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Ground motion parameters for the 2015 Nepal Earthquake

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Availability of data and material

Strong-motion data of the Gorkha earthquake sequence from the Central Indo-Gangetic Network (CIGN) are obtained collaboratively from D. S. Srinagesh at srinagesh@ngri.res.in. The ground motion records for stations KTP, PTN, TVU, and THM were made available by CESMD (Center for Engineering Strong Motion Data) (Takai et al. 2016), whereas station KATNP was made available by USGS (USGS, 2015b).

Code availability

Custom codes in MATLAB are used in the current study.

Authors' contributions

Formal analysis, investigation and writing-original draft preparation: Jahnabi Basu, Writing-review and editing and Supervision: Bhargavi Podili, Conceptualization, Methodology and Supervision: S. T. G. Raghukanth, Data provision: D Srinagesh. All authors read and approved the final manuscript.

Ethics approval

Compliance with Ethical Standards

No Human Participants and/or Animals are involved in the research.

Consent to participate

Not applicable

Consent to publication

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37 **ABSTRACT**

38

39 The 25th April 2015 Nepal earthquake is the first major event in the Himalayan orogeny that
40 provides a relatively well recorded dataset. This paper presents a comprehensive analysis of
41 the mainshock and its five major aftershocks through 21 well established ground motion
42 parameters. The analysis is presented for near field stations of the Kathmandu basin and far
43 field stations of the Indo-Ganga basin, including the site response behavior with varying
44 sediment thickness. In addition, a new ground motion model is derived for all the 21
45 parameters using moment magnitude, rupture distance, site class and sediment depth as
46 predictor variables.

47

48 *Keywords: Ground Motion Parameters; Rupture distance; Sediment depth; Instantaneous*
49 *Frequency and Ground Motion Prediction Equation*

50

51

52 **1. INTRODUCTION**

53

54 The Himalayan tectonic zone of the Indian subcontinent has been constantly subjected to
55 incessant seismic activities, wherein a total of eight major earthquakes of $M_w > 7$ are recorded
56 in the past two centuries. These include the 1866 Nepal (M_w 7.4), 1897 Shillong (M_w 8.1), the
57 1905 Kangra (M_w 7.8), the 1934 Bihar-Nepal (M_w 8.1), the 1947 Arunachal Pradesh (M_w
58 7.9), the 1950 Arunachal Pradesh (M_w 8.6), the 2005 Muzafferapur (M_w 7.6) and the 2015
59 Nepal (M_w 7.9) earthquakes. The most common and unfortunate aspect among all these
60 events is that none of these earthquakes are recorded, until the M_w 7.9 Nepal earthquake, thus
61 a thorough study of this 2015 event gained extreme importance. The 25th April 2015 Nepal
62 earthquake originated near the Main Frontal Thrust of the Himalayas, which is at the
63 subduction plate boundary between the Indian and Eurasian plates. The rupture originated in
64 the Gorkha district (28.147°N and 84.708°E), which is nearly 80 km to the northwest of the
65 capital city – Kathmandu and propagated eastwards for about 160 km. Ground failure was
66 reported in many densely populated areas, where the main event triggered landslides and
67 ground liquefaction in the areas surrounding Kathmandu. There was unprecedented damage
68 to buildings and heritage structures, where the economic loss is nearly 10 billion USD, which
69 is about half the GDP of Nepal and the human loss includes over 9000 fatalities. Several
70 moderate to strong aftershocks followed this main event. According to USGS, 227
71 earthquakes of $M_w > 4.1$ were recorded within 250km radius of the main event epicentre and
72 in three weeks prior and after the main event. Among these, there are 30 aftershocks of $M_w >$
73 5 and five aftershocks of $M_w > 6$ due to the M_w 7.9 rupture. In order to perform a thorough
74 study of the region, the five major aftershocks are also included in the analysis.

75

76 In the immediate aftermath, many studies were conducted on source rupture characteristics
77 (Hayes *et al.*, 2015, Galetzka *et al.*, 2015, Grandin *et al.*, 2015) and on artificial simulation of
78 ground motions (Koketsu *et al.*, 2016, Dhanya *et al.*, 2017). Some of the important
79 observations drawn from these studies were the significant influence of rupture directivity
80 and basin effects on ground motions. Among the few studies on ground motion datasets,

81 Goda *et al.* (2015) attributed the main cause of structural damage and collapse of buildings to
82 the short period peak, which is due to ground vibration induced by near source rupture.
83 Dhakal *et al.* (2016) confirmed that the basin sediments at the KATNP station amplified the
84 long period components of ground motions and have concluded that the GMPEs of the
85 similar tectonic environment such as Japan underestimate the recorded peak and spectral
86 acceleration values. The ground motion study conducted by Takai *et al.* (2016) with the five
87 time histories (KATNP, PTN, THM, KTP and TVU) in the Kathmandu valley has also
88 emphasized the amplification due to valley response of sedimentary sites. These studies
89 therefore emphasize the effect of the sedimentary basin response in the Kathmandu valley.
90 The ground motions of the Nepal earthquake and its aftershocks were also recorded by the
91 Central Indo-Gangetic Plains Network (CIGN) and the Indian Institute of Technology,
92 Roorkee (IITR) networks of the Indo-Gangetic Plain (IGP). Due to the presence of substantial
93 alluvial deposits, significant site amplifications in the IGP sites were noticed previously
94 (Srinagesh *et al.*, 2011; Singh *et al.*, 2020). Therefore, it would be interesting to study the site
95 response of both the Kathmandu valley and the IGP. Chadha *et al.* (2016) and Rajaure *et al.*
96 (2016) have obtained site amplifications for some of the IGP and Kathmandu stations for the
97 Nepal earthquake, using a Fourier based approach. However, an advanced non-stationary
98 approach including the evolutionary characteristics of the ground motion would yield better
99 results. Moreover, higher order parameters such as the Root-Mean-Square acceleration
100 (a_{RMS}), Cumulative absolute velocity (CAV), Arias Intensity (I_a) etc., and spectral parameters
101 extracted from Fourier spectrum, Power spectral density and evolutionary power spectral
102 density functions would extract important information from the ground motions, that was not
103 explored previously. Therefore, in the present study, the ground motions of the Nepal
104 earthquake and its aftershocks are studied using a total of 21 ground motion parameters
105 (GMPs) representing various amplitude, frequency and duration characteristics of the ground
106 motion.

107
108 Many local and regional ground motion models are available for the Himalayan region of
109 India (Sharma *et al.*, 2009, Nath *et al.*, 2005, 2012, Raghukanth and Iyengar, 2006, 2007,
110 Joshi *et al.*, 2013 etc.). Among these models, very few ground motion prediction equations
111 (GMPEs) are derived using strong motion datasets (Sharma *et al.*, 2009, Bajaj and
112 Anbazhagan, 2019) and the ground motion datasets that have been used to derive many of
113 these GMPEs via regression analysis, does not include ground motions of an earthquake of
114 magnitude M_w 7.9. Moreover, among the few available GMPEs that are derived by including
115 the Nepal earthquake data (Singh *et al.*, 2017), most were obtained for standard GMPs such
116 as PGA, PGV and Spectral accelerations. There are few global models available for higher
117 order parameters such as CAV and I_a (Campbell and Bozorgnia, 2010, Foulser-Piggot and
118 Goda, 2014). However, since these models are derived for a different tectonic environment,
119 the applicability as well as the goodness of these models to the Indian scenario is debatable.
120 Hence, there is a necessity to generate a new set of GMPEs using the Nepal mainshock and
121 aftershock datasets, especially for the higher order parameters. Therefore, this paper presents
122 a thorough analysis of the 2015 Nepal earthquake and its aftershocks through 21 well
123 established GMPs. The behavior of these GMPs with distance from the source, distance,
124 sediment depth and site characteristics is also studied. Since it is observed that the GMPEs

125 available for the Nepal region are unable to capture the exact trend in the data for the M_w 7.9
126 earthquake, a ground motion prediction model is also developed for all the GMPs and the
127 attenuation relations are validated through residual analysis and through comparison with the
128 local as well as global GMPEs that are applicable to the region. In addition, the hierarchy of
129 the current model against the entire local as well as the global models is validated through a
130 ranking system.

131

132

133 **2. STRONG MOTION DATABASE**

134

135 **2.1. Strong Motion Data**

136 Despite being seismically active, a nationwide seismic monitoring network is not available in
137 Nepal. However, four datasets of records are obtained here from four different sources: The
138 United States Geological survey (USGS), Hokkaido University, Japan and Tribhuvan
139 University, Nepal (HU-TU), Central Indo-Gangetic Plains Network (CIGN) and the Indian
140 Institute of Technology, Roorkee (IITR). As discussed in the literature, the first set of ground
141 motion records were provided by USGS at the KATNP (Kantipath, Kathmandu) station
142 (USGS, 2015b). Along with the main event, this station has recorded accelerograms of seven
143 aftershocks. Later, the collaborated effort of Hokkaido University, Japan and Tribhuvan
144 University, Nepal (HU-TU) provided four strong motion records for the 2015 main shock at
145 stations KTP, PTN, THM and TVU. Both these sets of acceleration time histories were made
146 available through CESMD (Center for Engineering Strong Motion Data) (Takai *et al.* 2016).
147 The relatively new CIGN network comprise of 26 strong motion stations, 19 of which are
148 located in the IGP (Chadha *et al.*, 2016); whereas the IITR seismic network of around 100
149 accelerographs in the IGP (Kumar *et al.* 2012). However, due to technical issues, the main
150 shock of the 2015 earthquake was recorded by just 15 CIGN stations and 13 IITR stations.
151 The number of recording stations is even lower for the succeeding aftershocks. Based on the
152 records available, ground motions of the M_w 7.9 2015 Nepal earthquake and its
153 accompanying five major aftershocks are considered in the present study. Table 1 gives the
154 details of these aftershocks, which range between M_w 5.3 - M_w 7.2 and were dated just within
155 a month of the main event. Fig. 1 shows the epicentres of all the earthquakes considered in
156 the current study, along with the distribution of the ground motion recording stations.

157

158 **2.2. Seismo-tectonics and site geology of the region**

159 The Himalayan front is divided into four major tectonic divisions from south to north, which
160 were demarcated by the presence of physiographic transition zones in the form of prominent
161 faults (Gansser, 1964; Validya, 1964). The four tectonic divisions include – Sub Himalaya,
162 Lesser Himalaya, High Himalaya and Tethyan Himalaya. The southernmost division i.e., the
163 Sub Himalaya is separated from the IG plain by the Himalayan Frontal Thrust (HFT), which
164 is sometimes also referred to as the Main Frontal Thrust (MFT). Other prominent faults from
165 south to north include Main Boundary Thrust (MBT), Main Central Thrust (MCT), and the
166 South Tibetan Detachment System (STDS). Nepal is diversified with all the four tectonic
167 divisions, wherein the Kathmandu valley is located in the Lesser Himalayan belt (Sakai *et al.*,
168 2002). It can be observed from the seismo-tectonic map given in Fig. 1 that the earthquake

169 rupture originated in the MCT and the effect of these tremors was felt even in the IGP
170 whereas the impact of these tremors was severe in the Kathmandu valley. It is to be noted
171 here that each of the tectonic divisions is characterised with similar physiography and the
172 thickness of sedimentary deposits usually decreases from north to south. Therefore, sediment
173 depth is considered as an important attribute to the ground motions considered in this study
174 and thus, the sediment depth of each ground motion recording station is obtained from 1D
175 velocity models given by Bijukchhen *et al.* (2017) and Dhakal *et al.* (2016). In addition, the
176 average velocity in the top 30 metres of the soil strata (V_{s30}) is regarded as the representative
177 of site characteristics. The V_{s30} values for the IGP sites are obtained from the SPT-N studies
178 of Anbazhagan and Bajaj (2020) and the V_{s30} values for the sites in the Kathmandu basin are
179 obtained from Bijukchhen *et al.* (2017) and USGS. Table 2 gives NEHRP Site classification
180 based on the V_{s30} , as given by IBC, 2009. The details of all the 40 ground motion recording
181 stations along with the respective sediment depth and NEHRP site class are given in Table 3.
182

183 **2.3. Preliminary Studies**

184 In the present study, 155 ground motions recorded at 40 stations due to a total of six
185 earthquakes are used for the analysis. Since it is established that the current study include
186 analysis of ground motions in the IGP as well, the range of distance measure is considered
187 from 10 to 1000 km. Fig. 2 shows the distribution of strong motion records used in the
188 present study with respect to moment magnitude (M_w) and rupture distance (R_{rup}). It can be
189 observed that the dataset consist of near field records with R_{rup} as low as 12 km and far field
190 records with R_{rup} as high as 896 km. Histograms showing the number of records with respect
191 to M_w , R_{rup} and hypocentral depth are given in Fig. 3.

192 The raw data recorded by CIGN and IITR stations with a sampling rate of 100 Hz is
193 subjected to a baseline correction of 0.1 to 25 Hz. On the other hand, HU-TU data with a
194 sampling rate of 200 Hz is corrected from 0.1 to 49.5 Hz. Overall, the highest PGA of the
195 dataset is observed to be 0.265g at the near-field station KTP. Interestingly, the observed
196 PGA values are unexpectedly lower for such an earthquake which is notorious for severe
197 structural damage. Fi. 4 shows the acceleration time histories (EW, NS and UD component)
198 of the main event for stations KTP, THM and AMT. It is observed that among the two near-
199 field stations KTP and THM, the PGA recorded is slightly higher for KTP due to lower R_{rup}
200 than that of THM (0.14g). In comparison to these stations, the station AMT which is located
201 in the IGP is observed to have significantly lower PGA value (0.017g). Spectral acceleration
202 plots of the respective ground motion time histories are obtained in Fig. 5, to observe any
203 underlying pattern with respect to the period. Predominant long period amplitudes are
204 observed clearly for Kathmandu basin stations – TVU, PTN and THM, wherein all of these
205 are located on sediment sites. On the other hand, low period predominant amplitude is
206 observed in the case of the station KTP, which is located on rock hard strata. Hence, sediment
207 depth is considered as an important factor to which ground motion amplifications of the
208 current dataset can be attributed. However, such long period amplifications at sediment sites
209 in the IG basin are not clearly identified from the spectral acceleration plot. Therefore,
210 detailed analysis of these ground motions are conducted via ground motion parameters.

211

212

213 3. GROUND MOTION PARAMETERS

214

215 Ground motion parameters are essential for describing unique characteristics of the ground
216 motions in a short, compact and quantitative form. Thus, parameters representing peak values
217 such as PGA, PGV and PGD are more prevalent among engineers for seismic design and
218 linear structural analysis. But, peak values solely are not sufficient to characterize damage
219 potential completely. Therefore, many other GMPs like Vertical-to-Horizontal (V/H) spectral
220 ratio, Root-Mean-Square acceleration (a_{RMS}), Arias Intensity (I_a), etc. were developed and
221 were correlated with structural damage. Bhargavi and Raghukanth (2018, 2019) have
222 presented a total of 21 such parameters and have rated the damage potential of a ground
223 motion using principal component analysis. The same 21 parameters are selected to analyze
224 the ground motions considered in the current study. Amplitude, frequency content and
225 duration are the three significant characteristics of a ground motion (Kramer, 1996). Thus,
226 based on the dependence of a parameter on the number of these essential ground motion
227 characteristics, GMPs are broadly categorized into three classes: First order, second order and
228 third order parameters. The basic definitions of all the ground motion parameters obtained
229 from the acceleration time histories are explained in table 4.

230

231 3.1. First Order Parameters

232 First order parameters which represent only one key characteristic include Peak Ground
233 Acceleration (PGA), Peak Ground Velocity (PGV), and Peak Ground Displacement (PGD).
234 Combinations of these parameters like V/H ratio of PGA and PGA/PGV ratio are also shown
235 as better characterisations of ground motions. Seismic damage is complicated and difficult to
236 interpret solely on the basis of peak values of the ground motion. Therefore parameters such
237 as PGA, PGV and PGD are insufficient to determine the damage potential of an earthquake.
238 Thus, frequency content based parameters are estimated through analysis of the time history
239 in the frequency domain using Fourier Amplitude Spectra (FAS) and Power Spectra.
240 Venmarke (1976) proposed central frequency and shape factor as the indicators of statistical
241 properties of Power Spectral Density (PSD) function, using spectral moments. Predominant
242 frequency is measured as the frequency corresponding to maximum value of the Fourier
243 amplitude spectrum. Similarly, a duration parameter – Significant duration, which depends on
244 the ground motion time history, is considered. Significant duration is defined as the time
245 interval between 5% and 95% energy thresholds of the of Husid plot (Husid, 1969).

246 Extremely low values of the entire peak GMPs including PGV and PGD, are
247 recorded for all the IG basin records, compared to that of the Kathmandu basin records. The
248 predominant frequencies of records at near field sediment sites are of the order 0.2, whereas
249 the predominant frequency obtained at the station KTP is of the order 3.4. Once again, these
250 values clearly indicate the presence of predominant low frequency i.e., long period
251 amplitudes in the ground motions recorded at sediment sites of the Kathmandu basin. On the
252 other hand, even though predominant frequencies as high as 3 and 7.5 are recorded at some
253 sites in the IG basin, no clear pattern is observed with the variation of these values with the
254 sediment depth. Interestingly, the central frequencies (Ω) of vertical component of ground
255 motion records are greater (double in most cases) to that of the horizontal component, for

256 both the Kathmandu and IG basins. However, the $V/H_{(PGA)}$ values are less than 1 for most of
257 these cases. Further, the parameter 'q', which represents bandwidth of the time history, is
258 almost same in both the components (Horizontal and Vertical), if not slightly higher for
259 horizontal component of the ground motions. Therefore, based on the definition of 'Ω' and
260 'q', it can be assumed that the horizontal component of these ground motions have registered
261 almost the same amount of energy at lower frequencies to that of the vertical component
262 concentrated at high frequencies. As expected, significant durations of near field records of
263 the Kathmandu basin are extremely low to that of the far field IG basin records. All the
264 primary parameters for two near field stations – 'KTP' and 'THM' in the Kathmandu basin
265 and one far field station – 'AMT' in the IGP basin are tabulated in table 5.

266

267 **3.2. Second Order Parameters**

268 First order parameters are not sufficient for complete characterisation of the ground motion.
269 Housner and Jennings(1964) proposed Root-Mean-Square acceleration (a_{RMS}) as an indicator
270 of damage as it considers the acceleration time history over significant duration of Husid
271 plot. Arias intensity was introduced to observe the cumulative effects of ground motion over
272 the complete Duration. Benioff (1934) proposed Response Spectra as a quantification of the
273 physical response of structures due to ground shaking. Housner (1952) derived spectrum
274 intensities from the pseudo velocity response spectrum to measure the intensity of ground
275 shaking between periods of 0.1s to 2.5s. Park *et al.* (1985) introduced a new secondary
276 parameter as the characteristic intensity in terms of a_{RMS} and significant duration, which is
277 linearly related to damage index due to maximum deformation and absorbed hysteretic
278 energy (Ang *et al.*, 1990). Acceleration Spectral Intensity (ASI) was developed by VonThun
279 *et al.* (1988) to characterize ground motion at higher frequencies. Electrical Power Research
280 Institute (EPRI, 1988) mathematically defined a ground motion intensity measure named
281 cumulative absolute velocity (CAV) as an efficient indicator of structural damage. Reed and
282 Kassawara (1990) proposed CAV for determining the exceedance of the Operating Basis
283 Earthquake (OBE) after the occurrence of an earthquake at a nuclear power plant.

284 In the preliminary analysis it was noted that maximum PGA of 0.265g was observed
285 at the station KTP, despite KTP being on a rock site. It should be noted that all the four
286 stations of the HU-TU network including the KTP station are located at almost similar
287 rupture distances. Therefore, this preliminary result would be considered an inconsistency
288 from the common notion that soft soil sites or sediment sites usually register larger ground
289 motion amplitudes. Upon the analysis of the second order parameters, it was revealed that the
290 sediment sites indeed recorded larger amplitudes. The highest a_{RMS} value of the dataset –
291 0.055g is obtained for the station TVU, whereas the a_{RMS} value for the highest recorded PGA
292 component of KTP is 0.036g. In fact, all the other second order amplitude parameters such as
293 I_a , CAV and I_c are higher for the record of station TVU over the record at KTP. This is due to
294 the reason that the significant duration of the former ground motion is greater than the latter.
295 Results like these emphasize the importance of second order parameters over first order
296 parameters such as PGA, PGV and PGD. Table 5 shows the secondary parameters i.e., a_{RMS} ,
297 I_a , CAV, I_c , ASI, VSI etc. for stations KTP, THM and AMT.

