

Chromitite layers require the existence of large, long-lived, and entirely molten magma chambers

Rais Latypov (✉ rais.latypov@wits.ac.za)

University of the Witwatersrand <https://orcid.org/0000-0002-8158-5949>

Sofya Chistyakova

University of the Witwatersrand

Stephen Barnes

CSIRO

Belinda Godel

CSIRO

Gary Delaney

CSIRO Data61, Clayton South

Paul Cleary

CSIRO Data61, Clayton South

Viktor Radermacher

Evolutionary Studies Institute <https://orcid.org/0000-0001-6524-7811>

Ian Campbell

Australian National University <https://orcid.org/0000-0001-8578-0154>

Kudakwashe Jakata

European Synchrotron Radiation Facility <https://orcid.org/0000-0001-7217-2774>

Article

Keywords: molten magma chambers, Bushveld Complex, stratiform chromitites

Posted Date: May 26th, 2021

DOI: <https://doi.org/10.21203/rs.3.rs-443338/v1>

License:  This work is licensed under a Creative Commons Attribution 4.0 International License.

[Read Full License](#)

1 **Chromitite layers require the existence of large, long-lived, and**
2 **entirely molten magma chambers**

3

4 Rais Latypov^{1*}, Sofya Chistyakova¹, Stephen J. Barnes², Belinda Godel², Gary W. Delaney³,
5 Paul W. Cleary³, Viktor J. Radermacher⁴, Ian Campbell⁵, and Kudakwashe Jakata^{2,6}

6

7 **An emerging and increasingly pervasive school of thought is that large, long-lived and**
8 **largely molten magma chambers are transient to non-existent in Earth's history¹⁻¹³.**
9 **These ideas attempt to supplant the classical paradigm of the 'big magma tank' chambers**
10 **in which the melt differentiates, is replenished, and occasionally feeds the overlying**
11 **volcanoes¹⁴⁻²³. The stratiform chromitites in the Bushveld Complex – the largest**
12 **magmatic body in the Earth's crust²⁴ – however, offers strong contest to this shifting**
13 **concept. Several chromitites in this complex occur as layers up to 2 metres in thickness**
14 **and more than 400 kilometres in lateral extent, implying that chromitite-forming events**
15 **were chamber-wide phenomena²⁴⁻²⁷. Field relations and microtextural data, specifically**
16 **the relationship of 3D coordination number and grain size, indicate that the chromitites**
17 **grew as a 3D framework of touching chromite grains directly at the chamber floor from**
18 **a melt saturated in chromite only²⁸⁻³⁰. Mass-balance estimates dictate that a 1 to 4 km**
19 **thick column of this melt^{26,31,32} is required to form each of these chromitite layers.**
20 **Therefore, an enormous volume of melt (>1,00,000 km³)^{24,25} must have been involved in**

¹School of Geosciences, University of the Witwatersrand, Johannesburg, South Africa; ²CSIRO Mineral Resources, Kensington, Perth, WA 6151, Australia; ³CSIRO Data61, Clayton South, VIC 3169, Australia; ⁴Evolutionary Studies Institute, University of the Witwatersrand, Private Bag 3, 2050, South Africa; ⁵Australian National University, Canberra ACT 0200, Australia; ⁶European Synchrotron Radiation Facility (ESRF), France.
*Corresponding Author, email: Rais.Latypov@wits.ac.za

21 **the generation of all the Bushveld chromitite layers, with half of this melt being expelled**
22 **from the magma chamber^{24,26}. We therefore argue that the very existence of thick and**
23 **laterally extensive chromitite layers in the Bushveld and other layered intrusions strongly**
24 **buttress the classical paradigm of ‘big magma tank’ chambers.**

25

26 For over a century, the classical paradigm of magma chambers has underpinned all models of
27 the Earth’s magmatism. This paradigm envisages a magma chamber as a large body of the
28 molten, long-lived, and slowly fractionating magma (‘a big magma tank’) enclosed in crustal
29 rocks^{14–23}. In recent years, this classic view of a magma chamber garnered increased scrutiny
30 from many geoscientists who contend that such long-lived and largely molten ‘big tank’
31 magma chambers are either very short-lived or never existed in Earth’s history^{1–13}. For
32 instance, some igneous petrologists maintain that tectonic processes that create open spaces in
33 the crust at any tectonic setting are so slow that the liquid-dominated magma bodies of any
34 composition can hardly form^{1,2}. Instead, they have suggested that plutons grow incrementally
35 from numerous sills/dykes a few meters thick that cool and may even become totally solidified
36 between injections^{1,2}. Similarly, most volcanologists have abandoned the classic paradigm
37 because the geophysical surveys have failed to detect any present-day eruptible magma bodies
38 in the Earth’s crust^{4,8}. As an alternative, they proposed the existence of transcrustal mushy
39 systems (including mushy reservoirs for mafic layered intrusions⁴) that are formed in the crust
40 from numerous collating intrusions. The transcrustal systems contain only small melt lenses
41 that are produced by compaction^{3,7} or tectonic destabilization^{4,8} of the crystal mush and exist
42 for only a very short period of time before accumulating and erupting as lavas on the Earth’s
43 surface^{1–8}. Yet another group of mafic plutonists, influenced by out-of-sequence zircon
44 geochronological data^{9,10,13}, have proposed that mafic plutons do not require the existence of
45 large magma chambers¹². These are rather produced as a stack of randomly-emplaced sills,

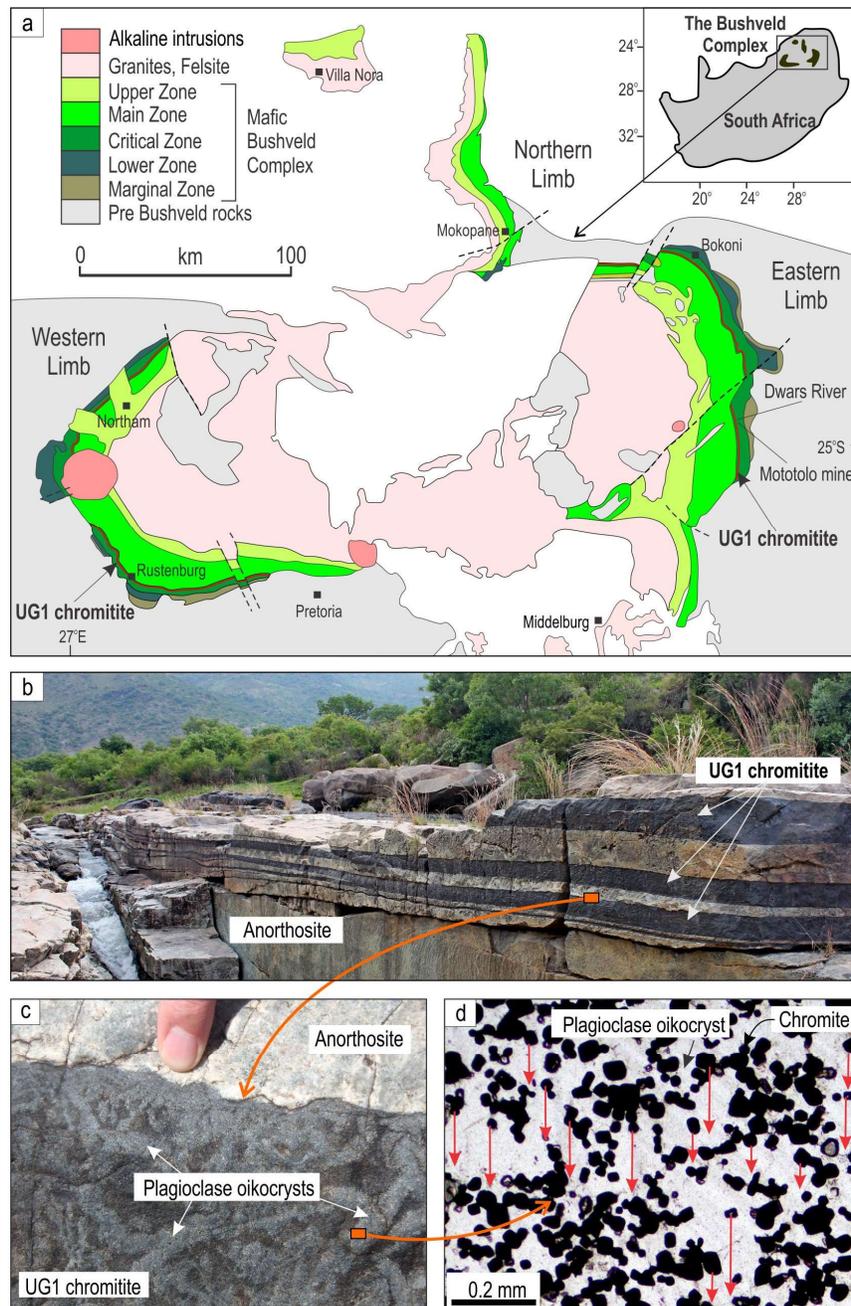
46 with successive crystal-rich pulses often invading pre-existing cumulates⁹⁻¹³. In stark contrast
47 to these ‘anti-magma-chamber’ approaches, we argue here that the existence of large magma
48 chambers is indicated by laterally extensive layers of chromite-rich cumulates, which require
49 many times their own volume of magma to supply the key component, chromium (Cr). We
50 present here field and microtextural data from massive chromitites of the Bushveld Complex
51 that indicate that ‘big magma tank’ chambers in the Earth’s crust are a reality that cannot and
52 should not be dismissed.

