{lum}$ .

403 Finally, it is worth to remind here, that whatever the spectroscopic instrument used, it is necessary to take  
404 into account its wavelength response and to correct the measured spectra by the corresponding device function.  
405 Otherwise, the relative intensity of the spectral bands can be wrong. This remains a delicate operation if we  
406 want to compare measurements obtained by two different setups. Indeed, each spectrum must be corrected by  
407 the instrument function with which it was obtained and the instrument functions themselves, must be recorded  
408 properly. If we consider the spectra of Camphorequinone and Yb<sup>3+</sup>-complexes, with large emission bands,  
409 (Figure 3 and Figure 10, respectively) we can see that, despite the wavelength response correction, there is still  
410 difference mainly in the relative intensities of the luminescence and CPL bands. However, as this error is the  
411 same on both measurements, it disappears in the  $g_{lum}$  and there, we can really compare the two methods more  
412 safely.

## 413 Conclusion

414 Measuring CPL is a delicate matter because it requires making a differential measurement with a relative  
415 precision of  $10^{-5}$  on the emission signal. Therefore, the smallest artifacts can have a value of the same order of  
416 magnitude as the CPL itself or even larger leading to false CPL spectra. For this reason, modulation systems  
417 (PEM + lock-in amplifier) and a step by step measurements, with a very long integration time are generally  
418 used to extract the good signal.

419 In order to reduce the acquisition time and still obtain reliable spectra, a CPL spectrophotometer based on  
420 only one CCD camera and the combination of two inverted polarization encoded channels has been theoretically  
421 analyzed, set up and tested on two different classes of molecules giving narrow or broad band spectra respec-  
422 tively. We studied the limitations of three different configurations of this set up: time, spatial and spatial-time  
423 separations of the polarization.

424 We demonstrated that unless using a rigorous and heavy calibration procedure, the artifacts related to the two  
425 first configurations cannot be completely suppressed and thus, inevitably lead to more or less false CPL spectra.  
426 This can range from a wavelength shift, deformation, false relative intensities to a sign reversal of the CPL  
427 bands. Consequently, a misinterpretation of stereochemical structures and associated transitions of the studied  
428 chiral system or a wrong evaluation of the  $g_{lum}$  is very likely. However, we showed that the third configuration,  
429 i.e. the spatial-time combination is the most efficient for obtaining reliable CPL spectra capable of measuring  
430  $g_{lum} > 10^{-4}$  in a robust, reproducible and fast way. It offers many advantages over CPL measurement by only  
431 time or only spatial separation of the polarization :

- 432 1. Accurate and fast measurement with standard optical elements for  $g_{lum} \geq 10^{-3}$
- 433 2. Nearly no need for calibration, the measured signal is directly the CPL times a correcting factor  $\in [0.92 -$   
434  $1.00]$  coming mostly from the residual straylight and the PBS imperfections;
- 435 3. Auto-compensation of the artifacts at first order. The remaining artifacts are set at a very low value  
436 around  $10^{-5}$  times the luminescence; No need of high quality QWP
- 437 4. Possibility to simply measure the artifacts on the *same sample* by just taking off the QWP.

438 Compared to a more standard mono-channel setup with polarization modulation, the recorded time is reduced  
439 by three order of magnitude, for the same signal to noise ratio, and a fast measurement on a whole spectral  
440 range is obtained in one shot.

441 The implementation of such a fast and robust CPL spectrophotometer opens interesting perspectives for the  
442 monitoring of dynamic processes such as chemical reactions that vary over time, depending on a particular  
443 parameter. Moreover, thanks to its rapid measurements, the stabilization of the external parameter, such as  
444 temperature or magnetic field, is not any more as critical as in the case of much longer measurements with  
445 PEM-based setups.

446 Finally, this method can also be valuable for an automatic screening of enantioselective syntheses.

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