298

299 **3.3. Third Order Parameters**

300 An acceleration time history is highly stochastic and non-stationary. Thus, the statistical
301 properties vary in both temporal and spectral domain simultaneously. Therefore, parameters
302 reflecting non-stationarity of ground motions became popular. The concept of evolutionary
303 power spectral density was introduced by Priestley (1965), which gives an account of the
304 temporal and frequency evolution of time history. To address non-stationarity in ground
305 motion, Liu (1970b) used instantaneous power spectrum models. However, PSD does not
306 exist for non-stationary processes. Huang *et al.* (1998) proposed that nonlinear and non-
307 stationary data can be handled through Empirical Mode Decomposition (EMD) with Hilbert-
308 Huang transformation (HHT). Wu and Huang (2009) modified the EMD as an Ensemble
309 EMD, which includes the shifting of an ensemble of white noise modulated signals to remove
310 signal intermittency. Raghukanth and Sangeetha (2012) have presented the detailed
311 procedure for obtaining EPSP for ground motions using EMD-HHT method. In the present
312 study, Ensemble Empirical Mode Decomposition (EEMD) is used to obtain EPSP where the
313 acceleration time history is decomposed into simple empirical oscillatory modes called
314 Intrinsic Mode Functions (IMFs). An IMF is defined as a time series where the number of
315 extrema and the number of zero crossings must either be equal or differ by one. Further, the
316 mean value of the envelopes defined by the local maxima and minima of the IMF is zero.
317 Thus obtained IMFs are orthogonal to each other. Initially, a white noise of negligible
318 amplitude is added to the acceleration time history. From the resulting signal, the consecutive
319 maxima and minima are identified. These maxima and minima are connected as cubic splines
320 to obtain the upper and lower envelopes. Thus IMFs are obtained by removing local median
321 in every step through a shifting procedure. This process is continued until a monotonic
322 function is obtained at the end. Once the IMFs are obtained, the Hilbert Transform of these
323 IMFs is used to construct EPSP. An Analytic Function (AF_j) is formed from the conjugate
324 pair of obtained IMF (IMF_j) and Hilbert Transform of the respective IMF ($hIMF_j$) to
325 calculate the corresponding amplitude (C_j) and instantaneous frequency (ω_j) of EPSP. The
326 amplitude of energy in time is combined over different frequencies for the extracted IMFs to
327 obtain the energy distribution in the time-frequency domain in the form of a surface plot of
328 EPSP ($G(\omega, t)$).

329

$$AF_j = IMF_j(t) + ihIMF_j(t) = C_j(t)e^{-i\omega(t)} \quad (1)$$

$$C_j(t) = \sqrt{[IMF_j^2(t) + hIMF_j^2(t)]} \quad (2)$$

$$\omega_j(t) = \frac{d}{dt} \left[\tan^{-1} \left(\frac{hIMF_j(t)}{IMF_j(t)} \right) \right] \quad (3)$$

$$G(\omega, t) = \sum_{i=1}^n \frac{1}{2} [C_j(IMF_i)]^2 \quad (4)$$

330

331 The EPSP plots thus obtained for the near field stations in Kathmandu valley (KTP, PTN,
332 THM and TVU) as well as the far field IGP stations (VNS, AMT, BSR and TDR) and their
333 respective time histories are shown in Figure 6. The energy distribution varying with time
334 and frequency can be visualized here and it can be observed that energy distribution in these
335 evolutionary spectra is concentrated over the frequency of 6-8 Hz for the near field stations.

336 The EPSD function is considered as the third order GMP and its statistical properties are
337 characterised using: Total Energy (E_{acc}) of the EPSD, first and second moments of the
338 frequency and time axes: Spectral Centroid (E_w), Spectral standard deviation (S_w) and
339 Temporal Centroid (E_t), Temporal standard deviation (S_t) and, the Correlation coefficient of
340 the time and frequency. All the 21 GMPs derived for the three stations – KTP, THM and
341 AMT are shown in Table 5.

342

343 **3.3.1 Analysis using IMFs and instantaneous frequencies**

344 The IMFs of the acceleration time history (EW) for four stations in the IG basin: VNS, AMT,
345 BSR and TDR and four stations in the Kathmandu basin: KTP, TVU, PTN and THM, along
346 with their respective percentage variances are shown in Figure 7. The original signal can be
347 expressed as the sum of these IMFs. The percentage of variance of each IMF is defined as the
348 contribution of each IMF to the total variance of the data. In majority of cases in the present
349 study, the original data can be reconstructed using the first 10 IMFs. It can be observed that
350 for VNS station, the first four IMFs are sufficient to recreate 93% of the ground motion
351 where the first IMF represents 52% of the original data. For the station AMT, first five IMFs
352 are required to regenerate 94% of the original time history with both first and second IMF
353 having percentage variance of 25%. For station BSR, 95% of the ground motion can be
354 obtained by adding up the first six IMFs. In this case, the second IMF is more dominant than
355 the first IMF, with 24% of the total variance. For station TDR, IMF4 is predominant with
356 19% of the variance of ground motion, and the first six IMFs are required to reproduce 93%
357 of the ground motion. Thus, it can be assumed that as the thickness of sediments in the IG
358 basin increases from south to north, the IMFs with higher modes becomes predominant.

359

360 In case of the station KTP, which is located at a hard rock site of the Kathmandu basin, first
361 five IMFs are essential for rebuilding the original ground motion. Here, the maximum
362 percentage variance is observed for the first IMF, representing 38% of the original data. For
363 stations PTN, THM, TVU and KATNP located at sediment sites, higher IMFs (third to sixth
364 IMF) are more dominant. For instance, IMF3 constitute 30% of the variance for station TVU.
365 On the other hand, IMF4 represents 49% of the ground motion for station THM and 30% of
366 the ground motion for station PTN. Moreover, IMF6 constitute 54% of the variance for
367 station KATNP. Therefore, it is inferred that IMFs with higher modes are indeed
368 predominant for ground motions recorded on sediment sites of the Kathmandu valley as well.
369 The estimated instantaneous frequencies from IMFs of acceleration time histories for both IG
370 basin and Kathmandu basin are shown in Figure 8. Table 6 shows the mean and standard
371 deviation of the instantaneous frequency, along with the percentage variances of the first 10
372 IMFs for all the ground motion records of the Mw 7.9 earthquake. According to the definition
373 of IMFs, the mean instantaneous frequency of the IMFs decreases in the higher modes.
374 Moreover, the instantaneous frequency corresponding to the dominant IMF i.e., the IMF with
375 maximum percentage of variance, is considered as the dominant frequency of the acceleration
376 time history. Consequently, it is observed that the dominant frequencies of the IG basin
377 stations -VNS, AMT, BSR and TDR are 7.2 Hz, 9.6 Hz, 5 Hz and 1.3 Hz respectively.
378 Therefore, the dominant mean instantaneous frequencies are decreasing as the R_{rup} increases
379 and the thickness of the sediment increases. In order to observe this property, the dominant

380 IMFs and their corresponding mean instantaneous frequencies are plotted against sediment
381 depth in Figure 9. Even though there is no underlying pattern in the dominant IMFs of the
382 data at intermediate rupture distances, it can be observed that the higher modes are prevalent
383 for larger sediment depths at larger distances as well as at near field distances. Further, a
384 decent trend can be observed in Figure 9(b), wherein the dominant mean instantaneous
385 frequencies are decreasing with the sediment depth, especially for the CIGN records.
386 Srinagesh *et al.* (2011) conducted analysis of the IG basin and provided the fundamental
387 frequencies of the basin, wherein these fundamental frequencies increases (from 0.13 to 0.8)
388 with a decrease in the sediment thickness (from north to south of the basin). Consequently,
389 significant amplifications due to sediment depth are expected at the site's fundamental
390 frequency. Therefore, it is assumed that significant amplifications due to the sediment depth
391 are not observed in many of the far field ground motions that were recorded in the IG basin.
392 However, in order to obtain amplifications at each of these sites, a HHT based procedure is
393 adopted.

394

395 **3.3.2 HHT based site amplification**

396 In this section, site amplification analysis is carried out to provide a greater understanding of
397 the effect of sediment depth on ground motions over short and long time periods. Several
398 techniques have been in use to perform the site response analysis, of which the most common
399 procedure involves comparing the Horizontal-to-Vertical Spectral Ratio (HVSr). However,
400 this method is infamous for not estimating the exact amplification value and therefore limits
401 its application. Various other methods based on Fourier spectral analysis such as the Standard
402 spectral ratio (SSR) technique and the Ratio of source spectrum (RSS) technique were
403 suitable for the present scenario and they have been successfully used previously to estimate
404 the site amplification in the IGP and Kathmandu basins (Chadha *et al.*, 2016; Singh *et al.*,
405 2020; Rajaure *et al.*, 2016). However, the studies considering Fourier-based amplification
406 assumes linear behavior of the sites sediments. In Fourier-based estimation, frequency is
407 constant for the entire length of data window where as HHT-based frequency content analysis
408 considers temporal non-stationarity as well. Therefore, the effect of site nonlinearity on the
409 ground motion time histories can be characterized by estimating site amplification based on
410 the marginal spectrum obtained from the HHT (Zhang, 2006). It is defined as the ratio of
411 marginal Hilbert amplitude spectra of sedimentary sites to hard rock sites, which is similar to
412 the SSR technique involving Fourier based analysis. The marginal Hilbert amplitude spectra
413 $A(\omega)$ over a temporal duration T of the ground motion and the HHT based amplification
414 factor ($FH_s(\omega)$) are given in equations 5 and 6 respectively.

415

$$A(\omega) = \int_0^T G(\omega, t) dt \quad (5)$$

$$FH_s(\omega) = \frac{\sqrt{A_{s,EW}^2 + A_{s,NS}^2}}{\sqrt{A_{r,EW}^2 + A_{r,NS}^2}} \quad (6)$$

416

417 Here, $G(\omega, t)$ is the EPSD; the subscripts EW and NS represent two horizontal components
418 and subscripts s and r represent the target sedimentary site and reference hard rock site
419 respectively. The HHT-based amplifications thus obtained at sedimentary sites of IGP and
420 Kathmandu basin for M_w 7.9 mainshock are shown in figure 10. Site amplification factors are
421 estimated for VNS, AMT, BSR and TDR stations in the IG basin considering hard reference
422 sites as ALB, ALM, ALM and RPG respectively. Similarly, the KTP station is considered as
423 a reference hard rock site for TVU, PTN, THM and KATNP in the Kathmandu basin. In
424 order to compare these values with the Fourier based amplifications, the corresponding plots
425 are obtained for the IGP and KB sites using the procedure given Chadha *et al.* (2016) and
426 Rajaure *et al.* (2016) respectively. It can be observed from the figure that HHT-based
427 dominant frequency is similar to the fundamental frequency obtained using Fourier based
428 method for most of the sites in IGP. A slight diminution can be observed at the HHT- based
429 dominant frequency at some stations in Kathmandu basin. However, a significant increase in
430 the amplification values can be observed at these frequencies for HHT- based method than
431 the Fourier based method.

432

433 The amplifications obtained for the four far field stations VNS, AMT, BSR and TDR of the
434 IG basin for the M_w 7.9 mainshock and the two aftershocks of M_w 6.7 and M_w 5.3 are shown
435 in figure 11. It can be observed that the HHT- based dominant frequency is same for M_w 7.9
436 mainshock and M_w 5.3 aftershock for IGP stations. The amplitude for the fifth aftershock
437 (M_w 5.3) is much lesser than the M_w 7.9 mainshock and its immediate aftershock (M_w 6.7) for
438 most of the stations in IGP. This is a fair observation since aftershocks are usually created
439 along a ruptured fault, and hence have lower amplitude than mainshock. The increased site
440 amplification at lower frequencies for the higher magnitude events can be considered as an
441 indication of soil nonlinearity at these sites. Moreover, it should be noted that the sediment
442 depth of the station VNS is much less compared to the other three sites. Therefore, the
443 amplifications noticed at this site are not quite high as that of the other sites. At the same
444 time, the amplifications for aftershocks noticed at this site are higher than mainshock. Thus,
445 comparing with the other sites the station VNS shows weak or no site nonlinearity under the
446 mainshock. However, the dataset used in the present study is considerably small and sparse to
447 obtain a relation between sediment depth and site amplification. In fact, the amplification
448 factors cannot be found for many sites in the IGP due to lack of information regarding a
449 sufficiently close hard rock site. Moreover, the aftershock records at the near-field stations
450 are not available for the Kathmandu basin. Therefore, given a comprehensive dataset, the
451 nonlinear behavior and the effect of sediment depth on amplification of the site can be
452 studied more comprehensively for insightful observations. Overall, the HHT based method
453 proved to be a better option for obtaining site amplification.

454

455

456 **4. GROUND MOTION MODEL**

457

458 In this section, a region specific ground motion prediction equation (GMPE) is developed. A
459 suitable GMPE is most important for predicting the future seismic potential of the region.

460 Due to lack of a comprehensive dataset, developing a regional GMPE for Himalayan orogeny
 461 is quite challenging. Despite of this, Singh *et al.* (2017) have developed empirical equations
 462 for the Himalayan region for three amplitude parameters: PGA, PGV and PGD. However, it
 463 will be interesting to develop new prediction equations for the traditional parameters as well
 464 as for those important for engineering applications like a_{RMS} , I_a , CAV etc. Due to small and
 465 sparse dataset, nonlinear mixed effect model (Abrahamson and Youngs, 1992) is used to
 466 develop GMPEs in the present study to incorporate both fixed and random effects. The
 467 algorithm for estimation of model parameters and variances involve estimating the
 468 parameters first using the fixed effects regression method. The most general framework of
 469 GMPE model is

$$\ln y_{ij} = f(M_i, R_{ij}, \theta) + \eta_i + \epsilon_{ij} \quad (7)$$

471
 472 Where Y_{ij} is a ground motion parameter, $\mu_{ij} = f(M_i, R_{ij}, \theta)$ is the attenuation equation, M is
 473 the measure of earthquake source, R is distance measure, θ is the vector of model parameters,
 474 η_i represents inter-event residual for event i and ϵ_{ij} represents intra-event residual for jth
 475 recording of event i.

476

477 **4.1. Ground motion prediction equation**

478 The predictor variables considered for the GMPE include M_w , R_{rup} , flags for Vs30 of the site
 479 (F_{Vs30}) and sediment depth variation (F_{sed}). In order to formulate attenuation relationships
 480 between a GMP and the predictor variables, the following simple functional form is selected
 481 for all the 21 GMPS:

482

$$\ln(GMP) = \ln(gmp) + \eta_i + \epsilon_{ij} \quad (8)$$

$$\ln(gmp) = f_{mag} + f_{dis} + f_{site} + f_{sed} \quad (9)$$

$$f_{mag} = c_0 + c_1 M \quad (10)$$

$$f_{dis} = c_2 R + c_3 \ln(R + c_4 \exp(c_5 M)) \quad (11)$$

$$f_{site} = c_6 F_{Vs30} \quad (12)$$

$$f_{sed} = c_7 F_{sed} \quad (13)$$

$$\tau^2 = var(\eta_i) \text{ and } \sigma^2 = var(\epsilon_{ij}) \quad (14)$$

483

484 Here, f_{mag} represents the magnitude term where coefficient c_0 corresponds to magnitude or
 485 absolute scale of the GMP (equation 8). The term ' $c_1 M$ ' represents effect on the ground
 486 motion concerning earthquake magnitude considered as source measure. The distance
 487 measure f_{dis} represents the term ' $c_2 R + c_3 \ln(R + c_4 \exp(c_5 M))$ ' corresponds to geometric
 488 and anelastic attenuation (equation 11). Since all the high amplitude records of the dataset are
 489 concentrated at distances less than 20 km from the source, near field magnitude saturation
 490 due to the M_w 7.9 mainshock is expected. In order to capture this property, the term

491 ' $c_4 \exp(c_5 M)$ ' is incorporated in the model; wherein these non-linear coefficients are
492 obtained using non-linear least square regression analysis. The f_{site} represents the term
493 ' $c_6 F_{Vs30}$ ' incorporating the local site conditions where ' F_{Vs30} ' includes flags for the site
494 classes as per Vs30 (equation 12). Further, it is already mentioned in the earlier sections that
495 the sediment depth varies from 0.6-5 km in the IG basin, whereas the sediment thickness
496 ranges between 0.04-0.46 km in the Kathmandu basin. Based on sediment thickness
497 variation, the sites are classified into three different classes. The sites for sediment depth <1
498 is marked with $F_{sed}=0$; for sediment depth from 1 km to 2.5 km with $F_{sed}=1$; for sediment
499 depth greater than 2.5 with $F_{sed}=3$. Therefore the term ' f_{sed} ' incorporates site spectral
500 amplification factor via the site's sediment depth (equation 13). The event-to-event (inter-
501 event, η_i) and within event (intra-event, ϵ_{ij}) residuals are independent and identically
502 distributed and are jointly normal (equation 14). The coefficients c_k ($k = 0, 1, \dots, 7$) in the
503 equation 8 and the corresponding inter-event and intra-event standard deviations are shown in
504 table 7. The estimates thus obtained are plotted against recorded data for all the 21 GMPs in
505 Fig. 12. It can be observed that the functional form considered in the study is able to capture
506 the trend in the behavior of GMPs with source and distance measures. In order to evaluate the
507 validity of the GMPE developed in this study, the inter-event and intra-event residuals
508 obtained from the mixed effect regression analysis are plotted against the predictor variables
509 in the model. The variation of the residuals plots are shown in Figure A7 (Electronic
510 supplement). In these residual plots, a positive residual indicates the underestimation of the
511 actual value by the predicted model and a negative residual indicates the overestimation of
512 the actual value by the predicted model. It is observed that neither inter-event nor intra-event
513 residuals shows trend with respect to the predictor variables.

514

515

516 **5. COMPARISON WITH ESTABLISHED GMPES**

517

518 In order to study the goodness of the model, the GMPEs developed here are compared with
519 previously established GMPEs that were applicable to the region considered in this study
520 (Douglas, 2020). Subsequently, a comparison of these already established models with the
521 present model is presented for the two events of magnitudes $M_w 7.2$ and $M_w 7.9$.

522

523 **5.1. Peak Ground Acceleration (PGA)**

524 Regionally established regression model proposed by Singh *et al.* (2017) – SSSA17 and well
525 established global GMPEs proposed by ASK14 – Abrahamson *et al.* (2014), Boore *et al.*
526 (2014) – BSSA14, Campbell and Bozorgnia (2014) – CB14, Chiou and Youngs (2014) –
527 CY14, Kanno *et al.* (2006) – K06 and Zhao *et al.* (2006) – Z06, are selected for comparison
528 of the GMPE derived for PGA. All the global models are derived for different tectonic
529 environments. ASK14 developed the ground motion model for mean horizontal component of
530 PGA for shallow crustal earthquakes in active tectonic regions, where the regression
531 equations were developed using the NGA-West2 database. BSSA14 includes the GMPE for
532 mean horizontal component of PGA, PGV and PSA (5 % damping ratio) for the same
533 database. Similarly, CB14 was also developed for the NGA-West2 database, which includes

534 period-dependent magnitude saturation, faulting effects, magnitude-dependent scaling with
535 hypocentral depth and fault dip, regionally dependent anelastic attenuation, updated hanging
536 wall effects and so on. K06 includes two regression models for shallow and deep events with
537 a V_{s30} of 300m/s. Z06 includes attenuation models considering tectonic source types and
538 focal mechanisms. SSSA17 was proposed for IGP based on an approximate solution of a
539 circular finite source model using the same database. The GMPE is derived using exponential
540 integral function where M_w and R_{rup} are considered as the predictor variables. Fig. 13 shows
541 the comparison of different GMPEs with the PGA values estimated using a flag value of 2 for
542 both F_{Vs30} and F_{sed} and for two moment magnitudes: $M_w 7.2$ and $M_w 7.9$. It is observed from
543 the figure that the global GMPEs are under-predicting the recorded data. It can be observed
544 that the developed model is able to predict the data. Remarkably, the regional prediction
545 model proposed by SSSA17 is able to capture the trend of the data. However, it can be
546 observed from table 7 that the inter-event and intra-event standard deviation of the present
547 model is lesser than that of SSSA17. Moreover, the present model also incorporates site
548 classes and sediment depth variations along with the M_w and R_{rup} in the functional form.
549 Thus, the GMPE derived in the present study provides better estimates compared to other
550 models.