53

54 **The enormous extent of chromitite layers**

55 The 2.05 Ga Bushveld Complex in South Africa (Fig. 1a) is the largest mafic-ultramafic
56 layered intrusion in Earth’s crust; totalling to about 600,000 km³ of igneous rocks^{24,25}. The
57 complex consists of several parts, the western, eastern and northern limbs being the largest,
58 and is subdivided stratigraphically into five major units - the Marginal, Lower, Critical, Main,
59 and Upper Zones, comprising a total thickness of 7 to 9 km^{24,25}. The Bushveld Complex
60 contains >80% of the Earth’s known chromium resources³³, an element critical to improving
61 the material properties of steel, making this magmatic body an object of perennial study. The
62 chromium is hosted within 14 principal layers of massive chromitites, mostly confined to the
63 Critical Zone^{26,27}. Three major groups of chromitites are recognised: the Lower (LG1 to LG7);
64 Middle (MG0 to MG4); and Upper Groups (UG1 to UG3)²⁷. The thickness of individual
65 chromitite layers ranges from a few decimeters to 2 meters. Mining activities have allowed
66 most of these layers to be traced across the entire Bushveld Complex²⁷. Remarkably, the
67 vertical distribution of platinum-group elements across some of these chromitites are nearly
68 identical in places that are separated laterally by over 300 km²⁶. The vast lateral extent and
69 mineralogical uniformity of chromitite layers indicate that the process responsible for their

70



71

72

73 **Fig. 1. A notable lack of textural evidence for gravity settling of chromite crystals within a UG1 chromitite**
 74 **of the Bushveld Igneous Complex, South Africa.** **a**, Location and generalized geological map of the Bushveld
 75 Complex emphasizing its enormous size. Note that this is just an erosional remnant of the complex so that its
 76 original size was even larger. The immense lateral extent of the chromitite layers (>350-400 km) in this complex
 77 is illustrated here by the UG1 chromitite that occurs at the top of the Critical Zone. Also indicated are places from
 78 which the studied samples were obtained. Modified from reference²⁸. **b**, Panoramic view of a few sublayers of
 79 the UG1 chromitite in the anorthosite footwall at the Dwars River, Eastern Limb. **c**, A close-up photograph of the
 80 top part of the UG1 chromite sublayer. Note large plagioclase oikocrysts enclosing numerous small crystals of
 81 cumulus chromite (chadacrysts). **d**, Photograph of a thin-section (under plane polarized light) of the UG1
 82 chromitite showing isolated chromite grains and their loose clusters enclosed by a single large oikocryst of
 83 plagioclase. Red arrows emphasize that chromite grains show no tendency to gravitate downwards despite a high
 84 porosity of the framework (~65 vol.%). Sample HX-07-153.33, Mototolo mine, Eastern Limb.

85 formation has been working synchronously in all parts of the superlarge chamber to produce
86 the same chromitite layer over lateral distances of up to 400 km (e.g., UG1 in Fig. 1a).

87

88 **Field and textural evidence for *in situ* growth of chromite**

89 The nature of this chamber-wide process can be constrained from field and textural features of
90 massive chromitite layers such as the 2 m-thick UG1 chromitite – the thickest and the best
91 exposed layer in the entire complex (Fig. 1b). This chromitite shows remarkable field
92 relationships with its respective footwall rocks. In addition to its occurrence on the planar
93 portions of the chamber floor, this chromitite develops within potholes, roughly circular
94 depressions in which footwall rocks are missing due to magmatic erosion^{28,34}. In these areas,
95 the planar UG1 chromitite that occurs along the periphery of the potholes commonly passes,
96 without any apparent changes in thickness and texture, into the steeply dipping, subvertical
97 and even overhanging UG1 chromitite in the interior of potholes^{28,34} (Extended Data Fig. 1).
98 This field observation strongly argues against the formation of the UG1 chromitite, both on the
99 planar and overhanging portions of the chamber floor, by processes involving gravity-induced
100 settling of chromite through either the resident melt^{35–38} or a crystal-rich mush^{39,40}. The
101 simplest alternative mechanism is *in situ* growth of chromite directly at the chamber floor from
102 a chromite-only-saturated melt^{28,30}. This is the only process that allows the chromitite layer to
103 cover all the planar and irregular margins, even the places where gravity-settling of chromite
104 grains is physically impossible (i.e., “gravity-settling shadows” in which dips are
105 overturned^{28,34}) (Extended Data Fig. 1).

106

107 An intriguing challenge here is to decipher how *in situ* growth of chromite is recorded in the
108 texture of massive chromitites themselves. We have re-visited the UG1 chromitite from the
109 classical Dwars River locality²⁹ (Fig. 1b) where it is composed of 25-50 vol% of cumulus

110 chromite that occurs as separate idiomorphic grains or clumps of grains that are smaller than
111 0.1 mm in size (Fig. 1c, d). The chromite grains are enclosed within much larger oikocrysts of
112 plagioclase (up to 5-10 cm in size) that are clearly visible in outcrops (Fig. 1c). The traditional
113 interpretation of such layers in the frame of gravity settling models is that chromite was the
114 first to settle on the chamber floor³⁵⁻³⁸ followed, after some period of post-depositional cooling,
115 by *in situ* growth of plagioclase oikocrysts from the interstitial melt in a mushy chromitite. An
116 important point is that settling chromite grains have enough time to reach the chamber floor
117 and start growing there. The subsequently forming oikocrysts may capture and arrest chromite
118 from experiencing further growth, producing snapshots of an immature solidification front.