551

552 **5.2. Peak Ground Velocity (PGV)**

553 Similarly, regression models proposed by Singh *et al.* (2017) – SSSA17, ASK14 –
554 Abrahamson *et al.* (2014), Boore *et al.* (2014) – BSSA14, Campbell and Bozorgnia (2014) –
555 CB14, Chiou and Youngs (2014) – CY14, and Kanno *et al.* (2006) – K06 are compared with
556 the GMPE derived for PGV. Consequently, corrections are applied to the model K06. Fig. 14
557 shows the comparison of these GMPEs, with the PGV estimated using a flag value of 2 for
558 both F_{Vs30} and F_{sed} and for two moment magnitudes: $M_w 7.2$ and $M_w 7.9$. The models
559 BSSA14, CB14, and CY14 are under-predicting the records whereas estimates of SSSA17
560 and K06 are close to the present model and provides a sufficiently close estimation of PGV.

561

562 **5.3. Peak Ground Displacement (PGD)**

563 The models given by Singh *et al.* (2017) – SSSA17 and Campbell and Bozorgnia (2008) –
564 CB08 are selected for the comparison of the GMPE derived for PGD. The model CB08 was
565 derived for crustal events, and includes hypocentral depth, shallow linear and nonlinear site
566 responses, 3-D basin response and hanging wall effects. Due to lack of information, the
567 present model does not consider fault mechanisms and basin responses. Fig. 15 shows the
568 comparison plot of the GMPEs using a flag value of 2 for both F_{Vs30} and F_{sed} and for
569 magnitudes $M_w 7.2$ and $M_w 7.9$. The model CB08 is over-predicting the estimated values of
570 PGD, whereas the model SSSA17 gives sufficiently closer estimates to the recorded data.

571

572 **5.4. Arias Intensity (I_a)**

573 There are very few GMPEs available for complex parameters like arias intensity, root-mean-
574 square acceleration and CAV. The GMPEs derived by Sharma and Srbulov (1998) – SS98,
575 Foulser-Piggot and Goda (2014) – FPG14 and Amiri *et al.* (2009) – A09 are selected for the
576 comparison of arias intensity (I_a). The model SS98 considered duration and number of pulses
577 as a function of acceleration, in addition to the usual predictor variables like magnitude and

578 source-to-site distance. Further, surface wave magnitude (M_s) was used as source measure,
579 and equations were derived for rock and soil site conditions. The model A09 also used
580 surface wave magnitude to derive four different equations for different ground types and
581 tectonic conditions for Iran. The comparison of these GMPEs with the present model for I_a
582 using a flag value of 2 for both F_{Vs30} and F_{sed} is shown in Fig. 16. It is observed that the
583 model A09 is not following the trend in the behaviour with predictor variables. However, the
584 model SS98 is over-predicting the data for M_w 7.2 while able to follow the trend for M_w 7.9.
585 FPG14 is under-predicting the arias intensity for far-field records and overestimating the
586 same for near field records. As expected, the regression equation derived in the present study
587 provides better estimates compared to other models.

588

589 **5.5. Cumulative Absolute Velocity (CAV)**

590 In this section, models proposed by Campbell and Bozorgnia (2010) – CB10, Foulser-Piggot
591 and Goda (2014) – FPG14 and Du and Wang (2013) – DW13 are used to compare the
592 estimates of CAV. The models CB10 and DW13 include intercontinental ground motions to
593 derive prediction equations for shallow crustal events. These models incorporate magnitude,
594 source-to-site distance, site conditions, faulting style hanging wall effect, linear and nonlinear
595 site response. However, both the models are not following the trend in the behaviour of
596 parameters and gives under-predicting estimates for CAV values. However, the model
597 FPG14 provides sufficiently close estimates of CAV for far-field data while overestimating
598 the near field data. Fig. 17 shows that the present model is relatively good in predicting the
599 estimates.

600

601

602 **6. RANKING OF GMPEs**

603 Since the model derived in the current study used a limited dataset of ground motion records,
604 the GMPEs cannot be used directly in any hazard studies. With the advent of statistical
605 approaches such as the likelihood methods, the goodness of GMPEs can be quantified such
606 that the most suitable model for a particular region can be selected with more ease. Therefore,
607 a statistical analysis is performed to facilitate the ranking of the established GMPEs
608 hierarchically for hazard and risk assessment. The most popular methods for this task are the
609 likelihood (LH) method (Scherbaum *et al.*, 2004), the log-likelihood (LLH) method
610 (Scherbaum *et al.*, 2009) and the Euclidean distance (EDR) method (Kale and Akkar 2013).
611 The basic concept in the LH method involves normalization of residuals obtained between
612 the recorded data and estimates, with standard deviation of the GMPE. The LLH method
613 measures the average log likelihood of the estimates given by the GMPE and the observed
614 data, both of which are considered as two continuous log normal probability density
615 functions. On the other hand, the EDR method makes use of the Euclidean distance concept
616 to obtain ranking of various GMPEs. All the above three methods may sometimes give
617 different rankings for a set of GMPEs, as the LH method is sensitive to standard deviations.
618 Therefore, the concept of normalized ranking index (RI_N) is opted here, which is obtained
619 from using all the above three methods to derive an unbiased ranking system for the GMPEs
620 (Kale, 2009). Table 8 gives the ranks obtained for all the GMPEs considered for comparison

621 of PGA in this study, using all the three statistical methods. It can be observed that the all the
622 GMPEs are placed in a different ranking order for each of the methods. Similar observations
623 are noticed in the cases of PGV, PGD, I_a and CAV, which are given in tables 9-12
624 respectively. Therefore the ranking indices are normalized for each method, by giving the
625 best performing GMPE a value of 1.0. Further, the final normalized ranking index (RI_N) is
626 obtained by taking a weighted combination (EDR-0.50, LH-0.25 and LLH-0.25) of the
627 corresponding normalized rankings (EDR_N , LH_N and LLH_N).

628 The final ranks thus obtained for the GMPEs of PGA, PGV, PGD, I_a and CAV are given in
629 tables 13 and 14 respectively. From the tables, it can be observed that the model derived in
630 the present study is ranked at number one for all the GMPs, the reason being the calibration
631 of the model exclusively to the recorded data used in this study. On that note, it can also be
632 observed that the other locally derived model SSSA17 is ranked at number 2 for PGA and
633 PGD. All the other global models register much higher RI_N values for PGA, verifying that the
634 Nepal earthquake is a truly unique event of the Himalayan tectonic zone, which cannot be
635 estimated effectively by many well established global models. Interestingly, the RI_N values
636 are much less for PGV and PGD, especially the K06 model that is derived for Japan, which
637 signifies that either some of the global models are effective in predicting these parameters or
638 that the current model still needs improvisation. The same pattern can be obtained from the
639 plots given in figures 13, 14 and 15. The rankings obtained for I_a and CAV also follows the
640 same pattern as observed from the plots of figures 16 and 17.

641

642 7. CONCLUSION

643

644 The 2015 M_w 7.9 Nepal earthquake is not only the most catastrophic but also the first well-
645 recorded event to have occurred in the Himalayan orogenic belt. This paper presents a
646 detailed analysis of this event using 21 GMPs, which represents the amplitude, frequency
647 content, duration and evolutionary characteristics of the ground motion. Preliminary analysis
648 indicates that very low PGA values are recorded wherein the near field station KTP, which is
649 located on a rock site with an R_{rup} of 12.05 km has recorded PGA of 0.265g. However, the
650 a_{RMS} value for this ground motion record is lower to that of others that were recorded on
651 sediment sites located at similar rupture distances. This behaviour is also seen in other second
652 order parameters of the Kathmandu basin. Moreover, analysis of spectral accelerations
653 emphasizes the presence of long period spectral amplitudes in near field ground motions of
654 Kathmandu basin, only at sediment sites. However, the same pattern is not observed in the far
655 field IG basin records of the earthquake. Therefore, a thorough study of the ground motion
656 characteristics through first, second and third order GMPs is conducted.

657

658 The behaviour of the parameters concerning the thickness of sediment is studied through
659 IMFs using HHT-EMD for all the ground motions of the M_w 7.9 earthquake. Interestingly,
660 the IMFs with higher modes become predominant as the thickness of sediments in the IG
661 basin increases from south to north, thereby resulting in low dominant mean instantaneous
662 frequencies for sites having larger sediment depths. Moreover, for the station KTP located at

663 a hard rock site of Kathmandu basin, maximum percentage variance is observed for the first
664 IMF; whereas higher order (third to sixth) IMFs are more dominant for sediment sites (PTN,
665 THM, TVU and KATNP). Therefore, a similar pattern is observed for both the Kathmandu
666 and IG basins with respect to dominant mean frequencies and sediment depths. Lack of any
667 noticeable predominant long period spectral amplitude for most of the records in the IG basin
668 is attributed to the overall low amplitudes resulting from the source mechanism, combined
669 with large rupture distances. The effect of sediment depth on ground motions over short and
670 long time periods is emphasised more through the HHT-based site amplification study. A
671 significant increase is observed in the amplification values at the dominant frequencies for
672 HHT- based method than the Fourier based method. Comparing the M_w 7.9 mainshock with
673 two of its aftershocks, the increased site amplification at lower frequencies for the higher
674 magnitude events can be considered as an indication of soil nonlinearity at these sites.
675 Remarkably, the station VNS with much lesser sediment depth compared to the other sites
676 shows weak or no site nonlinearity under the mainshock. However, a denser dataset would
677 provide in a visibly comprehensible pattern.

678

679 Additionally, in this study, GMPEs are derived for the 21 estimated GMPs, using nonlinear
680 mixed effect regression analysis to incorporate both fixed and random effects. The predictor
681 variables used for deriving GMPEs are Moment Magnitude (M_w), Rupture distance (R_{rup}),
682 flags for V_{s30} of the site (F_{Vs30}) and sediment depth variation (F_{sed}). The estimates thus
683 obtained from the GMPE are compared with the recorded values of the M_w 7.9 mainshock
684 and the M_w 7.2 aftershock records. Further, to ensure that the developed GMPEs are not
685 biased, the behaviour of inter-event and intra-event residuals against the predictor variables is
686 verified. Furthermore, the estimated GMPEs are compared with well established prediction
687 models that can be applicable to the study region. Among the few models applicable to
688 parameters such as PGA, PGV, PGD, I_a and CAV, superior prediction is observed from the
689 estimates of the current model. Similar class of estimates can be expected for all the second
690 and third order GMPs of the current model, which is a novel endeavour for the Nepal
691 Himalaya region. In addition, the evaluation of ranking indices (RI_N) also highlights the
692 efficiency of the current model. Since all the earthquakes considered in this study are
693 originated due to a similar source mechanism, the GMPE model given here is not only region
694 specific, but also source specific. Analysis using a wide range of data would further ensure
695 the goodness of this model and its applicability. The GMPEs developed in this study may not
696 be used directly in applications such as PSHA studies due to the constricted dataset used in
697 deriving the models. However, when used with a logic tree approach, all the GMPEs will
698 contribute quite significantly. Further, the GMPEs of higher order parameters can be even
699 used to generate ground motion time-histories using stochastic methods such as the one given
700 by Sabetta and Pugliese (2021), which can be used for a nonlinear dynamic analysis in
701 earthquake-resistant design.

702

703

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871

872

873 **Table 1.** Earthquake ground motion database used in the current study

Sl. No:	Name of the Earthquake	Date (Day/mon/year)	Time (hr:min)	Epicenter		Depth (km)	M _w	No. of Records
				Lat (°E)	Lon (°N)			
1	Mainshock	25-04-2015	06:11	27.92	85.325	12	7.9	32
2	Aftershock 1	25-04-2015	06:45	28.224	84.822	21	6.7	23
3	Aftershock 2	26-04-2015	07:09	27.771	86.017	20.6	6.7	25
4	Aftershock 3	12-05-2015	07:05	27.809	86.066	12	7.2	25
5	Aftershock 4	12-05-2015	07:36	27.625	86.162	18.1	6.2	26
4	Aftershock 5	16-05-2015	11:34	27.559	86.078	12	5.3	24

874

875 **Table 2.** Site classification based on V_{s30} ranges, as proposed by the International Building
876 Code (IBC 2009)

Soil shear wave velocity, V_{s30} (m/s)	Soil Type	Site Class	Flag used in the model
$V_{s30} > 1500$	Hard Rock	A	0
$760 < V_{s30} \leq 1500$	Rock	B	1
$360 < V_{s30} \leq 760$	Soft Rock	C	2
$180 \leq V_{s30} \leq 360$	Stiff soil	D	3
$V_{s30} < 180$	Soft soil	E	4

877

878 **Table 3.** Details of ground motion recording stations used in the present study.

Sl. No.	Station				Sediment Depth (km)	Network	Site class
	Name	Code	Lat (°E)	Lon (°N)			
1	Agra	AGR	27.1607	78.0192	0.6	CIGN	D
2	Allahabad	ALB	25.1257	82.125	0.9	CIGN	C
3	Aligarh	ALG	27.9709	77.7888	0.6	CIGN	D
4	Almora	ALM	29.5478	79.5488	5	CIGN	D
5	Amethi	AMT	26.1878	81.679	1.2	CIGN	C
6	Bagaha	BAG	27.137	84.06	4	IITR	C
7	Balrampur	BRP	27.505	82.0556	5	CIGN	C
8	Bisalpur	BSR	28.2098	79.8053	3.6	CIGN	D
9	Chandigarh	CDG	30.7609	76.7674	4.4	CIGN	C
10	Faizabad	FZB	26.7308	82.1467	3.1	CIGN	D
11	Gonda	GDA	27.1162	82.2371	2.7	CIGN	D
12	Gorakhpur	GRP	26.9504	83.3167	4.7	CIGN	D
13	Hardoi	HAR	27.396	80.133	2.9	IITR	D
14	Haridwar	HDR	29.9752	78.0529	3.7	CIGN	D
15	Hisar	HSR	29.445	75.608	1	CIGN	D
16	Jaunpur	JAU	25.733	82.693	1	IITR	E
17	Kanti Path	KATNP	27.712	85.316	0.47	USGS	D
18	Kanpur	KNP	26.2694	80.5007	0.6	CIGN	C
19	Kishanganj	KSN	26.095	87.947	4	IITR	C
20	Kirtipur	KTP	27.682	85.272	0.04	CESMD	C
21	Kushinagar	KUS	26.865	83.957	3.5	IITR	E
22	Lakhimpur	LAK	27.949	80.79	4	IITR	C
23	Loharpur	LEH	27.709	80.89	3.5	IITR	C
24	Maharajganj	MAH	27.141	83.539	4	IITR	C
25	Meerut	MRT	29.1542	77.6178	2.5	CIGN	D
26	Motihari	MTR	26.63	84.899	2.5	CIGN	D
27	Narora	NAR	28.227	78.4256	1.7	CIGN	D
28	Pihani	PHN	27.6506	80.1606	3.6	CIGN	D
29	Patan(IGP)	ptn	25.622	85.149	1	IITR	E
30	Patan(KB)	PTN	27.681	85.318	0.35	CESMD	D
31	Rath	RAT	25.6261	79.5722	0.7	CIGN	B
32	Raibarely	RBL	26.3633	81.2694	1.2	CIGN	D
33	Rudraprayag	RPG	30.5463	79.0634	5	CIGN	D
34	Sikandrarao	SDR	27.7388	78.4739	1.6	CIGN	D
35	Thakurdwara	TDR	29.1489	78.8551	4.6	CIGN	D
36	Sanothimi	THM	27.681	85.377	0.46	CESMD	D
37	Tarikhhet	TKT	29.6132	79.4047	5	CIGN	D
38	Tribhuvan University	TVU	27.681	85.288	0.325	CESMD	D
39	Utraula	UTR	27.31	82.415	3.5	IITR	C
40	Varanasi	VNS	25.5536	82.8541	0.7	CIGN	D

880 **Table 4.** Basic definitions of Ground Motion Parameters estimated in this study.

Sl. No.	Ground Motion Parameter	Mathematical expression	Key characteristics
1.	Peak Ground Acceleration (PGA)	$PGA = \max(\text{abs}(a(t)))$ $a(t)$ – acceleration time history	Amplitude
2.	Peak Ground Velocity (PGV)	$PGA = \max(\text{abs}(V(t)))$ $V(t) =$	Amplitude
3.	Peak Ground Displacement (PGD)	$PGA = \max(\text{abs}(a(t)))$ $a(t)$ – acceleration time history	Amplitude
4.	Vertical to Horizontal, V/H (PGA)	PVA/PHA	Amplitude
5.	Predominant Frequency, F_p	Frequency corresponding to $\max(\text{fft}(a(t)))$ $\text{fft}(a(t))$ – Fourier amplitude spectrum	Frequency content
6.	Central Frequency, Ω	$\Lambda_n = \int_0^{\omega_N} \omega^n G(\omega) d\omega$ $\Omega = \sqrt{\frac{\lambda_2}{\lambda_0}}$ $G(\omega)$ - Power Spectral Density	Frequency content
7.	Shape factor, q	$q = \sqrt{1 - \frac{\lambda^2}{\lambda_0 \lambda_2}}$	Frequency content
8.	Velocity to Acceleration ratio, V/A	PGV/PGA	Frequency content
9.	Significant duration, T_{sig}	$T_{sig} = T_{5-95\%}([a(t)]^2)$	Duration
10.	Root-Mean-Square Acceleration, a_{RMS}	$\alpha_{RMS} = \sqrt{\frac{1}{T_{sig}} \int_{t_1}^{t_2} [a(t)]^2 dt}$	Amplitude, Frequency content
11.	Arias Intensity, I_a	$I_a = \frac{\pi}{2g} \int_0^{\infty} [a(t)]^2 dt$	Amplitude, Frequency content, Duration
12.	Cumulative Absolute Velocity, CAV	$CAV = \int_0^{T_{sig}} a(t) dt$	Amplitude, Frequency content, Duration
13.	Characteristic Intensity, I_c		Amplitude, Frequency content, Duration
14.	Acceleration Spectrum Intensity, ASI	$ASI = \int_{0.1}^{0.5} PSA(\epsilon_g = 0.05, T) dT$	Amplitude, Frequency content

15.	Velocity Spectrum Intensity, VSI	$VSI = \int_{0.1}^{2.5} PSV(\epsilon_v = 0.05, T) dT$	Amplitude, Frequency content
16.	Total Energy, E_{acc}	$E_{acc} = \int_0^{\infty} \int_0^{\infty} G(t, \omega) d\omega dt$	Amplitude, Frequency content, Duration
17.	Spectral Centroid, E_w	$E_w(\omega) = \frac{\int_{t_1}^{t_2} \int_0^{N_{yq}} \omega G(t, \omega) d\omega dt}{\int_{t_1}^{t_2} \int_0^{N_{yq}} G(t, \omega) d\omega dt}$	
18.	Spectral Standard deviation, S_w	$S_w^2(\omega) = \frac{\int_{t_1}^{t_2} \int_0^{N_{yq}} [\omega - E(\omega)]^2 G(t, \omega) d\omega dt}{\int_{t_1}^{t_2} \int_0^{N_{yq}} G(t, \omega) d\omega dt}$	
19.	Temporal Centroid, E_t	$E_t(t) = \frac{\int_{t_1}^{t_2} \int_0^{N_{yq}} t G(t, \omega) d\omega dt}{\int_{t_1}^{t_2} \int_0^{N_{yq}} G(t, \omega) d\omega dt}$	
20.	Temporal Standard deviation, S_t	$S_t^2(t) = \frac{\int_{t_1}^{t_2} \int_0^{N_{yq}} [t - E(t)]^2 G(t, \omega) d\omega dt}{\int_{t_1}^{t_2} \int_0^{N_{yq}} G(t, \omega) d\omega dt}$	
21.	Correlation between time and frequency, $\rho(\omega, t)$	$\rho(\omega, t) = \frac{\int_{t_1}^{t_2} \int_0^{N_{yq}} [t - E(t)][\omega - E(\omega)] G(t, \omega) d\omega dt}{S(t)S(\omega) \int_{t_1}^{t_2} \int_0^{N_{yq}} G(t, \omega) d\omega dt}$	

882 **Table 5.** Ground motion parameters obtained for strong motion records of two near field
883 stations Kirtipur(KTP), Sanothimi (THM) and one IGP station Amethi (AMT): H-horizontal
884 and V-vertical components.

Sl. No.	GMP	Unit	GROUND MOTION CHARACTERISTICS					
			KTP		THM		AMT	
			H	V	H	V	H	V
1	PHA	g	0.209	0.141	0.142	0.185	0.016	0.010
2	PHV	cm/s	30.13	29.998	63.596	41.651	3.386	1.340
3	PHD	cm	27.857	31.540	59.137	34.595	1.241	0.646
4	F _p	s ⁻¹	0.153	0.216	0.459	0.229	0.215	0.138
5	CF	Hz	4.690	9.901	1.334	6.965	4.148	8.566
6	q	-	3.575	0.280	0.245	0.290	0.716	0.897
7	V/A	s	0.554	0.565	0.874	0.689	0.638	0.575
8	T _{sig}	s	37.984	40.959	40.637	36.253	97.85	109.8
9	a _{RMS}	g	0.033	0.023	0.045	0.032	0.004	0.002
10	I _a	cm/s	70.872	37.495	140.545	64.826	2.138	0.772
11	CAV	cm/s	890.17	694.285	1429.97	842.47	252.5	167.8
12	I _c	$\sqrt{cm^3 s^{-5}}$	1131.0	691.25	1865.28	1074.40	64.90	29.38
13	V/H	-	0.677		1.307		0.612	
14	ASI	cm/s	192.18	103.314	97.490	152.518	16.40	7.973
15	VSI	cm	64.663	52.933	139.835	92.841	9.901	4.782
16	E _{acc}	cm/s	140.88	37.857	312.318	72.288	4.351	0.763
17	E _w	Hz	3.552	7.592	0.541	4.578	2.893	6.666
18	S _w	Hz	2.523	5.596	1.430	4.776	2.584	5.031
19	E _t	s	40.444	38.416	46.178	39.918	239.4	230.4
20	S _t	s	12.836	13.325	12.762	11.755	59.53	71.44
21	ρ	-	0.196	0.013	0.0003	0.016	-0.01	-0.031

886 **Table 6.** Instantaneous frequency of the IMFs in Hz with mean – μ and standard deviation –
 887 σ and respective percentage variances contributing to the total variability of the data.

Station Code		IMF1	IMF2	IMF3	IMF4	IMF5	IMF6	IMF7	IMF8	IMF9	IMF10	
ALB	IF	μ	19.70	9.628	4.966	2.513	1.232	0.570	0.245	0.127	0.068	0.031
		σ	± 6.78	± 3.68	± 2.07	± 1.32	± 0.71	± 0.42	± 0.22	± 0.11	± 0.04	± 0.03
		% Var	32.79	27.23	14.72	12.52	7.497	2.860	1.660	0.238	0.111	0.048
VNS	IF	μ	16.58	7.300	3.687	1.903	0.877	0.397	0.183	0.096	0.051	0.025
		σ	± 8.59	± 3.72	± 2.03	± 1.14	± 0.65	± 0.33	± 0.15	± 0.06	± 0.04	± 0.02
		% Var	52.43	24.38	12.16	3.927	2.713	1.403	0.205	0.038	0.015	0.012
BRP	IF	μ	17.51	7.979	3.407	1.991	0.881	0.274	0.134	0.083	0.042	0.018
		σ	± 7.95	± 4.01	± 2.26	± 1.12	± 0.62	± 0.29	± 0.08	± 0.09	± 0.03	± 0.02
		% Var	23.97	32.65	26.16	10.45	4.750	1.426	0.306	0.050	0.038	0.016
GDA	IF	μ	17.48	7.720	3.703	1.645	0.882	0.256	0.140	0.081	0.043	0.019
		σ	± 6.65	± 3.74	± 2.17	± 1.27	± 0.68	± 0.26	± 0.09	± 0.05	± 0.03	± 0.02
		% Var	29.45	21.76	24.84	15.16	8.298	2.354	0.432	0.076	0.100	0.046
FZB	IF	μ	17.70	7.489	3.795	1.862	0.881	0.279	0.151	0.076	0.050	0.021
		σ	± 7.99	± 4.29	± 1.94	± 1.17	± 0.69	± 0.28	± 0.08	± 0.11	± 0.09	± 0.02
		% Var	29.14	30.67	22.64	10.12	4.240	2.579	0.630	0.184	0.019	0.017
AMT	IF	μ	19.04	8.486	3.775	2.094	0.920	0.349	0.184	0.097	0.044	0.027
		σ	± 7.69	± 4.20	± 2.35	± 1.25	± 0.74	± 0.33	± 0.15	± 0.12	± 0.07	± 0.02
		% Var	24.72	24.72	21.33	18.02	5.654	4.030	0.620	0.276	0.099	0.010
RAT	IF	μ	18.92	8.806	4.244	2.049	1.139	0.540	0.239	0.116	0.067	0.033
		σ	± 7.23	± 3.99	± 2.22	± 1.16	± 0.56	± 0.36	± 0.19	± 0.06	± 0.04	± 0.02
		% Var	26.67	24.69	44.79	4.426	0.985	0.356	0.153	0.113	0.036	0.022
NAR	IF	μ	22.83	10.72	5.059	2.479	1.200	0.435	0.218	0.111	0.058	0.029
		σ	± 7.67	± 5.10	± 2.65	± 1.33	± 0.83	± 0.44	± 0.16	± 0.09	± 0.05	± 0.02
		% Var	24.77	36.14	19.63	8.836	3.579	1.979	0.536	0.080	0.053	0.014
PHN	IF	μ	20.18	9.488	4.640	2.195	0.991	0.340	0.159	0.096	0.045	0.026
		σ	± 7.60	± 4.47	± 2.26	± 1.41	± 0.84	± 0.35	± 0.15	± 0.08	± 0.03	± 0.02
		% Var	27.94	27.62	18.05	10.69	6.264	5.163	1.364	0.206	0.057	0.020
BSR	IF	μ	20.00	9.475	4.318	2.138	0.896	0.281	0.140	0.077	0.048	0.021
		σ	± 7.30	± 4.49	± 2.52	± 1.37	± 0.79	± 0.29	± 0.12	± 0.06	± 0.03	± 0.02
		% Var	17.96	24.25	21.65	11.88	10.38	9.778	3.494	0.258	0.047	0.022
ALM	IF	μ	19.42	9.942	5.163	2.671	1.358	0.668	0.293	0.133	0.074	0.038
		σ	± 7.04	± 3.84	± 2.11	± 1.22	± 0.63	± 0.36	± 0.21	± 0.07	± 0.03	± 0.02
		% Var	36.36	35.99	19.40	4.533	1.400	0.655	0.362	0.110	0.038	0.007
TDR	IF	μ	20.35	9.867	4.806	2.334	0.966	0.289	0.131	0.073	0.041	0.017
		σ	± 7.38	± 4.32	± 2.35	± 1.27	± 0.87	± 0.29	± 0.18	± 0.08	± 0.04	± 0.02
		% Var	16.47	17.76	14.36	18.94	11.32	14.21	4.684	0.700	0.171	0.056
MRT	IF	μ	18.22	9.203	4.666	2.358	1.081	0.436	0.212	0.112	0.061	0.026
		σ	± 6.12	± 3.64	± 2.04	± 1.17	± 0.68	± 0.31	± 0.12	± 0.07	± 0.03	± 0.02
		% Var	25.88	24.01	16.52	13.73	8.328	9.395	1.538	0.295	0.056	0.019

HDR	IF	μ	19.84	9.646	4.729	2.357	1.124	0.398	0.175	0.098	0.050	0.025
		σ	± 7.34	± 4.25	± 2.17	± 1.16	± 0.68	± 0.41	± 0.15	± 0.06	± 0.03	± 0.02
		% Var	28.47	30.85	15.83	9.563	6.696	4.296	2.824	0.393	0.103	0.041
RPG	IF	μ	20.71	10.29	5.125	2.586	1.310	0.651	0.278	0.134	0.072	0.038
		σ	± 7.14	± 4.20	± 2.29	± 1.31	± 0.76	± 0.36	± 0.21	± 0.08	± 0.04	± 0.06
		% Var	23.76	33.56	22.73	9.678	4.705	1.205	0.743	0.223	0.033	0.013
CDG	IF	μ	20.38	10.08	4.979	2.335	1.025	0.379	0.178	0.095	0.051	0.027
		σ	± 6.80	± 3.98	± 2.25	± 1.30	± 0.80	± 0.32	± 0.12	± 0.06	± 0.03	± 0.02
		% Var	6.676	9.831	17.84	27.91	18.41	11.72	5.002	0.979	0.229	0.273
HSR	IF	μ	20.00	9.811	4.938	2.454	1.176	0.548	0.256	0.124	0.069	0.035
		σ	± 7.06	± 4.05	± 2.25	± 1.20	± 0.73	± 0.34	± 0.17	± 0.08	± 0.04	± 0.05
		% Var	28.83	20.21	25.46	10.46	13.44	1.630	0.803	0.244	0.066	0.019
BAG	IF	μ	6.426	2.845	1.253	0.762	0.259	0.159	0.062	0.031	0.017	0.003
		σ	± 5.93	± 2.62	± 1.11	± 0.71	± 0.38	± 0.27	± 0.04	± 0.02	± 0.02	± 0.01
		% Var	30.68	34.34	21.09	7.768	5.160	2.645	0.829	0.234	0.045	0.073
HAR	IF	μ	7.851	2.848	1.215	0.576	0.279	0.145	0.077	0.041	0.021	0.011
		σ	± 7.11	± 2.03	± 0.76	± 0.33	± 0.15	± 0.10	± 0.04	± 0.06	± 0.04	± 0.32
		% Var	7.389	21.83	27.76	25.53	12.42	5.613	1.087	0.291	0.036	0.058
JAU	IF	μ	7.990	3.371	1.606	0.774	0.369	0.215	0.113	0.058	0.027	0.009
		σ	± 6.63	± 2.31	± 1.13	± 0.51	± 0.24	± 0.38	± 0.08	± 0.06	± 0.01	± 0.004
		% Var	20.88	27.03	25.32	17.92	5.133	1.632	0.359	0.105	0.031	0.037
KSN	IF	μ	14.59	3.487	1.567	0.669	0.253	0.107	0.056	0.031	0.015	0.010
		σ	± 11.5	± 3.59	± 1.22	± 0.56	± 0.23	± 0.10	± 0.04	± 0.04	± 0.04	± 0.06
		% Var	10.08	9.004	12.53	18.87	13.70	19.98	2.349	0.307	0.030	0.031
LAK	IF	μ	5.829	2.656	1.282	0.610	0.300	0.146	0.076	0.038	0.025	0.014
		σ	± 6.03	± 1.69	± 0.66	± 0.30	± 0.18	± 0.09	± 0.04	± 0.04	± 0.12	± 0.03
		% Var	33.73	18.04	18.08	11.62	8.982	4.533	0.637	0.113	0.061	0.083
LEH	IF	μ	7.345	2.953	1.370	0.658	0.331	0.165	0.086	0.047	0.029	0.015
		σ	± 6.54	± 1.95	± 0.79	± 0.34	± 0.24	± 0.12	± 0.04	± 0.03	± 0.22	± 0.23
		% Var	16.13	30.27	22.75	12.75	7.288	2.032	0.887	0.197	0.008	0.013
MAH	IF	μ	11.26	3.167	1.557	0.716	0.323	0.158	0.071	0.044	0.023	0.014
		σ	± 9.91	± 2.36	± 1.11	± 0.64	± 0.21	± 0.35	± 0.06	± 0.03	± 0.04	± 0.02
		% Var	15.62	28.03	20.59	13.08	7.104	6.375	1.819	0.284	0.214	0.013
MTR	IF	μ	8.262	3.015	1.500	0.646	0.258	0.124	0.078	0.039	0.019	0.011
		σ	± 7.62	± 2.34	± 1.24	± 0.45	± 0.21	± 0.07	± 0.09	± 0.10	± 0.16	± 0.04
		% Var	7.804	21.29	26.96	18.31	10.27	10.28	1.537	0.133	0.039	0.013
PTN	IF	μ	8.820	3.439	1.336	0.792	0.405	0.215	0.092	0.048	0.018	0.009
		σ	± 7.99	± 3.36	± 1.05	± 0.79	± 0.45	± 0.21	± 0.15	± 0.17	± 0.03	± 0.02
		% Var	22.67	31.42	24.55	13.15	2.467	1.681	0.249	0.091	0.014	0.016
UTR	IF	μ	5.134	2.150	1.030	0.500	0.248	0.128	0.070	0.036	0.010	0.007
		σ	± 3.71	± 1.22	± 0.51	± 0.23	± 0.12	± 0.07	± 0.05	± 0.02	± 0.02	± 0.01
		% Var	19.71	38.59	20.95	12.17	5.921	0.873	0.161	0.047	0.083	0.034

KTP	IF	μ	11.56	4.338	2.023	1.110	0.562	0.235	0.151	0.073	0.044	0.019
		σ	± 8.54	3.28	1.44	0.85	0.56	0.17	0.18	0.05	0.03	0.02
		% Var	37.91	31.75	16.99	5.215	1.365	2.478	0.329	0.037	0.012	0.015
PTN	IF	μ	9.353	4.844	2.066	1.041	0.387	0.217	0.093	0.052	0.020	0.013
		σ	± 6.60	± 3.71	± 1.34	± 0.99	± 0.31	± 0.23	± 0.06	± 0.06	± 0.01	± 0.06
		% Var	9.515	14.57	11.99	19.51	30.81	6.901	0.689	0.730	0.626	0.440
THM	IF	μ	17.71	5.896	2.328	1.064	0.291	0.189	0.061	0.036	0.020	0.012
		σ	± 9.87	± 4.40	± 2.35	± 1.46	± 0.76	± 0.11	± 0.12	± 0.03	± 0.01	± 0.02
		% Var	0.336	4.102	13.48	12.13	49.46	7.612	1.116	0.219	0.145	0.111
TVU	IF	μ	10.06	3.029	1.151	0.607	0.275	0.133	0.066	0.051	0.031	0.014
		σ	± 9.50	± 2.31	± 1.11	± 0.69	± 0.47	± 0.38	± 0.10	± 0.04	± 0.05	± 0.10
		% Var	3.450	16.23	30.50	21.45	23.44	1.222	0.236	0.023	0.009	0.008
KATNP	IF	μ	38.11	10.38	4.506	2.125	0.846	0.300	0.194	0.117	0.050	0.026
		σ	± 30.3	± 9.17	± 3.36	± 1.43	± 0.80	± 0.20	± 0.14	± 0.15	± 0.04	± 0.04
		% Var	1.106	3.661	4.668	4.271	27.74	54.08	2.562	0.188	0.071	0.075

889 **Table 7.** Regression Coefficients of the GMPE for all the 21 GMPs.

SL. NO.	GMP	Co mp	c ₀	c ₁	c ₂	c ₃	c ₄	c ₅	c ₆	c ₇	σ _{inter}	σ _{intra}	σ
1	PGA	H	-2.733	0.640	-0.0029	-0.909	3.23E+06	-1.392	-0.028	0.106	0.184	0.427	0.465
		V	-5.951	0.867	-0.0020	-0.857	3.73E+06	-1.687	-0.047	0.071	0.159	0.431	0.459
2	PGV	H	5.577	0.664	-0.0024	-1.497	3.67E+06	-1.400	-0.089	0.262	0.338	0.483	0.589
		V	3.857	0.788	-0.0018	-1.502	2.95E+06	-1.405	-0.123	0.094	0.399	0.446	0.598
3	PGD	H	4.405	0.923	-0.0016	-1.729	2.90E+06	-1.386	-0.225	0.433	0.421	0.505	0.657
		V	-2.492	1.475	-0.0013	-1.362	1.45E+06	-1.386	-0.220	0.115	0.268	0.694	0.744
4	V/H		4.442	-0.875	-0.0002	0.277	4.45E+05	-1.402	0.203	-0.282	0.356	0.810	0.885
5	F _p	H	4.657	-0.967	0.0001	0.386	9.53E+04	-1.190	0.082	-0.151	0.503	1.049	1.163
		V	4.254	-0.538	0.0012	0.103	4.51E+06	-2.628	0.045	-0.207	0.194	0.333	0.385
6	Ω	H	5.105	-0.388	0.0014	-0.159	1.17E+06	-1.885	0.039	-0.099	0.146	0.264	0.302
		V	-0.783	0.093	0.0003	-0.069	1.14E+04	-1.038	-0.018	0.042	0.068	0.155	0.169
7	Q	H	-12.126	0.862	-0.0005	0.733	2.85E+06	-1.015	-0.005	0.004	0.130	0.162	0.208
		V	-3.236	0.409	0.0002	-0.296	1.71E+05	-1.511	-0.063	0.200	0.122	0.404	0.422
8	V/A	H	-5.202	0.580	-0.0002	-0.115	4.17E+05	-3.690	-0.095	0.093	0.217	0.537	0.579
		V	8.128	-0.660	0.0018	0.136	7.24E+05	-1.797	0.066	0.122	0.341	0.404	0.528
9	T _{sig}	H	7.131	-0.591	0.0015	0.291	2.12E+05	-1.557	0.102	0.077	0.248	0.445	0.510
		V	0.970	-0.104	0.0007	-0.187	6.75E+05	-1.748	-0.009	-0.044	0.074	0.338	0.346
10	a _{RMS}	H	-12.379	1.453	-0.0028	-0.630	4.09E+06	-2.028	-0.097	0.100	0.327	0.511	0.607
		V	15.956	-0.329	-0.0012	-3.093	1.78E+05	-0.845	-0.054	0.104	0.259	0.515	0.576
11	I _a	H	109.891	-0.984	0.0033	-14.714	8.57E+03	-0.315	-0.102	0.298	0.331	0.785	0.852
		V	39.312	-0.819	-0.0010	-5.305	1.78E+05	-0.842	-0.013	0.287	0.408	0.743	0.848
12	CAV	H	10.004	0.204	-0.0006	-1.049	2.90E+06	-1.386	-0.042	0.242	0.261	0.408	0.484
		V	18.599	-0.519	-0.0002	-1.624	1.82E+06	-1.205	0.028	0.216	0.317	0.353	0.474
13	I _c	H	-0.955	1.494	-0.0035	-0.964	4.02E+07	-2.026	-0.113	0.224	0.343	0.642	0.728
		V	10.418	0.507	-0.0022	-1.871	2.67E+06	-1.399	-0.047	0.212	0.372	0.614	0.718
14	ASI	H	0.968	0.838	-0.0030	-0.643	7.14E+06	-1.579	-0.013	0.102	0.188	0.415	0.456
		V	-0.221	1.086	-0.0018	-0.949	1.50E+04	-0.849	-0.013	0.058	0.200	0.397	0.445
15	VSI	H	1.566	0.984	-0.0027	-1.036	5.18E+06	-1.529	-0.077	0.341	0.325	0.555	0.644
		V	3.344	1.023	-0.0020	-1.438	2.19E+05	-1.013	-0.161	0.148	0.384	0.482	0.616
16	E _{acc}	H	17.227	0.217	-0.0041	-2.654	3.40E+06	-1.281	-0.113	0.323	0.391	0.815	0.904
		V	64.559	-0.847	0.0025	-9.115	1.75E+04	-0.492	-0.021	0.271	0.312	0.719	0.783
17	E _w	H	3.889	-0.597	0.0006	0.221	1.75E+05	-3.366	0.036	-0.219	0.228	0.400	0.461
		V	5.805	-0.534	0.0013	-0.158	1.05E+06	-1.434	0.070	-0.117	0.194	0.309	0.365
18	S _w	H	3.665	-0.441	0.0014	0.008	6.78E+06	-2.216	0.001	-0.144	0.129	0.387	0.408
		V	3.488	-0.208	0.0012	-0.164	1.20E+05	-1.451	0.047	-0.091	0.015	0.304	0.304
19	E _t	H	4.078	-0.205	0.0005	0.440	3.37E+06	-3.109	0.028	-0.019	0.172	0.293	0.340
		V	4.182	-0.253	0.0006	0.465	3.76E+05	-1.669	0.045	-0.030	0.173	0.343	0.384
20	S _t	H	5.042	-0.483	0.0009	0.392	3.45E+05	-1.646	0.072	0.009	0.265	0.318	0.414
		V	4.276	-0.435	0.0006	0.520	2.13E+05	-1.549	0.086	-0.005	0.196	0.356	0.406
21	ρ*	H	1.179	-0.160	0.0005	-0.036	-5.64E-03	-2.500	0.044	-0.069	0.050	0.172	0.179
		V	0.625	-0.085	0.0003	-0.007	-7.50E-05	0.893	0.025	-0.048	0.027	0.174	0.176