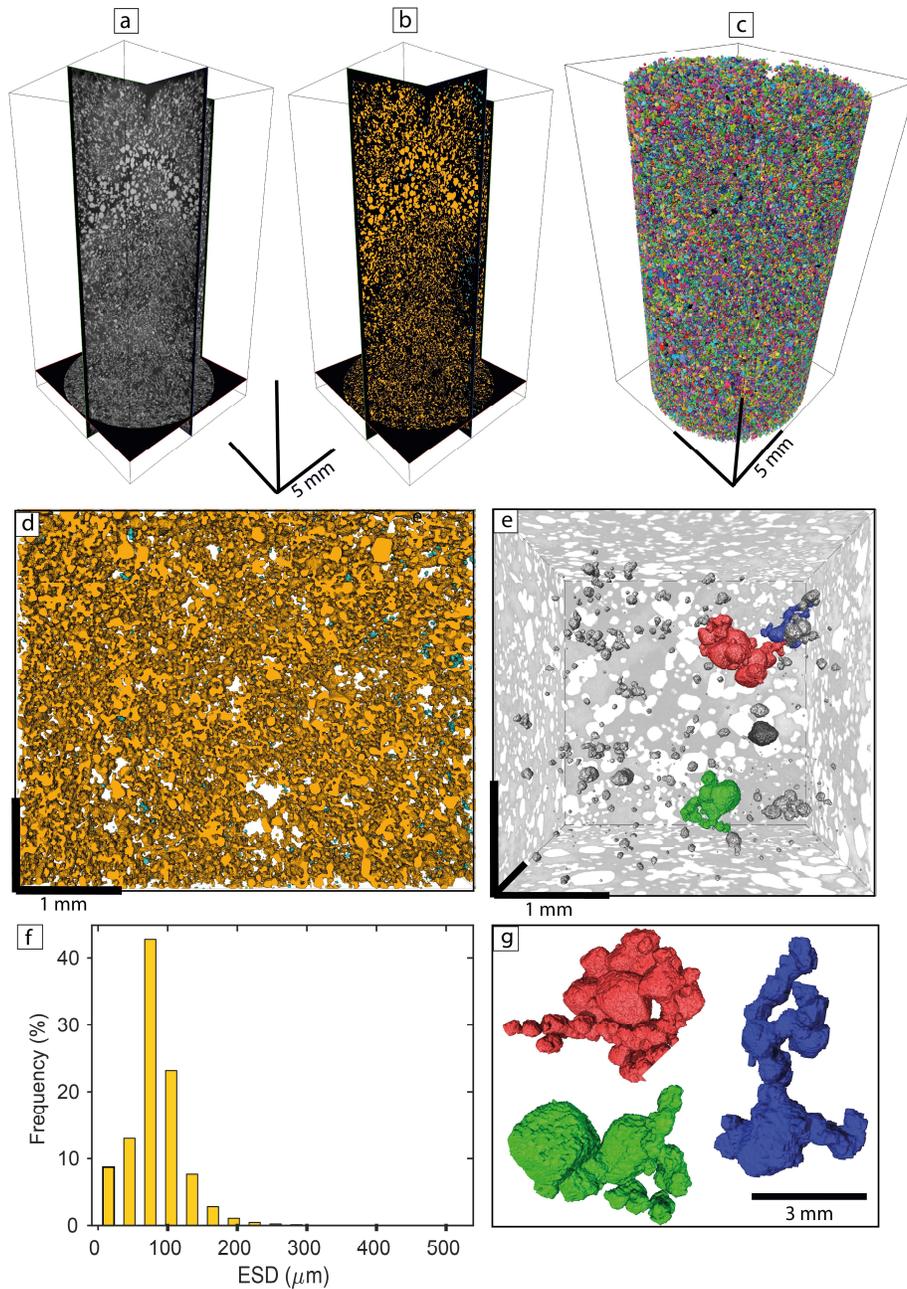
119

120 A close look at the UG1 texture (Fig. 1d) raises, however, a simple but fundamental quandary.
121 Chromite is almost twice as dense as a basaltic melt (4,800 kg/m³ and 2,600 kg/m³,
122 respectively) and is expected to settle to the chamber floor in a random closely-packed lattice
123 in which all adjacent chromite grains are touching each other. However, this is not the case as
124 chromite occurs as individual grains and clumps of grains that are ‘suspended’ within
125 plagioclase oikocrysts (Fig. 1d). This observation leads to a critical question: why have the
126 chromite grains/clumps failed to sink towards the chamber floor despite being much denser
127 than the host melt? A potential clue to this puzzle is that the chromite grains in the UG1 layer
128 appear to be arranged in chain-like aggregates^{41,42}.

129

130 **Three-dimensional framework of chromite crystals**

131 The analysis and quantification of chromitite in three-dimensions (3D) using high-resolution
132 X-ray computed tomography (HRXCT) revealed that nearly all chromite grains (97 vol%) from
133 the UG1 chromitite are interconnected to form a single continuous 3D framework composed
134 of many thousands grains that extend across multiple plagioclase and pyroxene oikocrysts (Fig.



135
 136 **Fig. 2. Results of high-resolution X-ray computed tomography revealing that nearly all chromite grains in**
 137 **UG1 chromitite are interconnected within a continuous 3D framework.** (a) Three orthogonal slices virtually
 138 cut through the UG1 sample (HX-07-153.33, Mototolo mine, Eastern Limb) showing chromite in light grey;(b)
 139 Segmented chromitite showing how a single interconnected network (coloured in yellow) covers the entire sample
 140 volume; (c) Volume rendering of chromite grains displayed using a false 256 colour scale; (d) Expanded view of
 141 a volume of interest showing details of the interconnected chromite network (orange) and the isolated chromite
 142 grains (in cyan); (e) and (g) Details of selected chromite cluster morphologies within the large interconnected
 143 chromite network where only a small number of grains are coloured to improve visibility; (f) Histogram showing
 144 the size distribution of chromite grains in the sample (ESD: equivalent sphere diameter).