890 *The coefficients of ρ are derived for non-logarithmic values

891 **Table 8.** Rankings and normalized average rankings obtained for the GMPEs of PGA

Rank	GMPE	EDR	EDR _N	GMPE	LH	LH _N	GMPE	LLH	LLH _N
1	This study	0.575	1.000	This study	0.536	1.071	BSSA14	1.701	1.000
2	SSSA17	0.806	1.402	CY14	0.526	1.051	ASK14	1.787	1.051
3	CB14	1.596	2.777	K06	0.435	0.870	This study	1.792	1.053
4	K06	1.754	3.052	ASK14	0.411	0.823	SSSA17	1.820	1.070
5	ASK14	1.783	3.102	BSSA14	0.409	0.817	CB14	1.879	1.105
6	CY14	1.883	3.276	SSSA17	0.377	0.755	CY14	1.895	1.114
7	BSSA14	1.968	3.424	CB14	0.325	0.650	Z06	2.348	1.381
8	Z06	2.563	4.461	Z06	0.109	0.218	K06	2.582	1.518

892

893 **Table 9.** Rankings and normalized average rankings obtained for the GMPEs of PGV

Rank	GMPE	EDR	EDR _N	GMPE	LH	LH _N	GMPE	LLH	LLH _N
1	This study	0.811	1.000	This study	0.536	1.071	K06	1.564	1.000
2	K06	0.969	1.195	CY14	0.508	1.016	This study	1.611	1.030
3	SSSA17	1.089	1.344	SSSA17	0.493	0.986	SSSA17	1.686	1.078
4	CB14	1.754	2.163	K06	0.440	0.879	CY14	1.885	1.205
5	CY14	1.834	2.263	ASK14	0.324	0.647	CB14	1.933	1.236
6	ASK14	1.856	2.289	BSSA14	0.312	0.625	ASK14	1.939	1.239
7	BSSA14	2.118	2.613	CB14	0.246	0.492	BSSA14	1.965	1.256

894

895 **Table 10.** Rankings and normalized average rankings obtained for the GMPEs of PGD

Rank	GMPE	EDR	EDR _N	GMPE	LH	LH _N	GMPE	LLH	LLH _N
1	This study	1.069	1.000	This study	0.517	1.035	This study	1.373	1.000
2	SSSA17	1.393	1.303	SSSA17	0.393	0.786	SSSA17	1.519	1.106
3	CB08	2.183	2.042	CB08	0.129	0.259	CB08	2.127	1.549

896

897 **Table 11.** Rankings and normalized average rankings obtained for the GMPEs of I_a

Rank	GMPE	EDR	EDR _N	GMPE	LH	LH _N	GMPE	LLH	LLH _N
1	This study	1.056	1.000	This study	0.588	1.176	FPG14	1.532	1.000
2	FPG14	2.180	2.063	FPG14	0.530	1.059	This study	1.583	1.034
3	SS98	3.466	3.281	SS98	0.215	0.431	SS98	2.076	1.355
4	A09	5.838	5.526	A09	0.004	0.008	A09	3.496	2.282

898

899 **Table 12.** Rankings and normalized average rankings obtained for the GMPEs of CAV

Rank	GMPE	EDR	EDR _N	GMPE	LH	LH _N	GMPE	LLH	LLH _N
1	This study	0.595	1.000	CB10	0.622	1.243	CB10	1.570	1.000
2	CB10	0.706	1.186	This study	0.498	0.995	This study	1.650	1.051
3	DW13	0.864	1.453	DW13	0.376	0.753	FPG14	1.655	1.054
4	FPG14	1.125	1.892	FPG14	0.330	0.659	DW13	1.932	1.231

900

901 **Table 13.** Final rankings for the GMPEs of PGA, PGV and PGD with respect to logic-tree-
 902 based normalized ranking indices

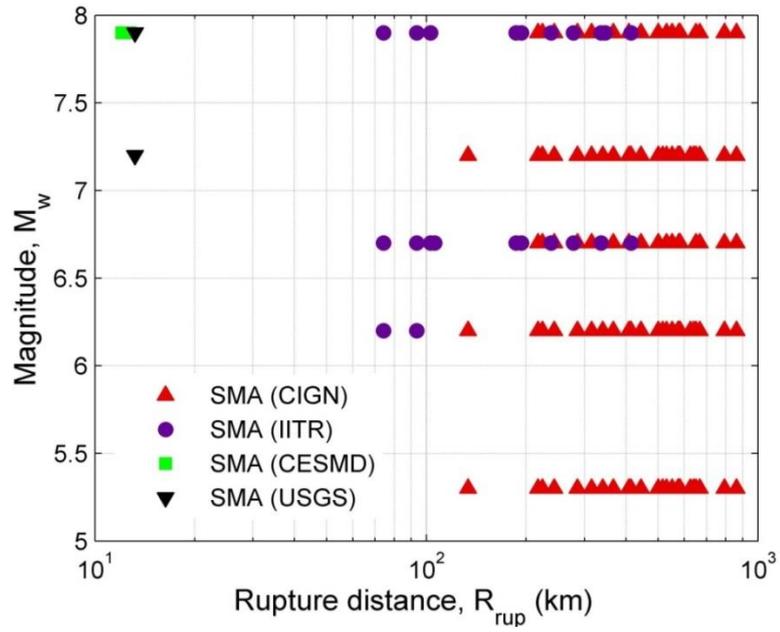
PGA			PGV			PGD		
Rank	GMPE	RI _N	Rank	GMPE	RI _N	Rank	GMPE	RI _N
1	This study	1.031	1	This study	1.025	1	This study	1.009
2	SSSA17	1.157	2	K06	1.067	2	SSSA17	1.125
3	CB14	1.827	3	SSSA17	1.188	3	CB08	1.473
4	ASK14	2.019	4	CB14	1.514			
5	K06	2.123	5	ASK14	1.616			
6	BSSA14	2.166	6	CY14	1.687			
7	CY14	2.179	7	BSSA14	1.777			
8	Z06	2.630						

903

904 **Table 14.** Final rankings for the GMPEs of I_a and CAV with respect to logic-tree-based
905 normalized ranking indices

I_a			CAV		
Rank	GMPE	RI_N	Rank	GMPE	RI_N
1	This study	1.052	1	This study	1.012
2	FPG14	1.546	2	CB10	1.154
3	SS98	2.087	3	DW13	1.222
4	A09	3.335	4	FPG14	1.374

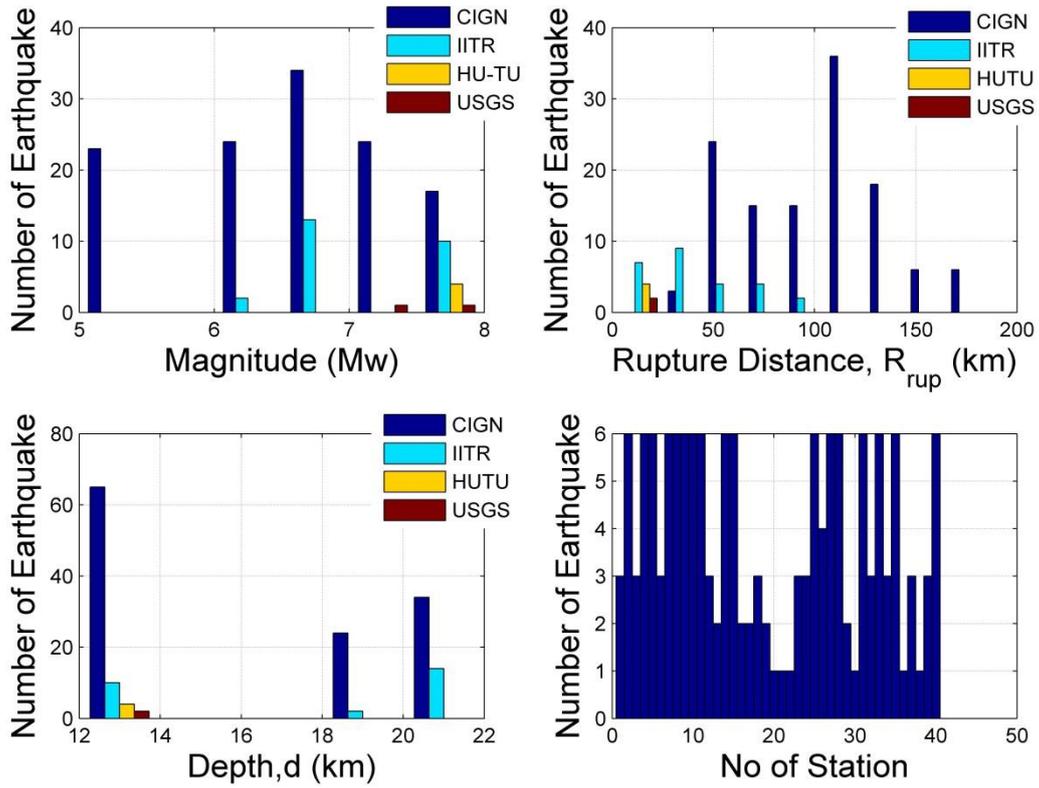
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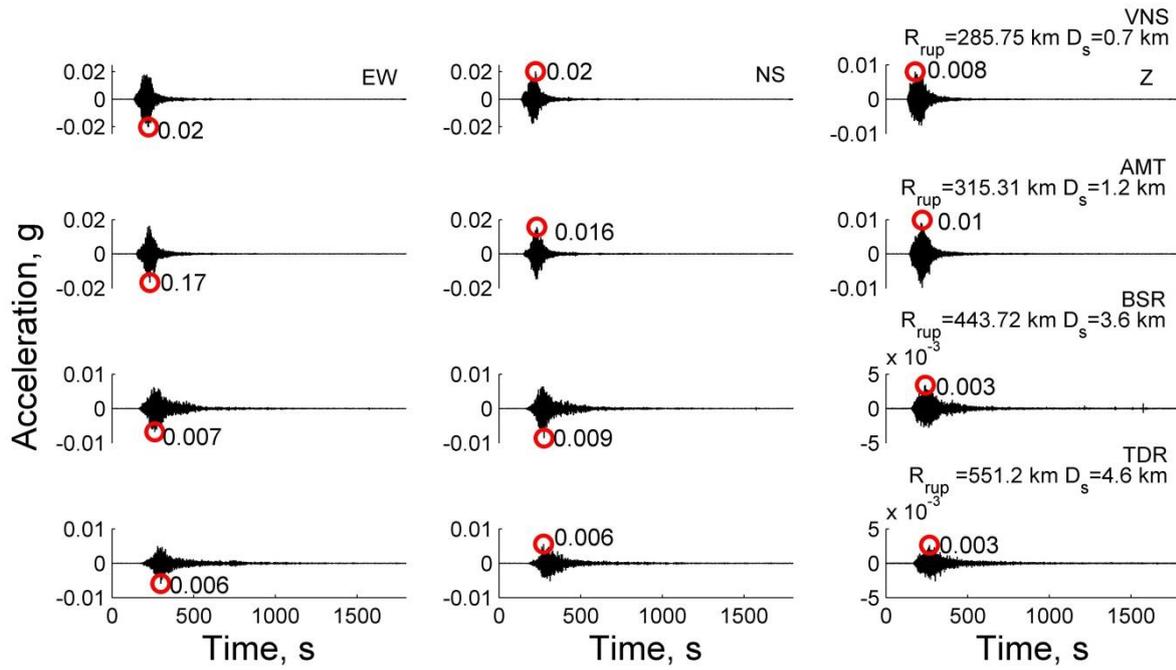
912 **Fig. 2** Distribution of ground motion records with respect to Moment magnitude (M_w) and

913 Rupture distance (R_{rup})



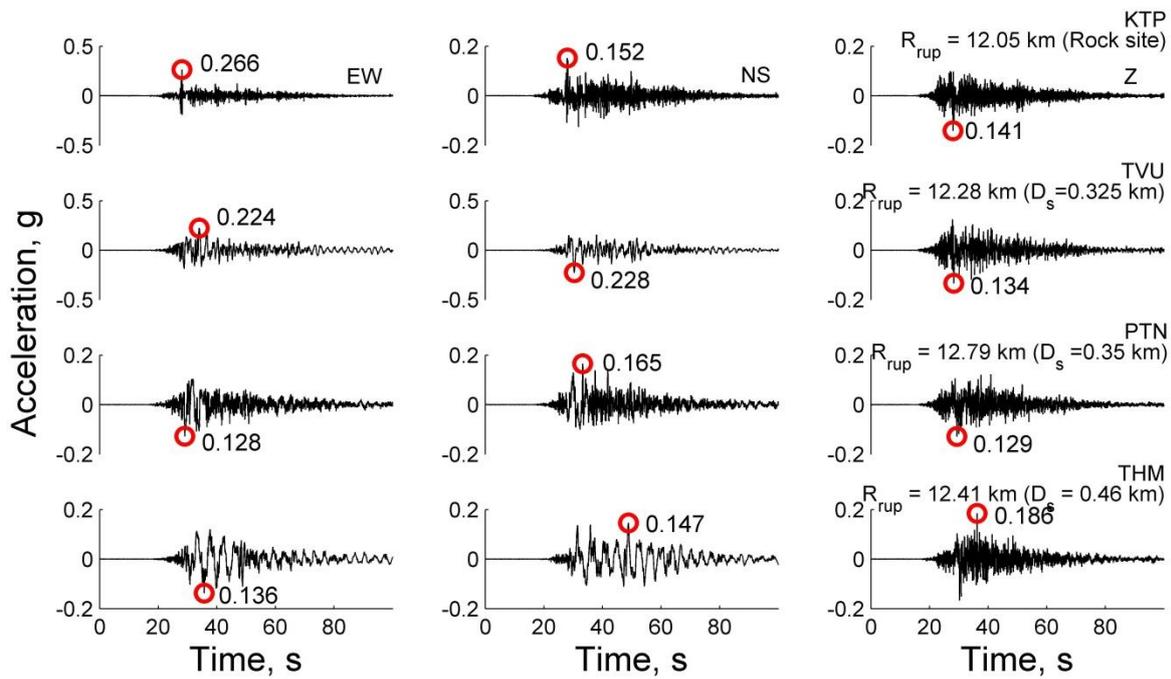
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915 **Fig. 3** Histograms showing the distribution of ground motion records used in the current
 916 study with respect to Moment magnitude (M_w), Rupture distance (R_{rup}) and Hypocentral
 917 depth (d)



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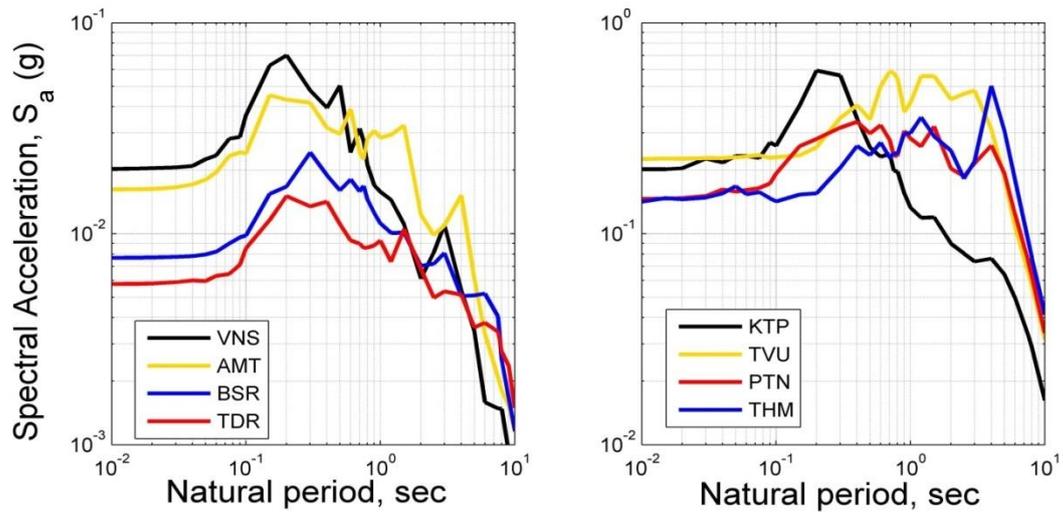
(a)



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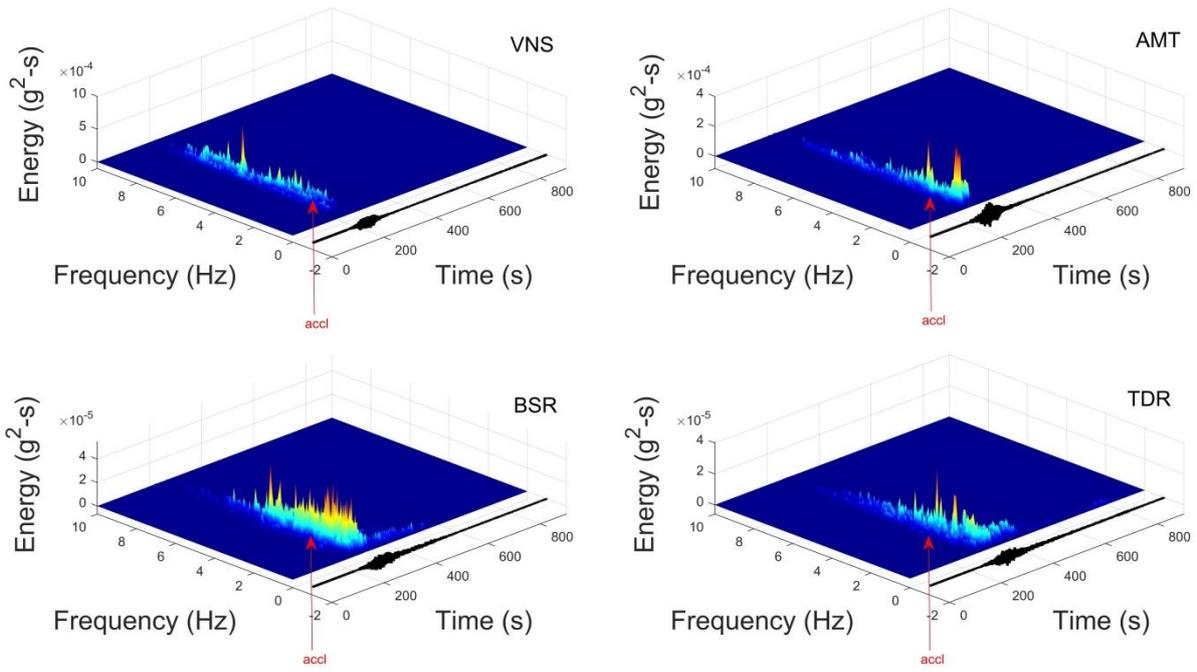
(b)

922 **Fig. 4** Acceleration time histories (EW, NS and UD component) of the M_w 7.9 Nepal
923 earthquake, recorded at stations: (a) VNS, AMT, BSR and TDR in the IG basin and (b) KTP,
924 TVU, PTN and THM in the Kathmandu basin



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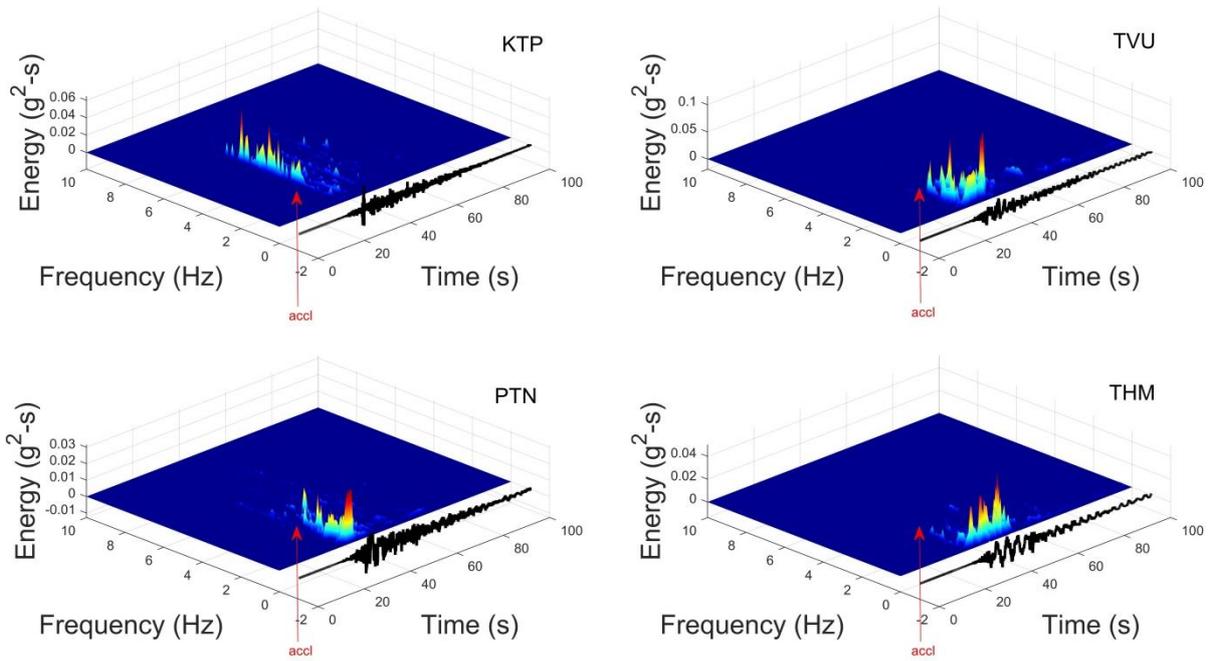
Fig. 5 Acceleration response spectra obtained for the time histories (of Figure 4) of stations: (a) VNS, AMT, BSR and TDR in the IG basin and (b) KTP, TVU, PTN and THM in the Kathmandu basin



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(a)

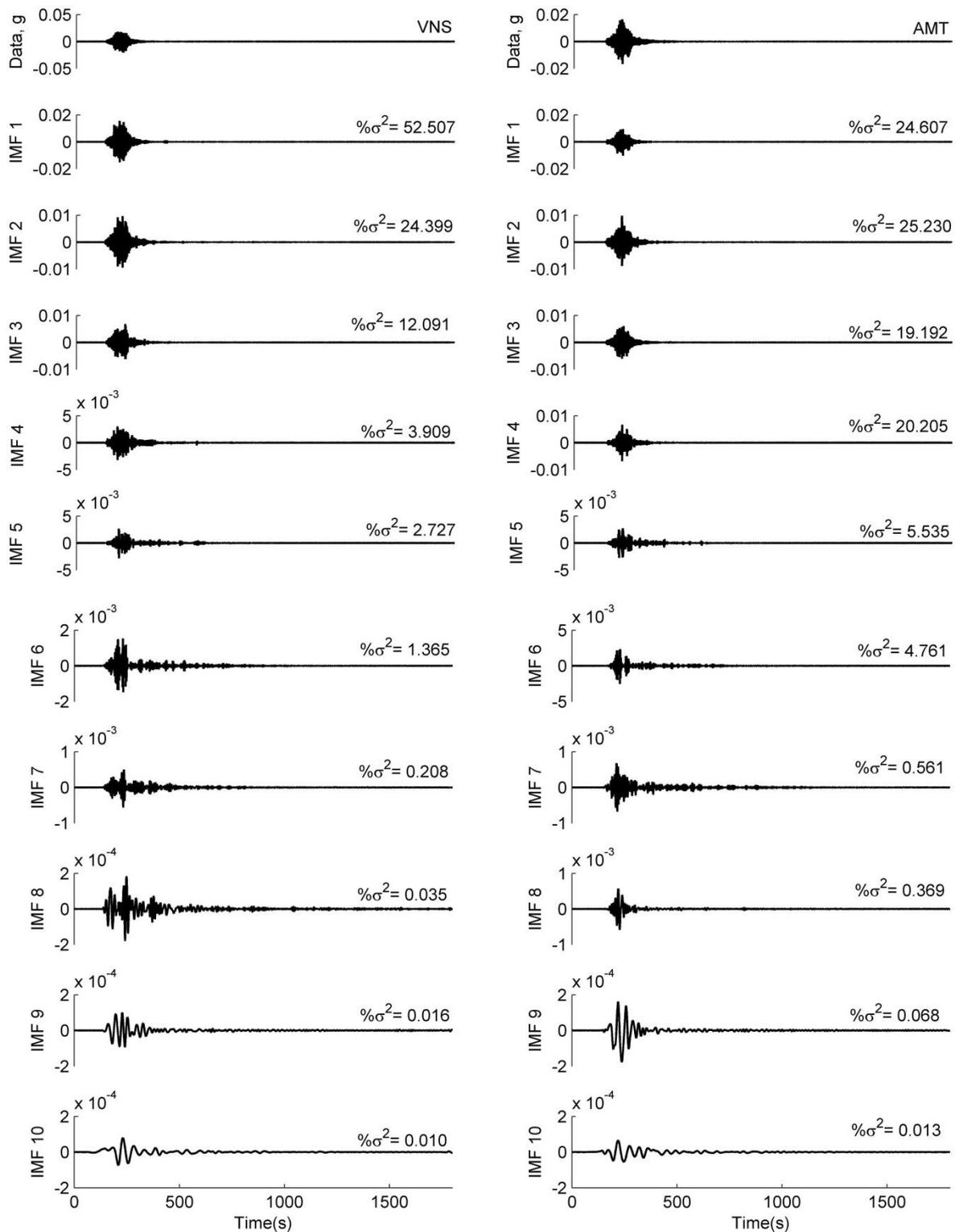


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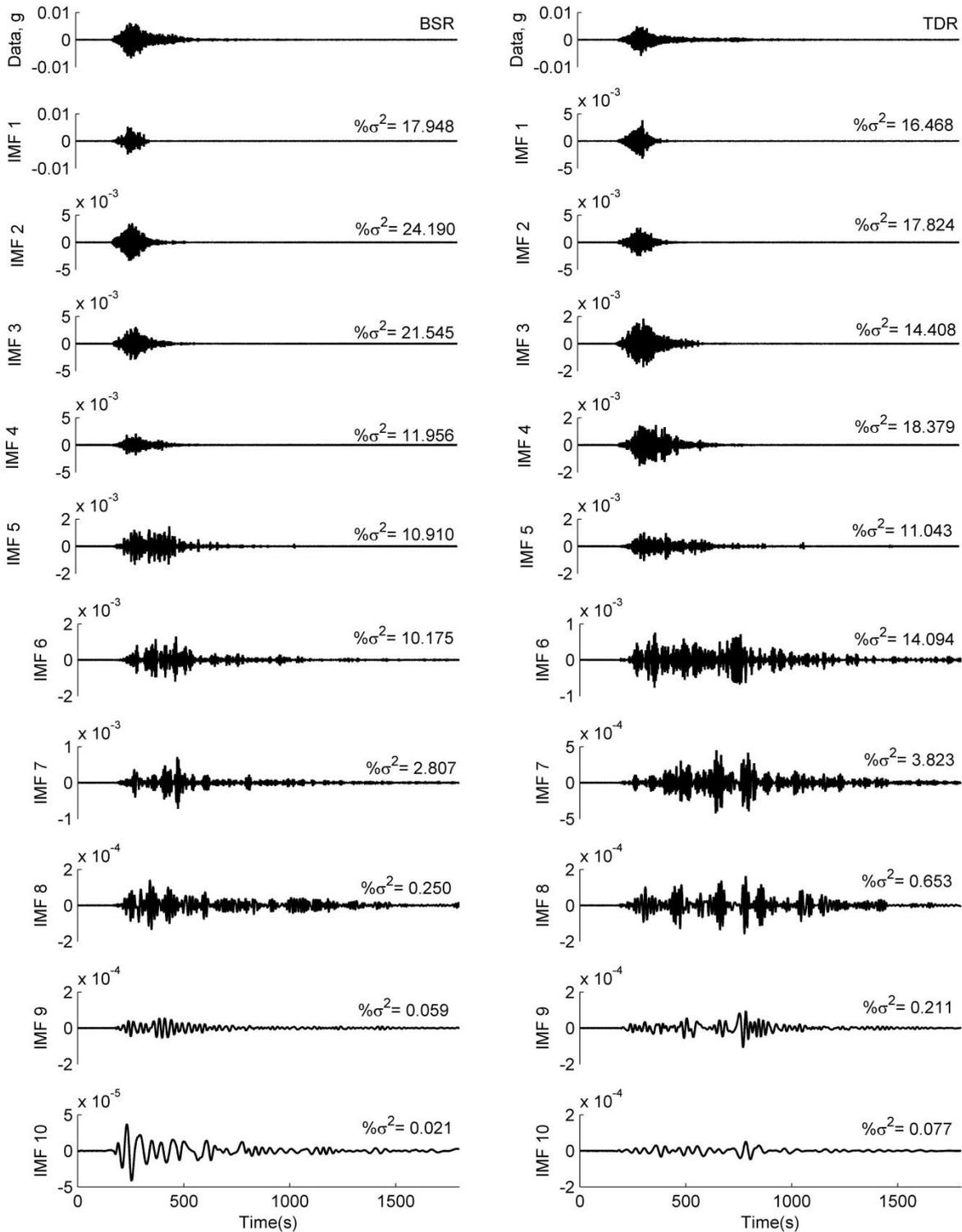
933 **Fig. 6** Hilbert Energy Spectrum for acceleration time histories (EW component) of the M_w
 934 7.9 Nepal earthquake, recorded at stations (a) VNS, AMT, BSR and TDR in the IG basin and
 935 (b) KTP, TVU, PTN and THM in the Kathmandu basin. Hilbert Energy Spectrum for
 936 acceleration time histories (NS and Vertical component) for stations in IG basin and
 937 Kathmandu basin are shown in electronic supplement (NS – Figure A5 and Vertical Figure
 938 A6)



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941 **Fig. 7** Intrinsic mode functions (IMFs) for acceleration time histories (EW component) of the
 942 M_w 7.9 Nepal earthquake, recorded at stations: VNS, AMT, BSR and TDR in the IG basin
 943 and KTP, TVU, PTN and THM in the Kathmandu basin. Intrinsic mode functions (IMFs) for
 944 acceleration time histories (NS and Vertical component) recorded at the mentioned stations in
 945 IG basin and Kathmandu basin is shown in electronic supplement. (NS – Figure A1 and
 946 Vertical Figure A2)

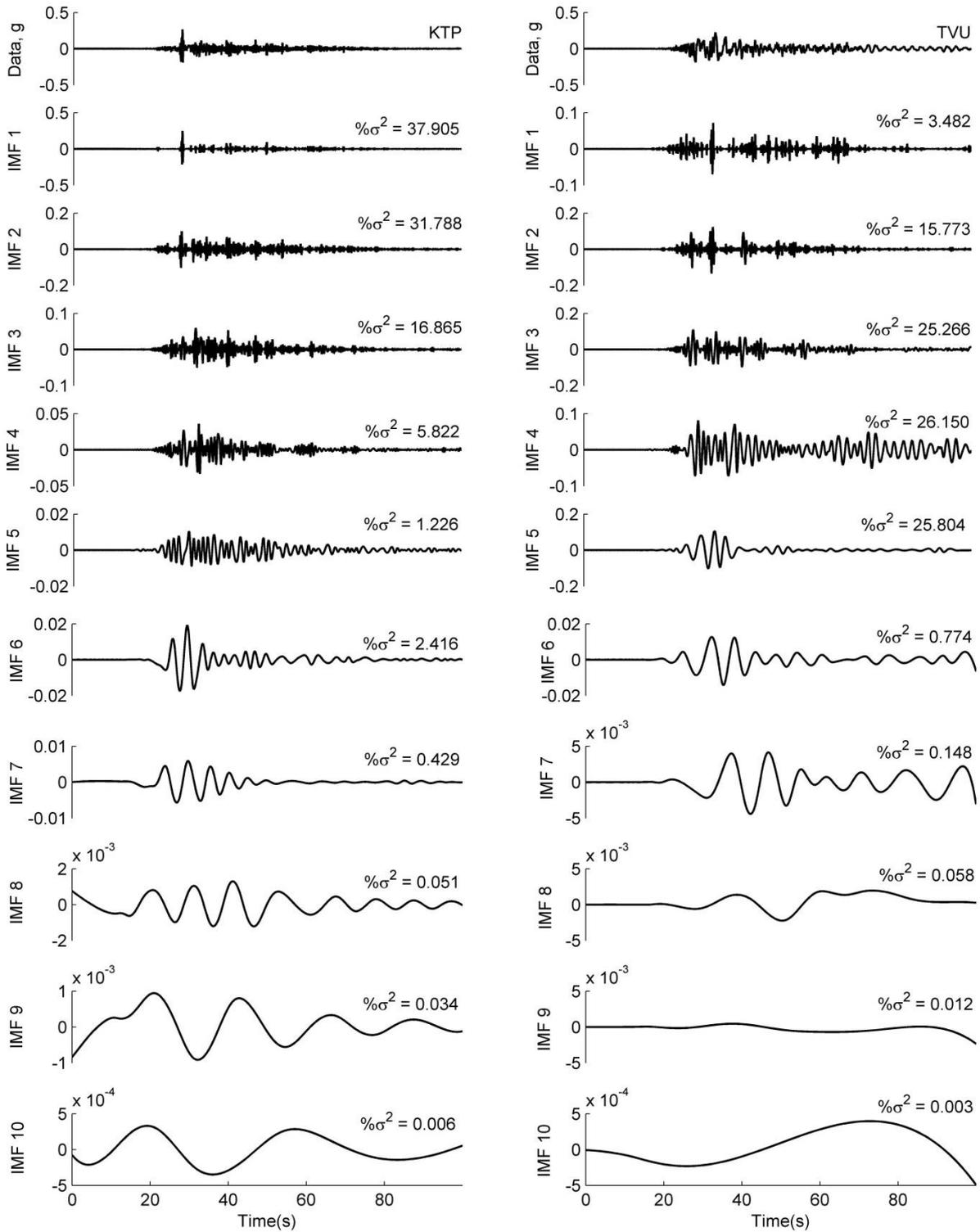


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949 **Fig. 7** [Continued] Intrinsic mode functions (IMFs) for acceleration time histories of the Mw
 950 7.9 Nepal earthquake, recorded at stations: VNS, AMT, BSR and TDR in the IG basin and
 951 KTP, TVU, PTN and THM in the Kathmandu basin. Intrinsic mode functions (IMFs) for
 952 acceleration time histories (NS and Vertical component) recorded at the mentioned stations in
 953 IG basin and Kathmandu basin is shown in electronic supplement. (NS – Figure A1 and
 954 Vertical Figure A2)

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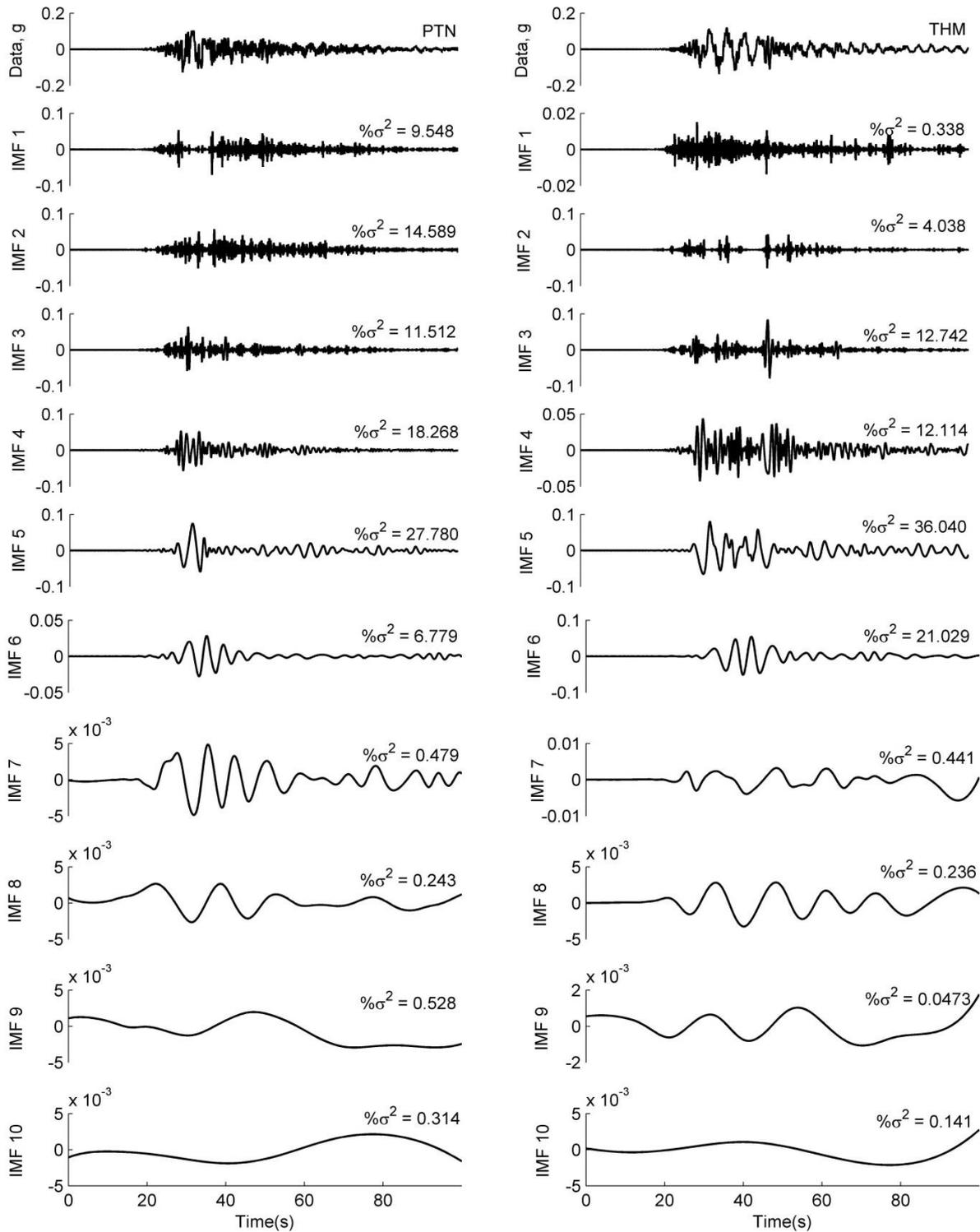
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Fig. 7 [Continued] Intrinsic mode functions (IMFs) for acceleration time histories of the Mw 7.9 Nepal earthquake, recorded at stations: VNS, AMT, BSR and TDR in the IG basin and KTP, TVU, PTN and THM in the Kathmandu basin. Intrinsic mode functions (IMFs) for acceleration time histories (NS and Vertical component) recorded at the mentioned stations in IG basin and Kathmandu basin is shown in electronic supplement. (NS – Figure A1 and Vertical Figure A2)