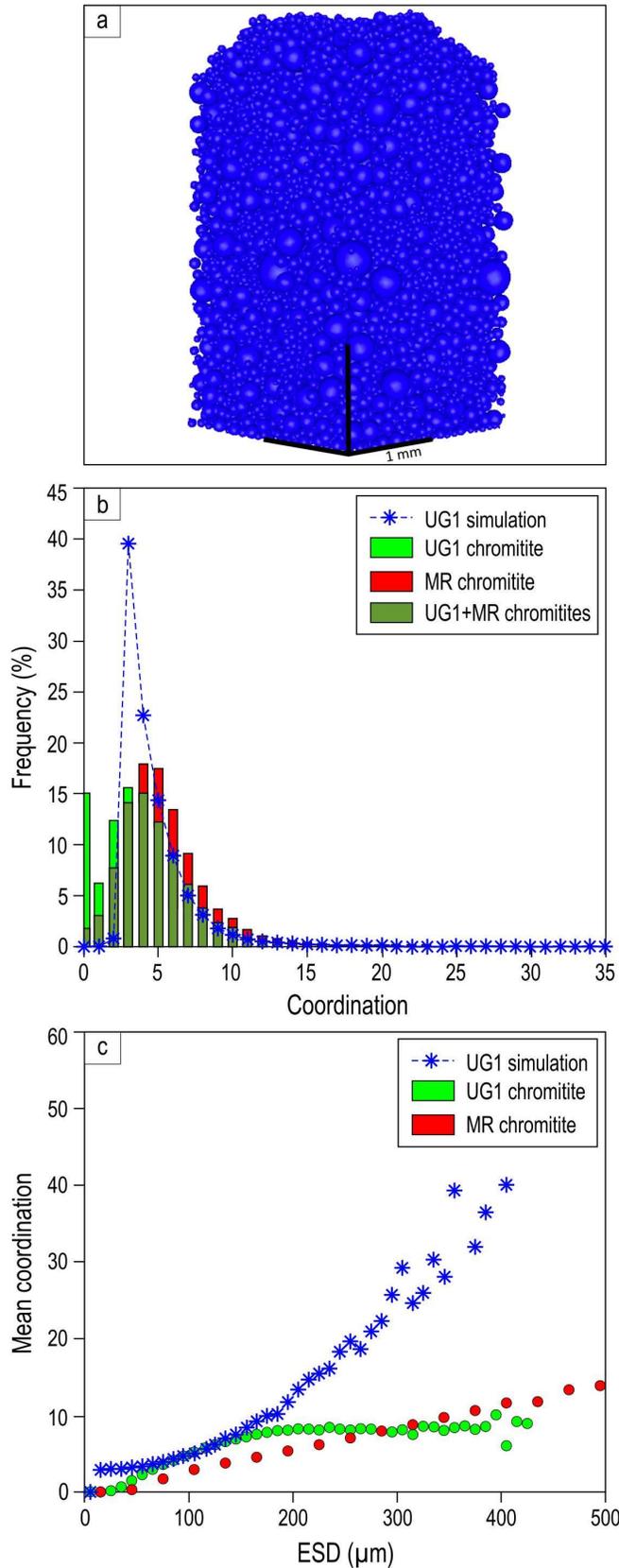
145

146 2; Extended Data Fig. 2; Supplementary Video 1). We can now consider whether these UG1
147 microstructures (Fig. 2) could be generated by random loose packing of non-interacting
148 particles, i.e. from small independent chromite grains settling from a basaltic melt by gravity
149 settling or kinetic sieving⁴³. Theoretical microstructures of random mechanical crystal packs
150 can be predicted by packing theory and characterised by two properties: the packing density
151 (inverse of porosity); and the distribution of coordination numbers, i.e., the number of nearest
152 neighbour grains in contact. Quantification of clustering and chain formation has typically used
153 assumptions of constant grain sizes⁴⁴ but these parameters are known to be sensitive to the
154 particle size distribution⁴⁵. No observations or simulations have been made to date using the
155 characteristic negative log-linear particle size distributions of crystals found in cumulates. To
156 address this gap, we measured chromite particle sizes and coordination number distributions
157 from the segmented HRXCT scan of the UG1 chromitite (Methods, Supplementary Data 1),
158 choosing a volume within which chromite grains are primarily enclosed within plagioclase or
159 pyroxene oikocrysts. This is to eliminate possible effects on microstructure caused by later
160 recrystallisation and annealing. We then compared the results with those of a Discrete Element
161 Method^{59,60} computer simulation of a random loose packing, generated by simple settling, of
162 an assemblage of crystals with the same size distribution as the UG1 sample (Methods). Results
163 show two distinct differences between the synthetic pack and the natural sample (Fig. 3):
164 firstly, the packing densities are greatly different, being much higher in the synthetic pack than
165 in the natural sample (60% vs. 27%). Secondly, the distribution of coordination numbers is
166 significantly different (Fig. 3b, c). In the random pack, coordination number increases
167 exponentially with the grain size. This happens because, for the geometrical reasons, larger
168 grains have a larger surface area and hence are likely to be in contact with a larger number of
169 smaller grains filling space between them. In the natural UG1 data set, however, the
170 coordination number flattens out and remains roughly constant at about 10 for grains larger

171 than 150 microns in size (noting that only a small proportion of the total number of grains falls
172 in this size range). This results from the chromite grains forming an open cage-like or chain-
173 like structure where gaps in the framework are not occupied by other grains, causing lower
174 coordination numbers at larger grain sizes. Furthermore, the natural sample contains a
175 significant proportion of isolated or nearly isolated grains with coordination numbers of 0, 1 or
176 2; these are absent in the simulation. We conclude that (a) the low packing density, (b) the
177 presence of isolated individual grains not supported by contact with any other chromite grains,
178 and (c) the relationship between coordination number and crystal size in the UG1 chromitite
179 are not consistent with random mechanical accumulation of non-interacting chromite grains,
180 be it crystal settling in a melt³⁵⁻³⁸ or kinetic sieving in a crystal mush^{39,40}.

181

182 There remains a possibility that chromite may settle in the form of chromite chains/clusters
183 produced either by heterogeneous nucleation against chromite grains suspended in the
184 convecting melt⁴⁶ or physical collision of isolated chromite crystals ‘swimming together’ in
185 this melt (i.e. “synneusis”)⁴⁷. The accumulation of such clustered chains on the chamber floor
186 may, in principle, give rise to the formation of a continuous 3D framework of touching
187 chromite crystals. However, we reject this scenario on the field evidence: neither individual
188 grains nor clustered chains can settle onto overhanging margins of potholes^{28,34} (Extended Data
189 Fig. 1). Sidewall crystallisation thus indicates that the 3D chromite framework has crystallized
190 *in situ*, i.e., directly at the chamber floor. This may only happen by one physical process –
191 heterogeneous/self-nucleation^{22,46,48} of chromite grains on the floor cumulates^{28,34}. To develop
192 this point further, we also compare the microstructure of a Merensky Reef chromitite seam that



occurs on a vertical to overhanging sidewalls of potholes (Extended Data Figs. 3 and 4; Supplementary Video 1), taken as a definitive example of a microstructure that could only have developed *in situ*^{34,49} (Fig. 3c).

Fig. 3. Results of numerical simulation compared to the observed natural data for the UG1 chromitite showing the contrasting relationship between grain size and coordination number between the randomised loose packing simulation (using the observed UG1 grain size distribution) and the two natural samples, the UG1 chromitite and MR overhanging chromitite. (a) 3D perspective view of a simulation pack of spheres having the same size distribution as chromite grains in the UG1 chromitite; (b) Histogram showing the distribution of coordination number in UG1 simulation, UG1 chromitite and MR overhanging chromitite; (c) Plot showing mean coordination number of all grains within each size range bin, as a function of size range of chromite grains in the UG1 simulation, UG1 chromitite and MR overhanging chromitite (ESD: equivalent sphere diameter).

226 The packing density is likewise lower than the random loose packing simulation (53% vs 60%)
227 (Supplementary Data 1). The coordination number vs grain size curve for this seam shows a
228 steady linear increase with grain size, but, like the UG1 chromitite, has systematically lower
229 coordination number values than the simulation along the entire length of the trend (Fig. 3c).
230 Significantly, the packing and coordination number characteristics of the UG1 chromitite are
231 much closer to those of the indisputably in-situ crystallised sidewall Merensky Reef chromitite
232 than to the random packing simulation. We deduce that the relatively low coordination number
233 values and low packing densities of both natural samples are the result of *in situ* growth of
234 chromite chains or cages by heterogeneous self-nucleation. We also conclude that chromite
235 grains (Fig. 1d) are not able to settle freely towards the chamber floor simply because they are
236 all bound together in self-supporting frameworks attached to the floor.