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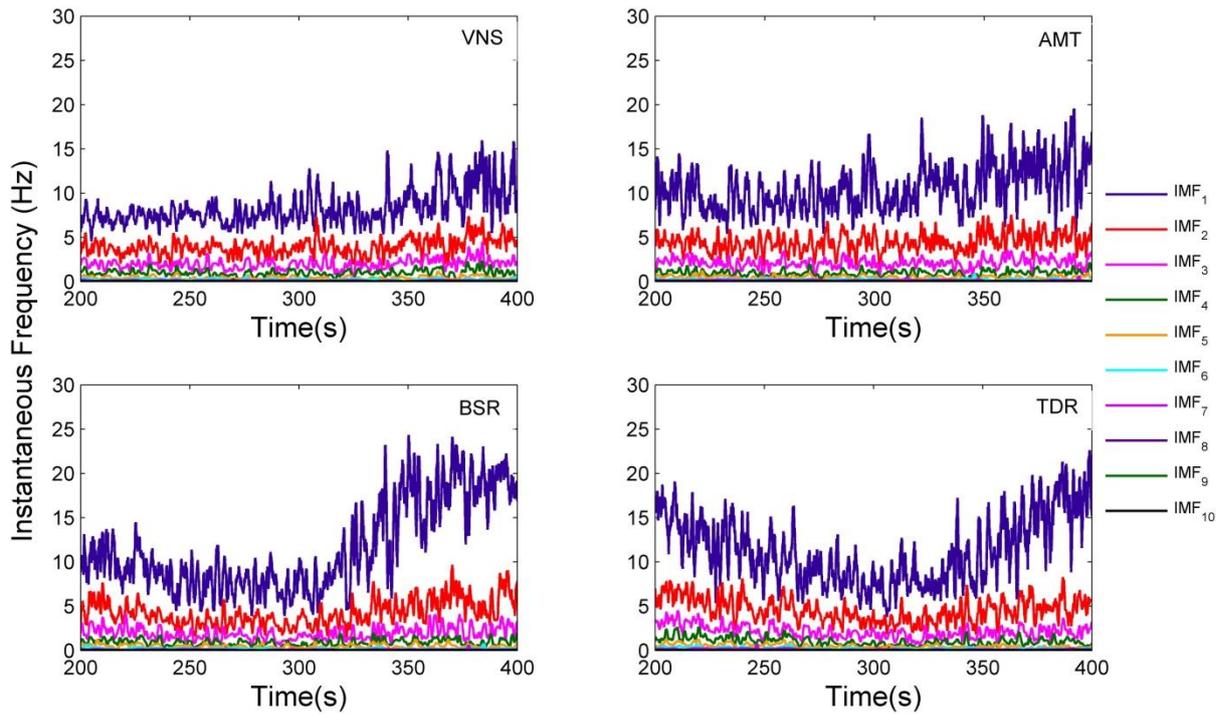
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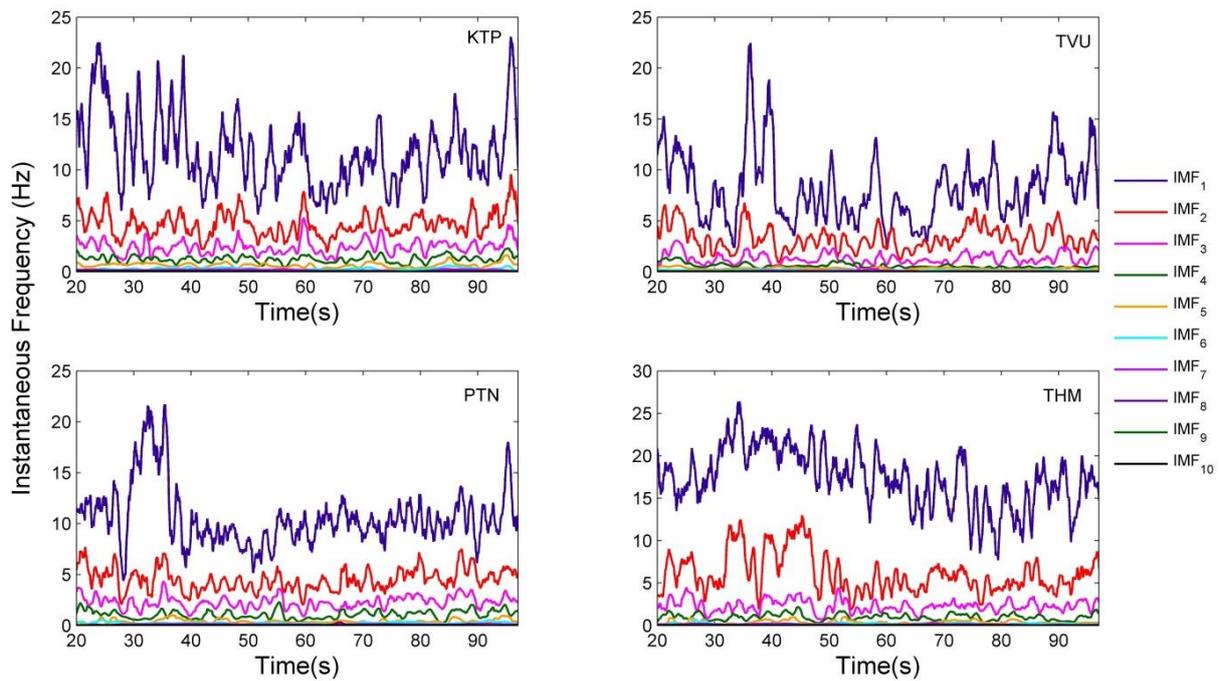
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Fig. 7 [Continued] Intrinsic mode functions (IMFs) for acceleration time histories of the Mw 7.9 Nepal earthquake, recorded at stations: VNS, AMT, BSR and TDR in the IG basin and KTP, TVU, PTN and THM in the Kathmandu basin. Intrinsic mode functions (IMFs) for acceleration time histories (NS and Vertical component) recorded at the mentioned stations in IG basin and Kathmandu basin is shown in electronic supplement. (NS – Figure A1 and Vertical Figure A2)



(a)

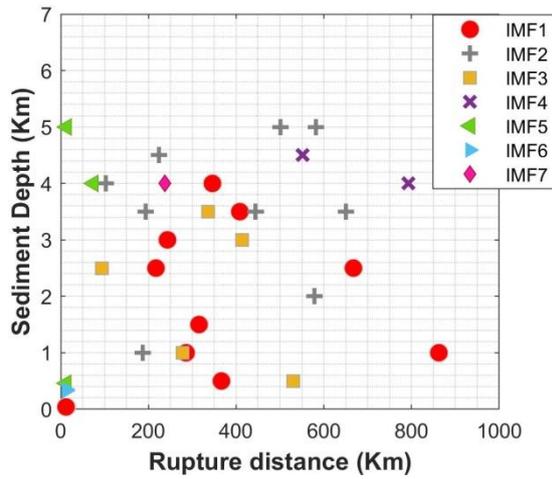


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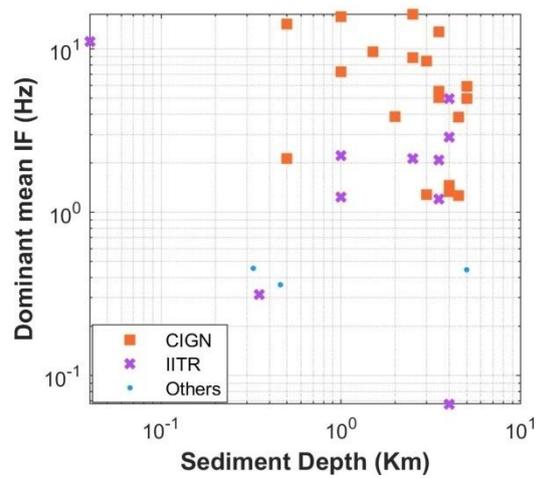
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975 **Fig. 8** Instantaneous frequencies of acceleration time histories (EW component) for stations
 976 (a) VNS, AMT, BSR and TDR in the IG basin and (b) KTP, TVU, PTN and THM in the
 977 Kathmandu basin. Instantaneous frequencies of acceleration time histories (NS and Vertical
 978 component) recorded at the mentioned stations in IG basin and Kathmandu basin is shown in
 979 electronic supplement. (NS – Figure A3 and Vertical Figure A4)

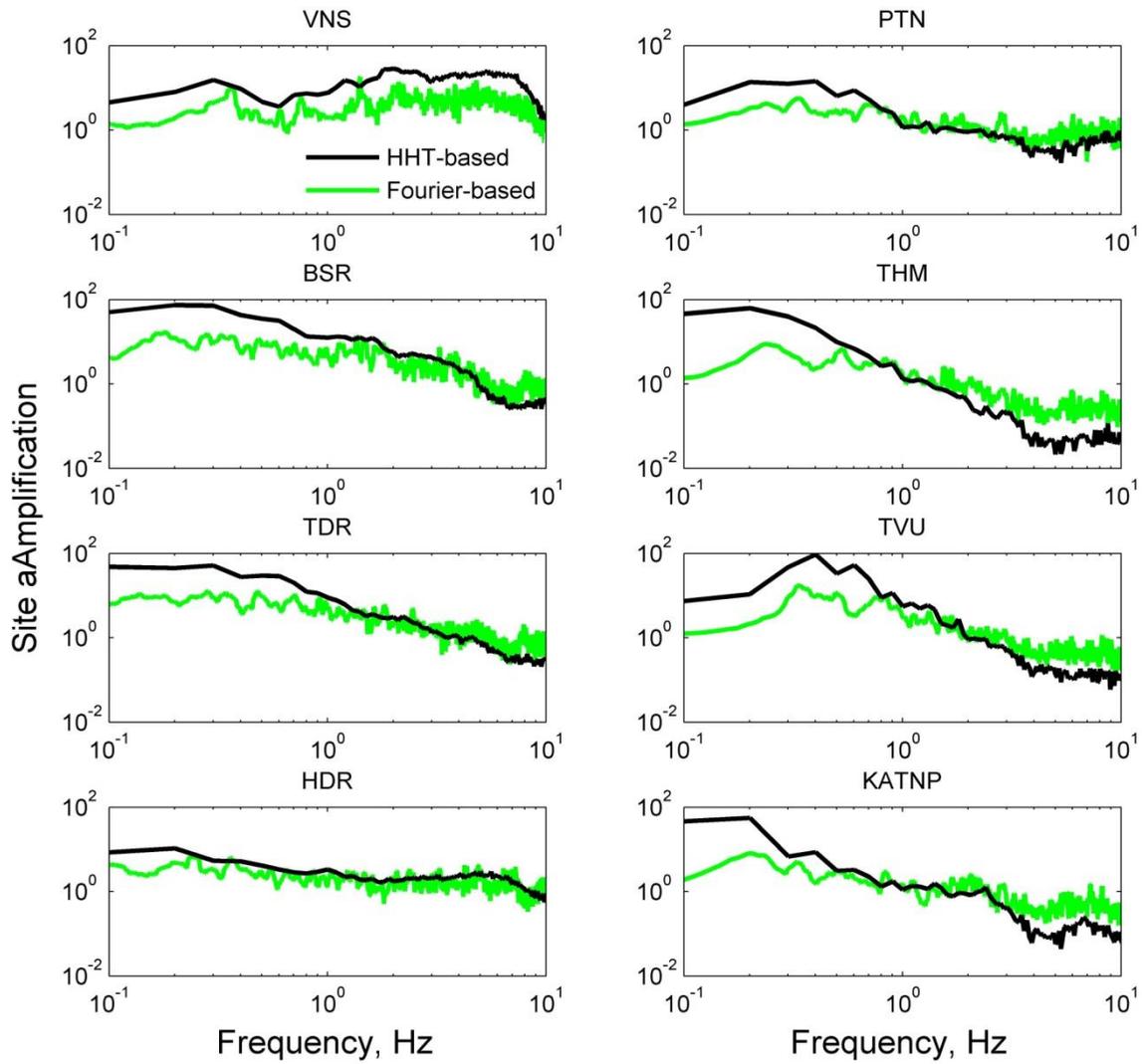


(a)



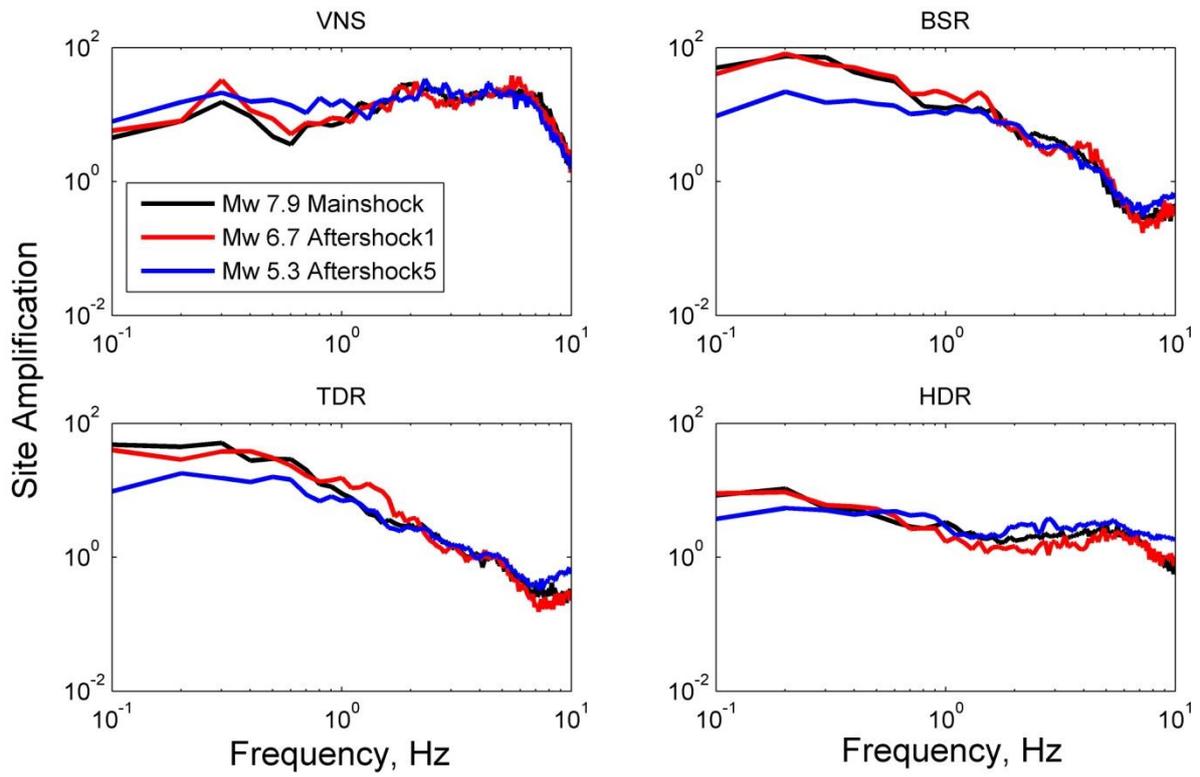
(b)

980 **Fig. 9** (a) Distribution of dominant Intrinsic Mode Functions (IMFs) of the ground motion
 981 time histories of the mainshock with rupture distance (Rrup) and Sediment depth (D); (b)
 982 Variation of dominant mean Instantaneous frequencies (Ifs) with respective site sediment
 983 depth values



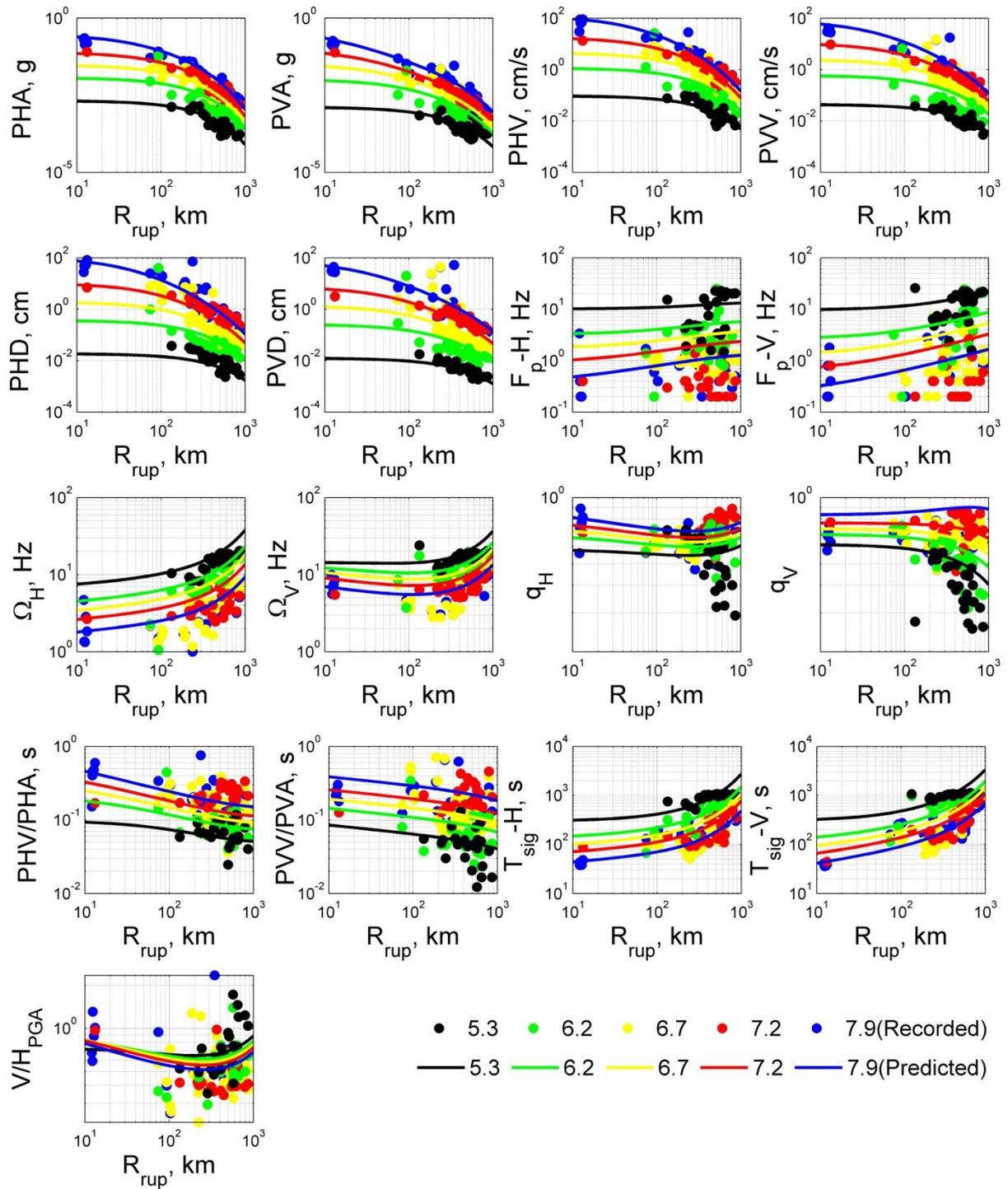
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985 **Fig. 10** Comparison between HHT-based and Fourier based site amplifications for stations (a)
 986 VNS, AMT, BSR and TDR with reference hard sites as ALB, ALM, ALM and RPG in the IG
 987 basin and (b) TVU, PTN, THM and KATNP with KTP as reference hard site in the
 988 Kathmandu basin



989

990 **Fig. 11** HHT-based site amplifications for stations VNS, AMT, BSR and TDR against the
 991 reference sites ALB, ALM, ALM and RPG of the IG basin, for the M_w 7.9 mainshock and the
 992 M_w 6.7 M_w 5.3) aftershocks