237

238 **A scenario for *in situ* growth of chromite on the chamber floor**

239 *In situ* growth of chromite requires crystallisation from a parental melt that was saturated in
240 chromite as the only liquidus phase. Such melts can be produced in response to decompression
241 during their ascent from a deep staging reservoir towards Earth's surface³⁰. We propose that
242 the Bushveld chamber has been replenished by such melts as basal flows that caused
243 thermochemical erosion of the floor cumulates²⁸, including the excavation of potholes (Fig.
244 4a). Upon cooling, the melt became saturated in chromite only³⁰ (Fig. 4b), with the first
245 chromite grains being nucleated heterogeneously on pre-existing plagioclase crystals of the
246 floor anorthosites. With further cooling, chromite started preferentially self-nucleating on
247 earlier-formed chromite grains to produce composite 3D clusters which subsequently merged
248 into a continuous 3D framework of touching chromite grains (Fig. 4c). New crystals emerged
249 in the system mostly by self-nucleation because the activation energy for this process is much
250 lower relative to other types of nucleation^{22,46}. A small portion of crystals (3 vol.%) that occur

251 as entirely discrete grains (Fig. 2d) has likely formed by homogeneous nucleation in the
252 interstitial space. We envisage that chemical differentiation of the resident melt in the chamber
253 at that time occurred by convective removal of a buoyant compositional boundary layer⁵⁰ from
254 *in situ* growing chromite crystals in a 3D framework (Fig. 4c). The differentiation is aided by
255 high porosity and permeability of a 3D crystal framework that permits the easy chemical
256 exchange of melts between the crystal framework and the main magma body. The remarkable
257 preservation of the nucleation/growth history in the UG1 chromitite is due to the early growth
258 of plagioclase and pyroxene oikocrysts which have ‘frozen in’ a 3D chromite framework at its
259 early immature stage. This continuous 3D framework of chromite grains (Fig. 2) is the first
260 documented example showing how a natural solidification front looks when it develops via
261 self-nucleation/heterogeneous nucleation⁴⁸. Under other circumstances, the initial framework
262 would have evolved into perfect chromitite adcumulate (up to 100% chromite; e.g. LG
263 chromitite in Extended Data Fig. 5) in which all primary information is lost. We propose that
264 most layers of monomineralic chromitites in the Bushveld Complex started their life as porous,
265 *in situ* produced 3D chromite frameworks that can no longer be seen due to adcumulus
266 overgrowth⁵¹. We further propose that the observed shape of the coordination number vs grain
267 size curve (Fig. 3c) may be diagnostic of chromite cumulates formed by *in situ* crystallisation
268 involving heterogeneous self-nucleation^{28,34}.

269

270 **Chromium budget requires a large magma volume**

271 The realization that massive chromitites form by *in situ* growth of chromite directly in the
272 chamber (Fig. 4a) – rather than from chromite phenocrysts brought into the chamber with
273 externally-derived crystal-rich mushes^{9,12,38,40} – logically brings us to a long-known Cr mass
274 balance issue^{32,52,53}. The stratiform chromitite layers in layered intrusions can be up to 2 m
275 thick and contain 40-50 wt.% Cr₂O₃, yet have evidently crystallized from a basaltic melt that

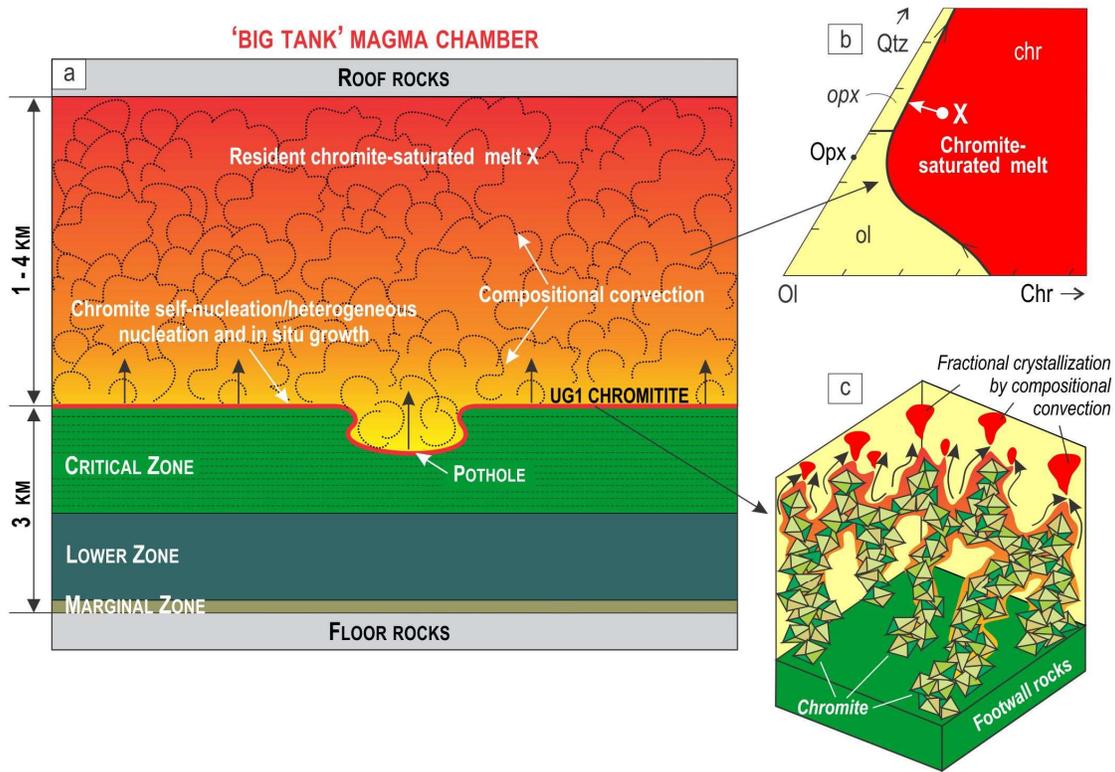
276 was unlikely to have contained more than 1000 ppm Cr³¹. An implication is that the formation
277 of a thick chromitite layer, such as the UG1 chromitite, requires extraction of Cr from a very
278 large volume of liquid that can be present either as a thick melt layer in the chamber or as the
279 melt flowing through the chamber, or both. To illustrate, given the thickness of a chromite
280 layer, from mass-balance considerations one can calculate the volume of the parental melt in
281 terms of the thickness of an equivalent layer. If a chromite layer crystallized at 1250°C and the
282 fO_2 equivalent to QFM buffer, then based on experimental data^{31,52}, the chromite should contain
283 ~45% Cr₂O₃, and the coexisting melt should carry about 0.10 wt% Cr₂O₃. Assuming that
284 chromitites formed from the overlying melt, it can be estimated that a 1 m thick layer of
285 chromitite will require a magma column of about 4 km thick (Fig. 4a). The thickness can be
286 reduced to 2 km³² or 1 km²⁶ if Cr solubility in a parental melt is to increase by its higher
287 temperature or lower fO_2 ³². These estimations assume 30% of the Cr removal from a parental
288 melt³². One cannot remove any more Cr from the melt than that because otherwise the melt
289 will reach a cotectic with other liquidus phases (e.g., olivine or orthopyroxene) terminating
290 chromitite formation (Fig. 4b).

291

292 This one-dimensional modelling illustrates the mass-balance issue. Applying this logic to the
293 Bushveld Complex, it has been estimated that the formation of its most prominent chromitites
294 would have required a column of 13 to 15 km of a parental chromite-saturated liquid^{25,26}. In
295 addressing this mass-balance requirement, Cawthorn and Walraven²⁵ modelled the Bushveld
296 chamber as a long-lived flow-through system (~75,000 years life-time) that developed via a
297 large number of injection events, partial crystallization of these magma batches, and removal
298 of their residual liquids from the chamber by the succeeding magma batches. They concluded
299 that the total volume of basaltic magma involved was 740-1200*10³ km³, with only ~50% of
300 this being represented by the cumulates now seen within the Bushveld Complex²⁵. The excess

301 magma has likely escaped from the chamber laterally²⁶ to form mafic intrusions, such as the
 302 Molopo Farms Complex located about 200 km west of the Bushveld Complex^{54,55}.

303



304

305 **Fig. 4. A 'big-tank' chamber of the Bushveld Complex filled with a resident melt that crystallizes into the**
 306 **UG1 chromitite at the chamber floor. a, Schematic cartoon of the Bushveld chamber that shows crystallization**
 307 **of the UG1 chromitite near the top of the Critical Zone. The formation of a 1 m thick layer of such chromitite**
 308 **requires an equivalent layer of chromite-only-saturated melt of at least 1 to 4 km in thickness^{31,32}. This large**
 309 **volume of liquid may be present in the chamber either as a melt column or as a melt that flows through the chamber**
 310 **for a long period of time. The chromite nucleates and crystallizes directly on the chamber floor, with the resident**
 311 **melt convecting turbulently to deliver Cr for *in situ* chromite growth. b, Ol–Chr–Qtz phase diagram illustrating**
 312 **the position of a chromite-only-saturated melt parental to the UG1 chromitite. The diagram is modified from**
 313 **reference³⁵. Ol, olivine; Opx, orthopyroxene, Chr, chromite; Qtz, quartz. c, A close-up view of the immature UG1**
 314 **chromitite that forms a 3D framework of touching chromite crystals which self-nucleate on the floor of a magma**
 315 **chamber. Note the low density packing and the low coordination number of chromite in the framework. A**
 316 **compositional boundary layer of buoyant liquid is produced around crystallizing chromite clusters that migrates**
 317 **towards their apex and is released into the overlying melt in the form of compositional plumes, thus causing**
 318 **chemical differentiation in the resident melt.**

319

320

321 **A classical magma chamber paradigm affirmed**

322 The enormous lateral extent of *in situ* formed chromitite layers and related mass-balance
323 considerations indicate that during the formation of massive chromitites the Bushveld chamber
324 has been operating as a giant magma body of more than 400 km in diameter, with a column of
325 the resident melt likely attaining a few km in thickness. Thus, starting from this stage the
326 Bushveld Complex has been developed as a large, long-lived and largely molten magma
327 chamber (a true ‘big tank’ reservoir) in Earth’s crust (Fig. 4a). The conclusion is further
328 supported by the remarkable homogeneity of Sr isotopes over an interval of more than 2.5 km
329 of the Upper Zone⁵⁶, which indicates a melt column thickness in the chamber being that thick
330 or even thicker^{57,58}. This is in contrast with the recent assertion, mostly based on out-of-
331 sequence geochronology^{9,13}, that depicts this giant complex as a stack of thin crystal-rich
332 sills^{9,11,13}. Field relationships indicate, however, that zircon isotopic ages in these studies were
333 almost certainly misinterpreted⁵⁹. We thus argue that the field and textural evidence from
334 massive chromitites disprove this radical re-interpretation of the Bushveld Complex. Our
335 inference may be extended to all large mafic-ultramafic layered intrusions that contained thick
336 and laterally extensive layers of monomineralic chromitites (e.g., Stillwater and Great Dyke).
337 Such intrusions are quite rare in the crust through the whole of geological time, so it is not
338 surprising that there are no known examples of equivalent magma chambers that are active and
339 detectable in the present-day Earth’s crust^{4,8}. We conclude that it is too early to discard the
340 classical paradigm of a magma chamber developed by several generations of petrological
341 luminaries¹⁴⁻¹⁹. Rather, we suggest re-directing our efforts to find out how new geophysical,
342 geochronological and thermal/diffusion modelling¹⁻¹³ can be logically reconciled with the
343 classical paradigm.

344

345

346 **Online content**

347 Any methods, additional references, Nature Research reporting summaries, source data,
348 extended data, supplementary information, acknowledgements, peer review information;
349 details of author contributions and competing interests; and statements of data and code
350 availability are available at.....

351

352 **References**

353

- 354 1. Glazner, A. F., Coleman, D. S., Gray, W. & Taylor, R. Z. Are plutons assembled over millions of
355 years by amalgamation from small magma chambers? *GSA TODAY* **14**, 4–12 (2004).
- 356 2. Coleman, D. S., Gray, W. & Glazner, A. F. Rethinking the emplacement and evolution of zoned
357 plutons: geochronologic evidence for incremental assembly of the Tuolumne Intrusive Suite,
358 California. *Geol* **32**, 433 (2004).
- 359 3. Bachmann, O. & Huber, C. The inner workings of crustal distillation columns; the physical
360 mechanisms and rates controlling phase separation in silicic magma reservoirs. *Journal of*
361 *Petrology* **60**, 3–18 (2019).
- 362 4. Cashman, K. V., Sparks, R. S. J. & Blundy, J. D. Vertically extensive and unstable magmatic
363 systems: a unified view of igneous processes. *Science* **355**, eaag3055 (2017).
- 364 5. Cooper, K. M. What does a magma reservoir look like? The “crystal’s-eye” view. *Elements* **13**,
365 23–28 (2017).
- 366 6. Edmonds, M., Cashman, K. V., Holness, M. & Jackson, M. Architecture and dynamics of magma
367 reservoirs. *Phil. Trans. R. Soc. A.* **377**, 20180298 (2019).
- 368 7. Jackson, M. D., Blundy, J. & Sparks, R. S. J. Chemical differentiation, cold storage and
369 remobilization of magma in the Earth’s crust. *Nature* **564**, 405–409 (2018).
- 370 8. Sparks, R. S. J. *et al.* Formation and dynamics of magma reservoirs. *Phil. Trans. R. Soc. A.* **377**,
371 20180019 (2019).
- 372 9. Mungall, J. E., Kamo, S. L. & McQuade, S. U–Pb geochronology documents out-of-sequence
373 emplacement of ultramafic layers in the Bushveld Igneous Complex of South Africa. *Nat*
374 *Commun* **7**, 13385 (2016).
- 375 10. Wall, C. J. *et al.* The Stillwater Complex: integrating zircon geochronological and geochemical
376 constraints on the age, emplacement history and crystallization of a large, open-system layered
377 intrusion. *Journal of Petrology* **59**, 153–190 (2018).
- 378 11. Robb, S. J. & Mungall, J. E. Testing emplacement models for the Rustenburg Layered Suite of the
379 Bushveld Complex with numerical heat flow models and plagioclase geospeedometry. *Earth and*
380 *Planetary Science Letters* **534**, 116084 (2020).
- 381 12. Yao, Z., Mungall, J. E. & Jenkins, M. C. The Rustenburg Layered Suite formed as a stack of mush
382 with transient magma chambers. *Nat Commun* **12**, 505 (2021).
- 383 13. Scoates, J. S. *et al.* Dating the Bushveld Complex: timing of crystallization, duration of
384 magmatism, and cooling of the world’s largest layered intrusion and related rocks. *Journal of*
385 *Petrology* (2021) doi:10.1093/petrology/egaa107.
- 386 14. Daly, R. A. The nature of volcanic action. *Proceedings of the American Academy of Arts and*
387 *Sciences* **47**, 48–119 (1911).
- 388 15. Bowen, N. L. *The evolution of the igneous rocks.* (Oxford University Press, 1928).
- 389 16. Wager, L. R. & Deer, W. A. Geological Investigations in East Greenland, Part III: The Petrology of
390 the Skaergaard Intrusion, Kangerdlugssuaq, East Greenland. *Med. Greenland* **105**, 352 (1939).
- 391 17. Wager, L. R. & Brown, G. M. *Layered igneous rocks.* (Oliver & Boyd, Edinburgh & London, 1968).