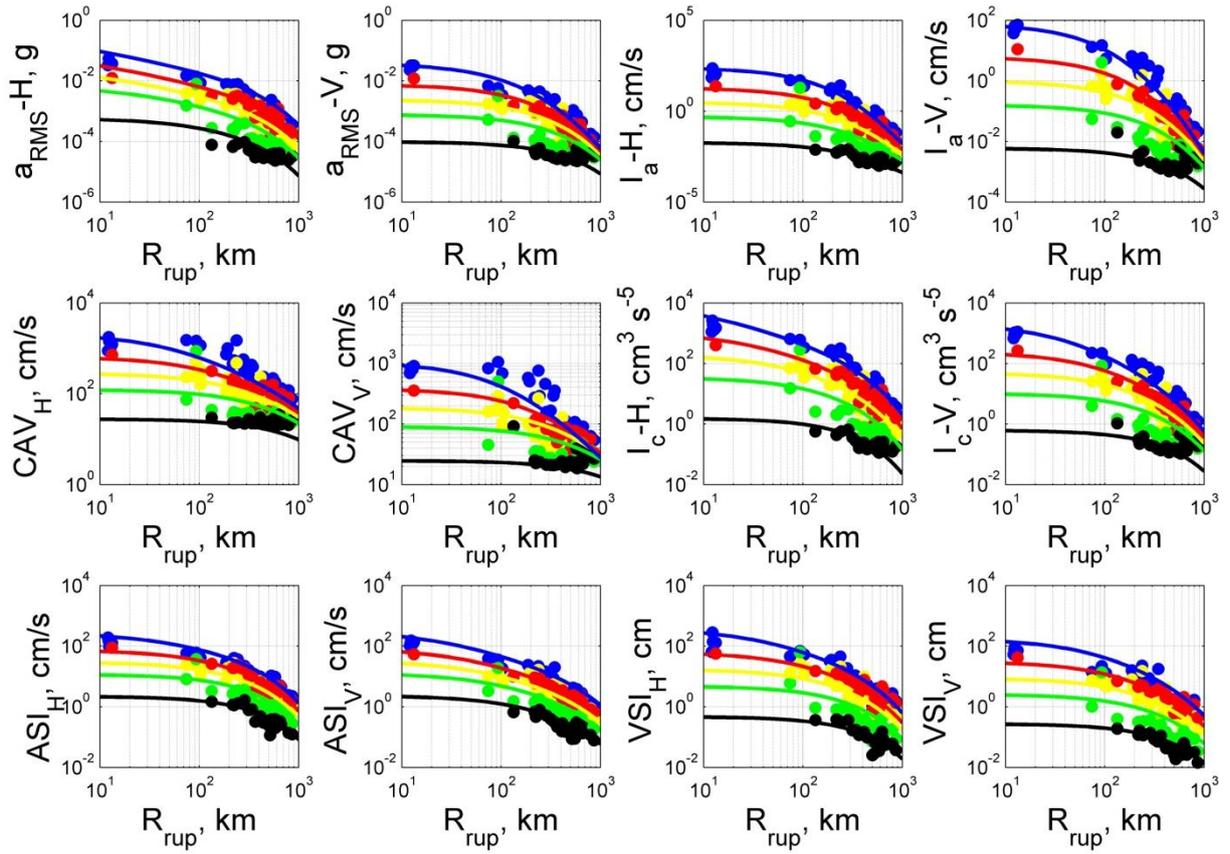


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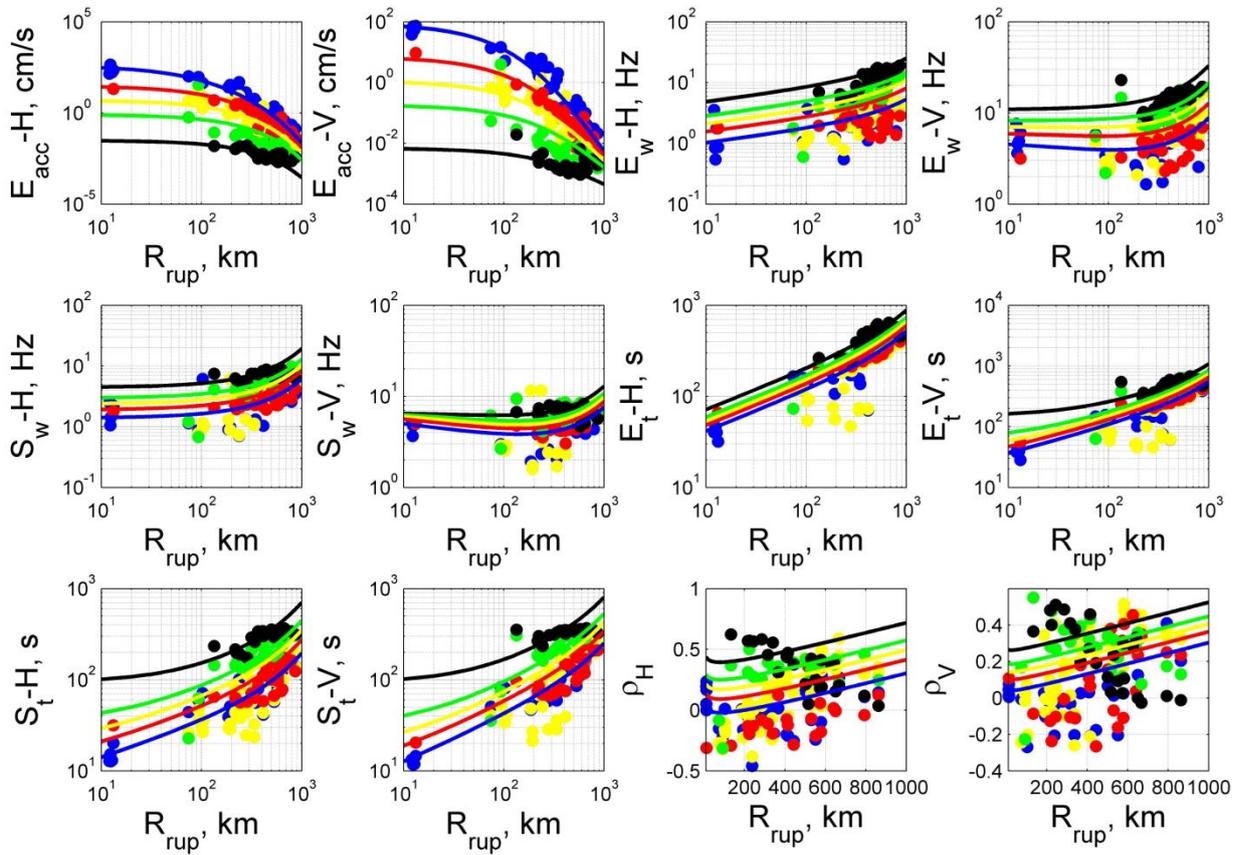
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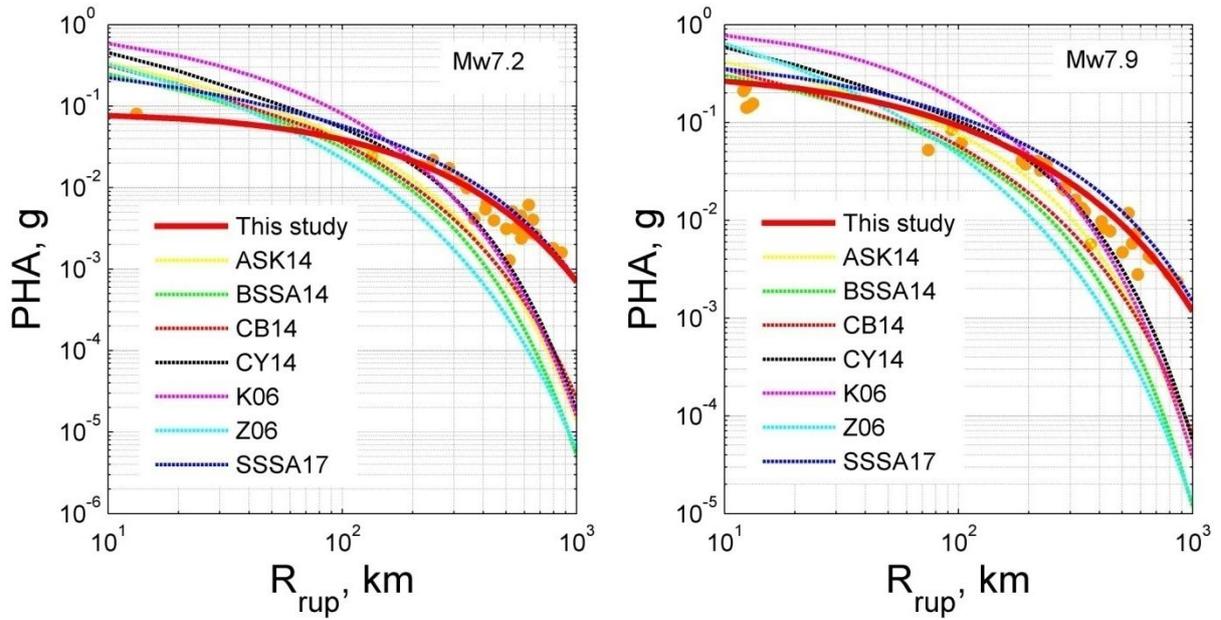
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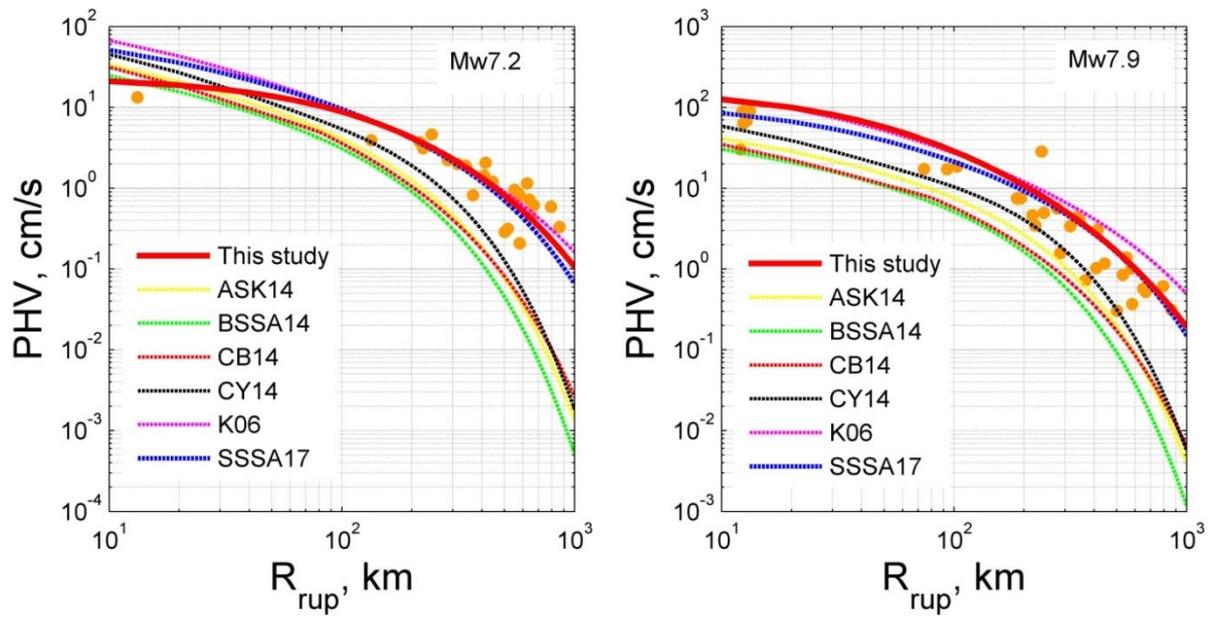
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1000 **Fig. 12** Variation of horizontal and vertical components of the 21 GMPs: (a) First order
1001 parameters, (b) Second order parameters and (c) Third order parameters, with rupture
1002 distance (R_{rup}), showing the recorded data (circular markers) and the predicted values (solid
1003 lines) for different moment magnitudes: M_w 7.9 – blue, M_w 7.2 - red, M_w 6.7 – yellow, M_w
1004 6.2 -green, M_w 5.3 – black



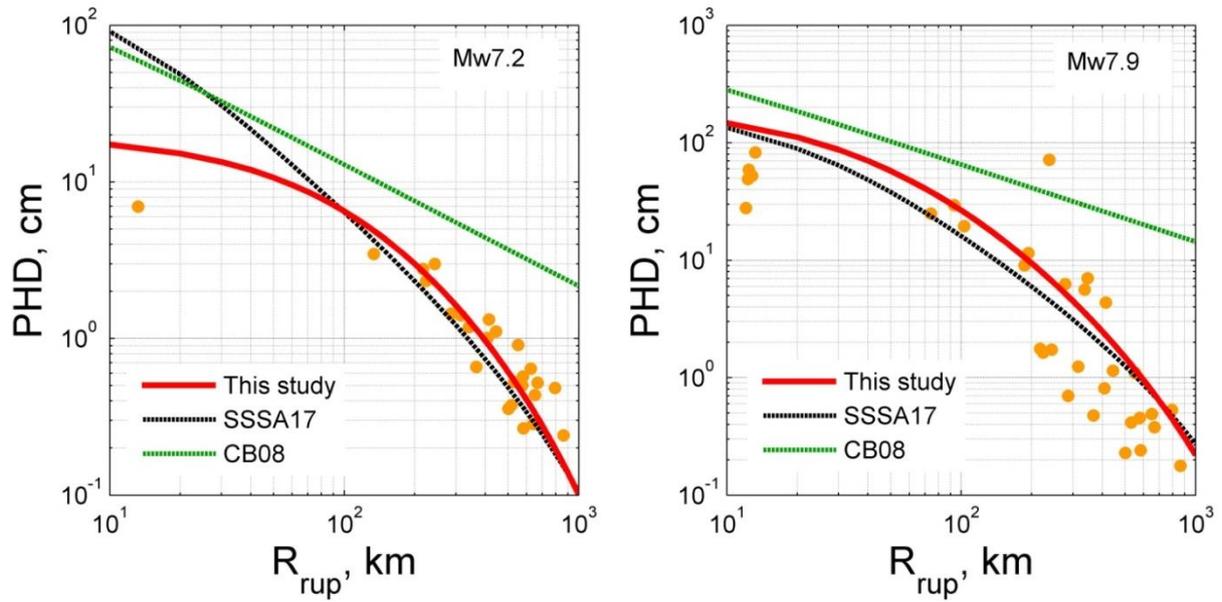
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1006 **Fig. 13** Comparisons of the GMPE developed for PGA for IG basin with other empirical
 1007 prediction equations: ASK14 – Abrahamson et al. (2014) BSSA14 – Boore et al. (2014);
 1008 CB14 – Campbell and Bozorgnia (2014); CY14 – Chiou and Youngs (2014); K06 – Kanno et
 1009 al. (2006), Z06 – Zhao et al. (2006) and SSSA17 – Singh et al. (2017)



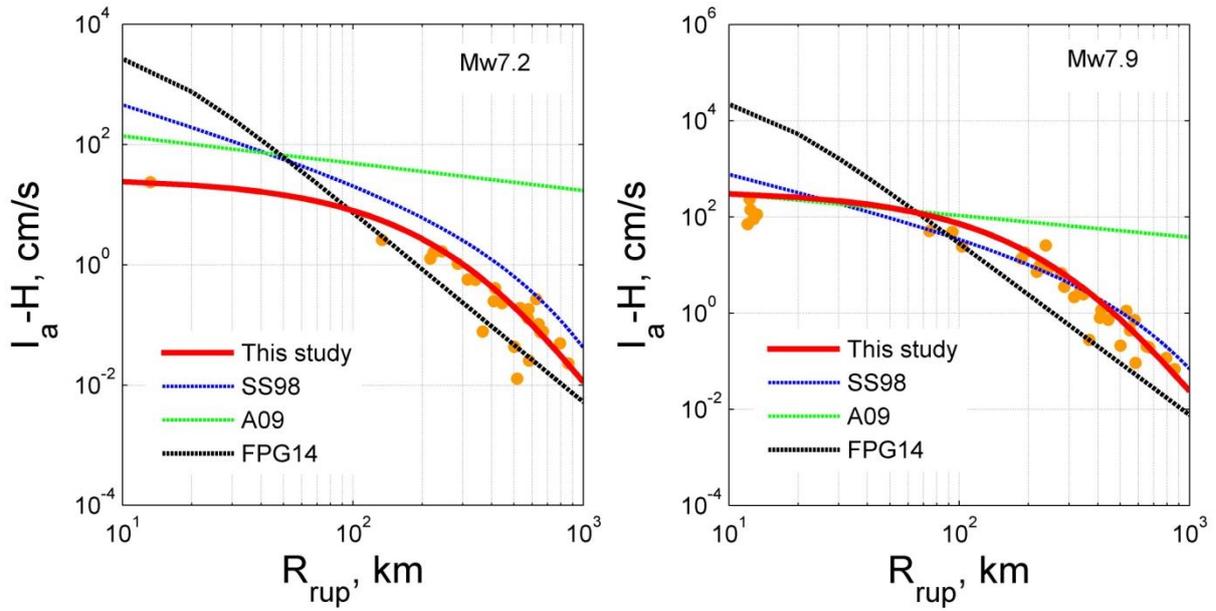
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1011 **Fig. 14** Comparisons of the GMPE developed for PGV for IG basin with other empirical
 1012 prediction equations: ASK14 – Abrahamson et al. (2014) BSSA14 – Boore et al. (2014);
 1013 CB14 – Campbell and Bozorgnia (2014); CY14 – Chiou and Youngs (2014); K06 – Kanno et
 1014 al. (2006) and SSSA17 – Singh et al. (2017)



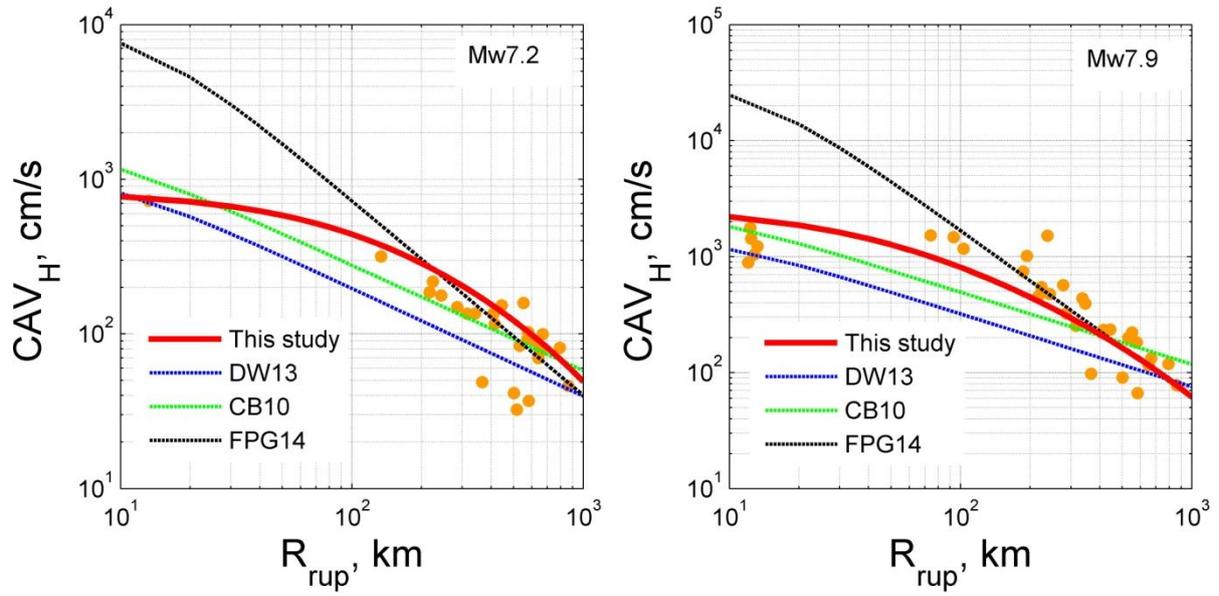
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1016 **Fig. 15** Comparisons of the GMPE developed for PGD for IG basin with other empirical
 1017 prediction equations: SSSA17 – Singh et al. (2017) and CB08 – Campbell and Bozorgnia
 1018 (2008)



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1020 **Fig. 16** Comparisons of the GMPE developed for Arias Intensity (I_a) for IG basin with other
 1021 empirical prediction equations: SS98 – Sharma and Srbulov (1998), A09 – Amiri et al. (2009)
 1022 and FPG14 – Foulser-Piggot and Goda (2014)



1023

1024 **Fig. 17** Comparison of the GMPE developed in this study for CAV for IG basin with other
 1025 empirical prediction equations: CB10 – Campbell and Bozorgnia (2010), FPG14 – Foulser-
 1026 Piggot and Goda (2014) and DW13 – Du and Wang (2013)

Supplementary Files

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- [ElectronicSupplementGMPsfor2015NepalEarthquake.docx](#)