- 392 18. Parsons, I. *Origins of Igneous Layering*. vol. 196 (Springer, Dordrecht, 1987).
- 393 19. Cawthorn, R. G. *Layered Intrusions*. vol. 15 (Elsevier, Amsterdam, 1996).
- 394 20. Charlier, B., Namur, O., Latypov, R. & Tegner, C. *Layered Intrusions*. (Springer Netherlands,
- 395 2015).
- 396 21. Marsh, B. D. Solidification fronts and magmatic evolution. *Mineralogical Magazine* **60**, 5–40
- 397 (1996).
- 398 22. Campbell, I. H. Fluid Dynamic Processes in Basaltic Magma Chambers. in *Developments in*
- 399 *Petrology* vol. 15 45–76 (Elsevier, 1996).
- 400 23. Gudmundsson, A. Magma chambers: formation, local stresses, excess pressures, and
- 401 compartments. *Journal of Volcanology and Geothermal Research* **237–238**, 19–41 (2012).
- 402 24. Cawthorn, R. G. The Bushveld Complex, South Africa. in *Layered Intrusions* (eds. Charlier, B.,
- 403 Namur, O., Latypov, R. & Tegner, C.) 517–587 (Springer Netherlands, 2015). doi:10.1007/978-94-
- 404 017-9652-1_12.
- 405 25. Cawthorn, R. G. & Walraven, F. Emplacement and crystallization time for the Bushveld Complex.
- 406 *Journal of Petrology* **39**, 1669–1687 (1998).
- 407 26. Naldrett, A. J., Wilson, A., Kinnaird, J., Yudovskaya, M. & Chunnett, G. The origin of chromitites
- 408 and related PGE mineralization in the Bushveld Complex: new mineralogical and petrological
- 409 constraints. *Mineralium Deposita* **47**, 209–232 (2012).
- 410 27. Schurmann, L. W., Grabe, P.-J. & Steenkamp, C. J. Chromium. in *The Mineral Resources of South*
- 411 *Africa* vol. 16 90–105 (Council for Geosciences, 1998).
- 412 28. Latypov, R., Chistyakova, S. & Mukherjee, R. A novel hypothesis for origin of massive chromitites
- 413 in the Bushveld Igneous Complex. *Journal of Petrology* **58**, 1899–1940 (2017).
- 414 29. Pebane, M. & Latypov, R. The significance of magmatic erosion for bifurcation of UG1 chromitite
- 415 layers in the Bushveld Complex. *Ore Geology Reviews* **90**, 65–93 (2017).
- 416 30. Latypov, R. *et al.* Platinum-bearing chromite layers are caused by pressure reduction during
- 417 magma ascent. *Nature Communications* **9**, (2018).
- 418 31. Murck, B. W. & Campbell, I. H. The effects of temperature, oxygen fugacity and melt
- 419 composition on the behaviour of chromium in basic and ultrabasic melts. *Geochimica et*
- 420 *Cosmochimica Acta* **50**, 1871–1887 (1986).
- 421 32. Campbell, I. H. & Murck, B. W. Petrology of the G and H Chromitite Zones in the Mountain View
- 422 Area of the Stillwater Complex, Montana. *Journal of Petrology* **34**, 291–316 (1993).
- 423 33. Cawthorn, R. G. The platinum group element deposits of the Bushveld Complex in South Africa.
- 424 *platin met rev* **54**, 205–215 (2010).
- 425 34. Latypov, R., Chistyakova, S., Barnes, S. J. & Hunt, E. J. Origin of platinum deposits in layered
- 426 intrusions by in situ crystallization: Evidence from undercutting Merensky Reef of the Bushveld
- 427 Complex. *Journal of Petrology* (2017) doi:10.1093/petrology/egx032.
- 428 35. Irvine, T. N. Origin of chromitite layers in the Muskox intrusion and other stratiform intrusions: A
- 429 new interpretation. *Geology* **5**, 273–277 (1977).
- 430 36. Spandler, C., Mavrogenes, J. & Arculus, R. Origin of chromitites in layered intrusions: Evidence
- 431 from chromite-hosted melt inclusions from the Stillwater Complex. *Geology* **33**, 893–896 (2005).
- 432 37. Kinnaird, J. A., Kruger, F. J., Nex, P. A. M. & Cawthorn, R. G. Chromitite formation—a key to
- 433 understanding processes of platinum enrichment. *Applied Earth Science* **111**, 23–35 (2002).
- 434 38. Leshner, C. M., Carson, H. J. E. & Houlé, M. G. Genesis of chromite deposits by dynamic upgrading
- 435 of Fe ± Ti oxide xenocrysts. *Geology* **47**, 207–210 (2019).
- 436 39. Maier, W. D., Barnes, S.-J. & Groves, D. I. The Bushveld Complex, South Africa: formation of
- 437 platinum–palladium, chrome- and vanadium-rich layers via hydrodynamic sorting of a mobilized
- 438 cumulate slurry in a large, relatively slowly cooling, subsiding magma chamber. *Mineralium*
- 439 *Deposita* **48**, 1–56 (2013).
- 440 40. Mondal, S. K. & Mathez, E. A. Origin of the UG2 chromitite layer, Bushveld Complex. *Journal of*
- 441 *Petrology* **48**, 495–510 (2007).

- 442 41. Jackson, E. D. *Primary textures and mineral associations in the ultramafic zone of the Stillwater*
443 *complex, Montana. Professional Paper 106* <https://pubs.er.usgs.gov/publication/pp358> (1961)
444 doi:10.3133/pp358.
- 445 42. Eales, H. V. & Reynolds, I. M. Cryptic variations within chromitites of the upper critical zone,
446 northwestern Bushveld Complex. *Economic Geology* **81**, 1056–1066 (1986).
- 447 43. Delaney, G. W., Hilton, J. E. & Cleary, P. W. Defining random loose packing for nonspherical
448 grains. *Phys. Rev. E* **83**, 051305 (2011).
- 449 44. Jerram, D. A. Quantifying the Building Blocks of Igneous Rocks: Are Clustered Crystal
450 Frameworks the Foundation? *Journal of Petrology* **44**, 2033–2051 (2003).
- 451 45. He, D., Ekere, N. N. & Cai, L. Computer simulation of random packing of unequal particles. *Phys.*
452 *Rev. E* **60**, 7098–7104 (1999).
- 453 46. Campbell, I. H. Some problems with the cumulus theory. *Lithos* **11**, 311–323 (1978).
- 454 47. Philpotts, A. R., Shi, J. & Brustman, C. Role of plagioclase crystal chains in the differentiation of
455 partly crystallized basaltic magma. *Nature* **395**, 343–346 (1998).
- 456 48. Latypov, R. M., Chistyakova, S. Yu., Namur, O. & Barnes, S. Dynamics of evolving magma
457 chambers: textural and chemical evolution of cumulates at the arrival of new liquidus phases.
458 *Earth-Science Reviews* **210**, 103388 (2020).
- 459 49. Latypov, R., Chistyakova, S., Page, A. & Hornsey, R. Field evidence for the in situ crystallization of
460 the Merensky Reef. *Journal of Petrology* egv023 (2015) doi:10.1093/petrology/egv023.
- 461 50. Kruger, W. & Latypov, R. Fossilized solidification fronts in the Bushveld Complex argue for liquid-
462 dominated magmatic systems. *Nat Commun* **11**, 2909 (2020).
- 463 51. Campbell, I. H. Distribution of orthocumulate textures in the Jimberlana intrusion. *The Journal of*
464 *Geology* **95**, 35–53 (1987).
- 465 52. Barnes, S. J. The distribution of chromium among orthopyroxene, spinel and silicate liquid at
466 atmospheric pressure. *Geochimica et Cosmochimica Acta* **50**, 1889–1909 (1986).
- 467 53. Eales, H. V. Implications of the chromium budget of the Western Limb of the Bushveld Complex.
468 *South African Journal of Geology* **103**, 141–150 (2000).
- 469 54. Prendergast, M. D. The Molopo Farms Complex, southern Botswana - a reconsideration of
470 structure, evolution, and the Bushveld connection. *South African Journal of Geology* **115**, 77–90
471 (2012).
- 472 55. Kaavera, J., Rajesh, H. M., Tsunogae, T. & Belyanin, G. A. Marginal facies and compositional
473 equivalents of Bushveld parental sills from the Molopo Farms Complex layered intrusion,
474 Botswana: Petrogenetic and mineralization implications. *Ore Geology Reviews* **92**, 506–528
475 (2018).
- 476 56. Kruger, F. J., Cawthorn, R. G. & Walsh, K. L. Strontium isotopic evidence against magma addition
477 in the Upper Zone of the Bushveld Complex. *Earth and Planetary Science Letters* **84**, 51–58
478 (1987).
- 479 57. Vantongerren, J. A., Mathez, E. A. & Kelemen, P. B. A felsic end to Bushveld differentiation.
480 *Journal of Petrology* **51**, 1891–1912 (2010).
- 481 58. Vantongerren, J. A. & Mathez, E. A. Incoming Magma Composition and Style of Recharge below
482 the Pyroxenite Marker, Eastern Bushveld Complex, South Africa. *Journal of Petrology* **54**, 1585–
483 1605 (2013).
- 484 59. Latypov, R. M. Misleading interpretation of zircon ages in layered intrusions: A discussion on
485 “Dating the Bushveld Complex: Timing of Crystallization, Duration of Magmatism, and Cooling of
486 the World’s Largest Layered Intrusion and Related Rocks” by James S. Scoates et al. (*J. Petrology*,
487 2021). (2021).

490

491 **Methods**

492 **Rock sampling and petrography**

493 Documentation of field observations of the UG1 chromitite was undertaken at the Dwars River
494 locality and its sampling at the nearby Mototolo Mine from the HEX 076 drill-core in the
495 Eastern Bushveld Complex. Thin sections and polished blocks were cut from orientated sample
496 blocks to be as close as possible to the original vertical position. Thin sections and polished
497 blocks were studied using a polarised light microscope with a circular stage and photographed
498 using the Olympus 224 BX-63 OM/FM optical microscope housed at the MMU (Microscopy
499 and Microanalysis Unit) of the University of the Witwatersrand, Johannesburg, South Africa.

500

501 **High resolution X-ray computed tomography and 3D image analysis and quantification**

502 The UG1 sample was scanned using the Zeiss Versa XRM 520 3D x-ray microscope installed
503 at the Australian Resources Research Centre (CSIRO Mineral Resources, Kensington,
504 Western-Australia). The instrument was set to maximize the contrast between chromite and
505 silicates (plagioclase and pyroxene) present in the sample. Two scans at a voxel size of 5 μm
506 were performed along the vertical axis of the samples and were stitched in 3D to maximize the
507 volume of sample used for further analysis. A total of 1601 projections were recorded over
508 360° degrees rotation for each scan and were used for volume reconstruction. Beam hardening
509 and ring artefacts were minimized during data acquisition and corrected (if necessary) during
510 image reconstruction. After reconstruction, the sample is represented by a regular grid (1998 x
511 2046 x 3748 voxels) where each voxel has a unique greyscale value. Chromite, plagioclase and
512 pyroxene were segmented from the volumes using a 3D gradient watershed algorithm⁶⁰ to
513 produce binary images. The separation of touching chromite crystals in 3D was done using a
514 modified version of the algorithm used to separate touching chromite in komatiites⁶¹ and
515 chromite from the normal Merensky Reef⁶² using Avizo2020.1TM and MatlabTM software. The

516 shape and size characteristics of chromite network and individual grains were computed to
517 provide quantitative measure of chromite grains above 15 μm equivalent sphere diameter
518 (ESD). Chromite grains were defined as touching each other using an 18-voxel connectivity
519 threshold (i.e., voxels are connected if their faces or edges touch). The coordination number of
520 each chromite grains (i.e., the total number of other chromite grains touching in 3D a given
521 grain) was also calculated. All results are summarized in Figs. 2 and 3.

522

523 **Random packing simulation**

524 The simulated Random Packing of UG1 chromite was generated using the Discrete Element
525 Method as described in ref⁶³. The individual chromite particles were modelled as spheres with
526 a size distribution as measured from the UG1 chromite sample ranging from 15 μm to 420 μm ,
527 and an interparticle friction coefficient of 0.9. The simulation box has dimensions 5 mm by 2
528 mm by 2 mm, with periodic boundaries in the two directions normal to gravity. Particles are
529 initially distributed randomly in the simulation box and then allowed to slowly settle under
530 gravity subject to a Stokes' drag force to form a random loose packing⁴³. The packing is then
531 analysed to determine the packing density and the distribution of inter-particle contacts
532 between particles.

533

534 **References**

535

- 536 60. Godel, B. High-Resolution X-Ray Computed Tomography and Its Application to Ore Deposits:
537 From Data Acquisition to Quantitative Three-Dimensional Measurements with Case Studies from
538 Ni-Cu-PGE Deposits. *Economic Geology* **108**, 2005–2019 (2013).
- 539 61. Godel, B., Barnes, S. J., Gürer, D., Austin, P. & Fiorentini, M. L. Chromite in komatiites: 3D
540 morphologies with implications for crystallization mechanisms. *Contrib Mineral Petrol* **165**, 173–
541 189 (2013).
- 542 62. Vukmanovic, Z., Barnes, S. J., Reddy, S. M., Godel, B. & Fiorentini, M. L. Morphology and
543 microstructure of chromite crystals in chromitites from the Merensky Reef (Bushveld Complex,
544 South Africa). *Contributions to Mineralogy and Petrology* **165**, 1031–1050 (2012).
- 545 63. Cleary, P. W. Large scale industrial DEM modelling. *Engineering Computations* **21**, 169–204
546 (2004).

547

548 **Data availability**

549 The authors declare that all relevant data are available within the article and its Supplementary
550 Information Files.

551

552 **Acknowledgements** Several versions of this manuscript benefited from the constructive peer reviews by
553 Catherine Annen, Allen Glazner, Chris Hawkesworth, Olivier Namur, Jill VanTongeren, Zoja Vukmanovic,
554 Steve Prevec and Willem Kruger. The study was supported by research grants to R.L. from the National
555 Research Foundation (NRF) and Department of Science and Technology (DST)-NRF Centre of Excellence
556 for Integrated Mineral and Energy Resource Analysis (CIMERA) of South Africa. V.J.R. was supported by
557 the DST-NRF Centre of Excellence in Palaeosciences (CoE-Pal), and the Palaeontological Scientific Trust
558 (PaST). Any opinion, finding and conclusion or recommendation expressed in this contribution is that of the
559 authors and the DST-NRF CIMERA, CoE-Pal, PaST and NRF do not accept any liability in this regard. We
560 thank an artist Inna Laur (Finland) for helping us depict the 3D framework of touching chromite crystals in
561 Fig. 4c.

562

563 **Author Contributions** R.L. and S.C. undertook field work, mapping and geochemical sampling of outcrops
564 as well as conceptualized the original idea and wrote a draft of the paper. V.R. and K.J. performed X-ray
565 microtomography of samples and data processing, took part in conceptualizing the original idea and
566 participated to the text writing. B.G. collected, processed, and analysed the HRXCT data. G. D., P. C. and
567 S.B. have conducted the randomised loose packing simulation of chromite grains in the UG1 chromitite.
568 S.B. and I.C. participated in interpretation of data, editing the paper as well as in improving clarity of figures.
569 All co-authors discussed the results and problems and contributed to producing a final draft for peer reviews.

570

571 **Supplementary Information** is available in the online version of the paper.

572

573 **Competing interests** The authors declare no competing interests.

574

Supplementary Files

This is a list of supplementary files associated with this preprint. Click to download.

- [SupplementaryData1.xlsx](#)
- [ExtendedData.pdf](#)
- [SupplementaryVideo1UG1chromitite.zip](#)
- [SupplementaryVideo2MRchromitite.zip](#)